HYPERBOLIC CONSERVATION LAWS ON MANIFOLDS: TOTAL VARIATION ESTIMATES AND THE FINITE VOLUME METHOD*

PAULO AMORIM[†], MATANIA BEN-ARTZI[‡], AND PHILIPPE G. LEFLOCH[†]

Abstract. This paper investigates some properties of entropy solutions of hyperbolic conservation laws on a Riemannian manifold. First, we generalize the Total Variation Diminishing (TVD) property to manifolds, by deriving conditions on the flux of the conservation law and a given vector field ensuring that the total variation of the solution along the integral curves of the vector field is non-increasing in time. Our results are next specialized to the important case of a flow on the 2-sphere, and examples of flux are discussed. Second, we establish the convergence of the finite volume methods based on numerical flux-functions satisfying monotonicity properties. Our proof requires detailed estimates on the entropy dissipation, and extends to general manifolds an earlier proof by Cockburn, Coquel, and LeFloch in the Euclidian case.

Key words. hyperbolic conservation law, Riemannian manifold, entropy solution, total variation, finite volume method

AMS subject classifications. 35L65, 74J40, 58J, 76N10

1. Introduction. In this paper, following [1], we investigate some properties of entropy solutions to the Cauchy problem for hyperbolic conservation laws on a manifold:

(1.1)
$$\partial_t u + \nabla_q \cdot f(u, \cdot) = 0, \qquad u = u(t, x) \in \mathbb{R}, \quad t \ge 0, x \in M.$$

Here, M is a d-dimensional, smooth manifold endowed with a Riemannian metric g and, for each constant $\overline{u} \in \mathbb{R}$, the map $x \mapsto f(\overline{u}, x)$ is a smooth vector field on M, i.e. a section of the tangent bundle TM. The well-posedness theory for the hyperbolic conservation law was recently established in Ben-Artzi and LeFloch [1]. As in Kruzkov's theory [5] which applies to conservation laws in the Euclidian setting $M = \mathbb{R}^d$, one is interested in weak solutions of (1.1) in the sense of distributions that are constrained by entropy inequalities. For instance, in the case that f is geometry compatible [1], in the sense that

(1.2)
$$\nabla_{a} \cdot f(\overline{u}, \cdot) = 0 \quad \text{for every constant } \overline{u},$$

the entropy inequalities read

$$\partial_t U(u) + \nabla_a \cdot F(u, \cdot) \le 0,$$

for every convex $U: \mathbb{R} \to \mathbb{R}$, where the vector field $x \mapsto F(\overline{u}, x)$ is

$$\partial_u F(\overline{u}, x) := \partial_u U(\overline{u}) \partial_u f(\overline{u}, x), \quad \overline{u} \in \mathbb{R}, x \in M.$$

Let us emphasize that the equation (1.1) is a geometric partial differential equation which naturally depends on the geometry of the manifold, only. In particular, all estimates derived on solutions to (1.1) should take a form that is completely independent of any particular system of local coordinates of M, although in practice, for

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[†]Laboratoire Jacques-Louis Lions & Centre National de la Recherche Scientifique, UMR 7598, University of Paris 6, BC 187, 75252 Paris (amorim@ann.jussieu.fr; lefloch@ann.jussieu.fr).

[‡]Institute of Mathematics, Hebrew University, Jerusalem 91904, Israel (mbartzi@math.huji.ac.il).

the proofs it will often be convenient to introduce a particular chart to represent the manifold. Throughout, the convention of implicit summation over repeated indices will be used. In local coordinates $x = (x^j)_{1 \le j \le d}$, we will use the short-hand notation $\partial_j := \partial/\partial x^j$. Recall that the divergence operator arising in (1.1) takes the form

$$\nabla_g \cdot f(u(t,x),x) := \frac{1}{\sqrt{|g(x)|}} \, \partial_j \left(\sqrt{|g(x)|} \, f^j(u(t,x),x) \right),$$

where (g_{ij}) are the coordinates of the metric tensor g and $|g| := \det(g_{ij})$. Here, $f^j(\overline{u}, x)$ are, for every \overline{u} , the coordinates of the vector field $f(\overline{u}, \cdot)$ at the point x.

The present paper supplements the well-posedness results established in the companion paper [1] and has two main objectives.

First, we investigate some properties of the total variation of solutions of (1.1) on a manifold. Precisely, we generalize the Total Variation Diminishing (TVD) property to manifolds, by deriving conditions on the flux f and a given vector field X ensuring that the total variation of the solution along the integral curves of X is non-increasing in time. Recall that TVD schemes for nonlinear hyperbolic problems play a central role in scientific computation, for instance in gas dynamics. Our diminishing total variation properties provide certain a priori estimates which, for instance, can be tested numerically and may help in designing robust schemes that are consistent with large-time asymptotics. Note that solutions of a conservation law on a manifold need not have a total variation that is bounded uniformly in time. The geometric effets may contribute to amplify the wave strengths, and the total variation of a solution may blow-up in the large as $t \to \infty$. We also investigate here in some detail the case where the manifold is the sphere S^2 embedded in \mathbb{R}^3 . Examples of flux (as vector fields in \mathbb{R}^3 tangent to the sphere) are discussed from both the intrinsic and embedded standpoints.

Second, we consider the numerical approximation of the entropy solutions of (1.1) and we establish the convergence of the finite volume scheme on manifolds when the numerical flux-functions depend monotonically upon their arguments. Our result is an extension to general manifolds of a theorem due to Cockburn, Coquel, and LeFloch [2] (see also [3]) in the Euclidian case. While the convergence of the finite volume scheme is easily established on a cartesian mesh, due to the invariance by translation of both the equation and the mesh, the convergence proof for non-cartesian meshes is more involved and, as was pointed out in [2], requires DiPerna's concept of measure-valued solutions [4]. To handle (1.1), we therefore rely on Ben-Artzi-LeFloch's extension to manifolds [1] of DiPerna's theorem.

An outline of this paper is as follows. In Section 2, we derive conditions on the flux and the vector field X ensuring that the function X(u) satisfies a conservation law whenever u is a solution of (1.1). This leads us to a diminishing total variation property for a class of flux and vector fields. In Section 3, we discuss the structure and general properties of conservation laws on the sphere S^2 . In Section 4, we apply the framework of Section 2 to the case of the sphere and derive total variation bounds. Next, Sections 5 and 6 are devoted to the statement and to the proof of the convergence of the finite volume scheme.

2. Total variation diminishing estimates on a general manifold. In the present section, we derive conditions on the flux f of the conservation law (1.1) which ensure that the total variation (at least along certain vector fields) of the entropy solutions of (1.1) is non-increasing in time. We will first state the main results (Propo-

sition 2.1 and Theorem 2.3) and then recall some elementary notions from differential geometry, before giving the proofs of the results.

2.1. Statement of the estimate. Throughout, M is a smooth, d-dimensional, Riemannian manifold (M, g), which has no boundary and need not be compact. The following result provides us with a key identity on directional derivatives of solutions of (1.1).

PROPOSITION 2.1. Let $u : \mathbb{R}_+ \times M \to \mathbb{R}$ be a smooth solution of the conservation law (1.1) on M where, $f = f(\overline{u}, \cdot)$ is a (smooth) vector field depending on the parameter \overline{u} . Then, given any (smooth) vector field X on M the function $w := X(u) : \mathbb{R}_+ \times M \to \mathbb{R}$ satisfies the (linear) hyperbolic equation

$$(2.1) \partial_t w + \nabla_g \cdot (w f_u(u, \cdot)) = -g(\nabla_g u, (\mathcal{L}_X f_u)(u, \cdot)) - X(\nabla_g \cdot f)(u, \cdot),$$

where ∇_g denotes the covariant derivative operator associated with g and $(\mathcal{L}_X f_u)(u,\cdot)$ denotes the Lie bracket of the vectors fields X, f_u , that is

$$\mathcal{L}_X f_u(\overline{u},\cdot) := [f_u(\overline{u},\cdot), X], \quad \overline{u} \in \mathbb{R}.$$

The first term in the right-hand side of (2.1) depends on the gradient of the unknown function u while the second term depends on the function itself. Observe that our formula is non-trivial even when M is the Euclidian space \mathbb{R}^d . In that case, the equation (2.1) becomes

$$\partial_t w + \nabla \cdot (w f_u(u, x)) = -\nabla u \cdot [X, f_u](u, x) - X \cdot \nabla (\nabla \cdot f)(u, x),$$

where $w = X \cdot \nabla u$. When the flux is independent of x it is obvious that for any constant vector field the right-hand side of (2.1) vanishes and we can recover the total variation diminishing property for scalar conservation laws in \mathbb{R}^d .

We denote by TV the total variation functional for functions defined on the Riemannian manifold (M,g). We introduce here a generalization, which extends to discontinuous functions the formula, valid for smooth u,

$$TV_X(u) := \int_M |X(u(t,x))| \, dv_g,$$

where dv_q is the volume element on M.

Definition 2.2. 1. To any function u and vector field X defined on the manifold M one associates the quantity

$$TV_X(u) := \sup_{\phi} \int_M u \, \nabla_g \cdot (\phi \, X) \, dv_g,$$

where the supremum is taken over all smooth functions $\phi: M \to \mathbb{R}$ satisfying $\|\phi\|_{L^{\infty}} \le 1$. When $TV_X(u) < \infty$ the function u is said to have bounded total variation along X.

2. A flux $f = f(\overline{u}, \cdot)$ on M is said to be geometry compatible [1] if

$$\nabla_g \cdot f(\overline{u}, \cdot) = 0, \quad \overline{u} \in \mathbb{R}.$$

It follows from Proposition 2.1 and the existence theory in [1] that :

Theorem 2.3. (Total variation estimates.) Let u be an entropy solution of the conservation law (1.1) on the manifold M where f is a smooth flux.

1. The total variation of u along a vector field X satisfies

(2.2)
$$TV_X(u(t)) \leq TV_X(u(0)) + \sup_{(0,t) \times M} |(\mathcal{L}_X f_u)(u, \cdot)|_g \int_0^t TV_X(u(\tau)) d\tau + ||X(\nabla_g \cdot f)(u, \cdot)||_{L^1((0,t) \times M)}.$$

2. Hence, when f is divergence-free and the Lie bracket of f_u and X vanishes,

$$\mathcal{L}_X f_u(\overline{u},\cdot) = 0, \quad \overline{u} \in \mathbb{R},$$

then the solution has bounded total variation along X for all times t > 0, if this property holds at the time t = 0, and moreover,

(2.3)
$$TV_X(u(t)) \le TV_X(u(0)), \quad t \ge 0.$$

At this juncture, it is important to recall from [1] that entropy solutions of (1.1) have bounded variation (in the standard sense) at every time t, if this is true at the time t = 0. Actually, one has the general estimate

$$TV(u(t)) \le e^{C_1 t} TV(u(0)) + C_2,$$

where C_1, C_2 depend on the metric g and on derivatives (up to second order) of the flux f. In contrast, our result in Theorem 2.3 solely assumes that $TV_X(u(0))$ is finite and concludes that $TV_X(u(t))$ is finite for all times. The property (2.3) should be useful when designing approximation schemes for (1.1), and it is natural to require (2.3) to hold at every time-step of the discretization.

Let us now recall some basic notions and notations that will be used throughout this paper. The differential of a function $u: M \to \mathbb{R}$ is the field of 1-forms defined by

$$du(X) := X(u),$$

for every vector field X. It is convenient to choose a basis of the tangent space T_xM at a point $x \in M$ together with its dual basis associated with the cotangent space T_x^*M , so that we can introduce coordinates for vectors and covectors relative to these bases. As usual, to each vector $(X^i) \in T_xM$ we associate its covector $(g_{ij}X^i)$ by lowering indices using the metric (implicit summation on repeated indices being used). Denoting (g^{ij}) the components of the inverse matrix associated with (g_{ij}) , then, to each covector $(\eta_i) \in T_x^*M$ we can also associate the vector $(g^{ij}\eta_j)$.

Recall also that the gradient of a (smooth) function $u: M \to \mathbb{R}$ is the vector field $\nabla_g u$ associated with the field of differential forms du. The divergence of a (smooth) vector field f on M is the function $\nabla_g \cdot f: M \to \mathbb{R}$ defined by

$$\int_{M} u \nabla_{g} \cdot f \, dv := -\int_{M} du(f) \, dv_{g}$$

for all smooth $u: M \to \mathbb{R}$, where dv_q is the volume element on M.

Consider now a coordinate chart (x^i) together with the associated basis of vectors $(\partial/\partial x^i)$ and covectors (dx^i) . The differential of a function u is given by

$$du = \frac{\partial u}{\partial x^i} dx^i.$$

In view of the relations $\nabla_g^i u := (\nabla_g u)^i = g^{ij} (du)_j$, we see that the components of $\nabla_g u$ in the basis $\partial/\partial x^i$ of $T_x M$ are

$$(\nabla_g u)^i = g^{ij} \, \frac{\partial u}{\partial x^j}.$$

If the support of a function u is included in the domain of definition of the coordinate chart (x^i) , recalling that |g| denotes the determinant of the metric tensor we can write

$$\int_{M} du(f) dv_{g} = \int_{\mathbb{R}^{d}} \left(\frac{\partial u}{\partial x^{i}} f^{i} \right) \sqrt{|g|} dx^{1} \cdots dx^{d}$$

$$= -\int_{\mathbb{R}^{d}} \frac{u}{\sqrt{|g|}} \frac{\partial}{\partial x^{i}} \left(f^{i} \sqrt{|g|} \right) \sqrt{|g|} dx^{1} \cdots dx^{d}$$

$$= -\int_{M} \frac{u}{\sqrt{|g|}} \frac{\partial}{\partial x^{i}} \left(f^{i} \sqrt{|g|} \right) dv_{g},$$

so that

(2.4)
$$\nabla_g \cdot f := \frac{1}{\sqrt{|g|}} \frac{\partial}{\partial x^i} \left(f^i \sqrt{|g|} \right).$$

The expression of $\nabla_g \cdot f$ follows also from computing the covariant derivatives of the vector field, as follows. Given a local chart (x^i) for M, the covariant derivative of the vector field f is the (1,1)-tensor field $\nabla_k f^j$ defined by

$$\nabla_k f^j := \partial_k f^j + \Gamma^j_{kl} f^l,$$

where Γ_{kl}^{j} are the *Christoffel symbols* given by

$$\Gamma_{kl}^{j} := \frac{1}{2} (\partial_k g_{li} + \partial_l g_{ki} - \partial_i g_{kl}) g^{ij}.$$

In particular, we have $\Gamma_{kj}^j = \frac{1}{2}g^{ij}\partial_k g_{ij}$ and, since $\partial_i |g| = |g|g^{kl}\partial_i g_{kl}$,

$$\Gamma_{ik}^k = \frac{1}{2|g|} \partial_i |g| = \partial_i \log \sqrt{|g|} = \frac{1}{\sqrt{|g|}} \partial_i (\sqrt{|g|}).$$

The divergence of a vector field f is the trace of the covariant derivative, i.e.

$$\begin{split} \nabla_g \cdot f &:= \nabla_k f^k = \partial_k f^k + \Gamma^k_{ik} f^i = \partial_k f^k + \frac{1}{\sqrt{|g|}} \partial_i \left(\sqrt{|g|} \right) f^i \\ &= \partial_k f^k + \frac{1}{\sqrt{|g|}} \left(-\sqrt{|g|} \partial_k f^k + \partial_k \left(\sqrt{|g|} f^k \right) \right) \\ &= \frac{1}{\sqrt{|g|}} \, \partial_k \left(\sqrt{|g|} f^k \right), \end{split}$$

which is (2.4).

2.2. First approach. The proof of Proposition 2.1 will follow from the following two technical lemmas.

LEMMA 2.4. For every smooth function $u: M \to \mathbb{R}$ the following identities hold $(x \in M)$:

(2.5)
$$\nabla_q \cdot (f(u,x)) = du(f_u(u,x)) + (\nabla_q \cdot f)(u,x)$$

(2.6)
$$\mathcal{L}_X(Y(u,x)) = X(u)Y_u(u,x) + (\mathcal{L}_XY)(u,x)$$
 for all vector fields $Y = Y(u,x)$.

(2.7)
$$X(h(u,x)) = X(u)h_u(u,x) + X(h)(u,x) \text{ for all functions } h.$$

Proof. In local coordinates we have

$$\nabla_g \cdot (f(u, x)) = \frac{1}{\sqrt{|g|}} \partial_i \left(\sqrt{|g|} f^i(u, x) \right)$$

$$= \frac{1}{\sqrt{|g|}} \partial_i \left(\sqrt{|g|} f^i \right) (u, x) + f_u^i(u, x) \partial_i u$$

$$= (\nabla_g \cdot f)(u, x) + du(f_u(u, x)),$$

which proves (2.5). To prove (2.6) we note that

$$\begin{split} \left(\mathcal{L}_X(Y(u,x))\right)^i &= [X,Y(u,x)]^i = X^j \partial_j (Y^i(u,x)) - Y^j(u,x) \partial_j X^i \\ &= \left(X^j \partial_j Y^i - Y^j \partial_j X^i\right) (u,x) + X^j Y_u^i(u,x) \partial_j u \\ &= [X,Y]^i (u,x) + X^j \partial_j u Y_u^i(u,x). \end{split}$$

This shows (2.6). The proof of (2.7) is completely similar. \square

LEMMA 2.5. For any (smooth) function $u: M \to \mathbb{R}$, vector field X, and flux $f = f(\overline{u}, \cdot)$, the following identity holds:

(2.8)
$$X(\nabla_g \cdot (f(u,x)))$$

$$= \nabla_g \cdot (X(u)f_u(u,x)) + g(\nabla_g u, (\mathcal{L}_X f_u)(u,x)) + X(\nabla_g \cdot f)(u,x).$$

Proof. Using (2.5) and (2.7) we find

$$X(\nabla_g \cdot (f(u,x))) = X((\nabla_g \cdot f)(u,x)) + X(g(\nabla_g u, f_u(u,x)))$$

= $X(\nabla_g \cdot f)(u,x) + X(u)(\nabla_g \cdot f_u)(u,x) + X(g(\nabla_g u, f_u(u,x))).$

On the other hand,

$$\nabla_g \cdot \big(X(u) f_u(u, x) \big) = X(u) \nabla_g \cdot (f_u(u, x)) + g(\nabla_g(X(u)), f_u(u, x))$$

= $X(u) (\nabla_g \cdot f_u)(u, x) + X(u) g(\nabla_g u, f_{uu}(u, x)) + g(\nabla_g(X(u)), f_u(u, x)),$

using (2.5). Thus, we arrive at

$$X(\nabla_g \cdot (f(u,x))) - \nabla_g \cdot (X(u)f_u(u,x))$$

= $X(\nabla_g \cdot f)(u,x) + X(g(\nabla_g u, f_u(u,x)) - X(u)g(\nabla_g u, f_{uu}(u,x)) - g(\nabla_g(X(u)), f_u(u,x)).$

Expressing the Leibnitz rule for the Lie derivative we obtain

$$X(g(\nabla_g u, f_u(u, x))) = (\mathcal{L}_X g)(\nabla_g u, f_u(u, x)) + g(\mathcal{L}_X \nabla_g u, f_u(u, x)) + g(\nabla_g u, \mathcal{L}_X (f_u(u, x))).$$

Therefore, we have

$$X(\nabla_g \cdot (f(u,x))) - \nabla_g \cdot (X(u)f_u(u,x)) = X(\nabla_g \cdot f)(u,x) + (\mathcal{L}_X g)(\nabla_g u, f_u(u,x))$$
$$+ g(\mathcal{L}_X \nabla_g u - \nabla_g (X(u)), f_u(u,x))$$
$$+ g(\nabla_g u, \mathcal{L}_X (f_u(u,x)) - X(u)f_{uu}(u,x))).$$

Using (2.6), the last term above equals $g(\nabla_q u, (\mathcal{L}_X f_u)(u, x))$.

The proof of the lemma will be complete once we check that, for all vector fields X, Z and functions u,

(2.9)
$$g(\mathcal{L}_X \nabla_q u - \nabla_q(X(u)), Z) = -(\mathcal{L}_X g)(\nabla_q u, Z).$$

To this end, in local coordinates we can write

$$\begin{split} \left(\mathcal{L}_{X}\nabla_{g}u - \nabla_{g}(X(u))\right)^{i} &= X^{j}\,\partial_{j}\nabla^{i}u - \nabla^{j}u\,\partial_{j}X^{i} - \nabla^{i}(X^{j}\,\partial_{j}u) \\ &= X^{j}\left(\partial_{j}\nabla^{i}u - \nabla^{i}\partial_{j}u\right) - \nabla^{j}u\,\partial_{j}X^{i} - \nabla^{i}X^{j}\,\partial_{j}u. \end{split}$$

For the first term above we have

$$X^{j} \left(\partial_{j} \nabla^{i} u - \nabla^{i} \partial_{j} u \right) = X^{j} \left(\partial_{j} (g^{ik}) \partial_{k} u + g^{ik} (\partial_{j} \partial_{k} u - \partial_{k} \partial_{j} u) \right)$$
$$= X(g^{ik}) g_{kl} \nabla^{l} u$$
$$= -X(g_{kl}) g^{ik} \nabla^{l} u.$$

For the remaining two terms, a tedious but straightforward computation gives

$$-\nabla^{j} u \,\partial_{j} X^{i} - \nabla^{i} X^{j} \,\partial_{j} u = -\nabla^{j} u \,g^{il} \big((\mathcal{L}_{X} g)_{jl} - X(g_{lj}) \big),$$

where the components of $\mathcal{L}_X g$ are computed from the formula

$$X(g(Y,Z)) = (\mathcal{L}_X g)(Y,Z) + g(\mathcal{L}_X Y,Z) + g(Y,\mathcal{L}_X Z).$$

Therefore, we find

$$\left(\mathcal{L}_X \nabla_g u - \nabla_g (X(u))\right)^i = -g^{ki} \nabla^j u (\mathcal{L}_X g)_{jk},$$

and, finally,

$$g(\mathcal{L}_X \nabla_g u - \nabla_g(X(u)), Z) = -g_{im} g^{ki} \nabla^j u(\mathcal{L}_X g)_{jk} Z^m$$
$$= -\nabla^j u(\mathcal{L}_X g)_{jk} Z^k$$
$$= -(\mathcal{L}_X g)(\nabla_g u, Z),$$

which proves (2.9). This completes the proof of Lemma 2.5. \square

Proof of Proposition 2.1. Applying the vector field X to the conservation law (1.1), we find

$$\partial_t X(u) + X(\nabla_q \cdot (f(u,x))) = 0.$$

The first term is precisely $\partial_t w$. Using Lemma 2.5 to evaluate the second term, we find the desired identity (2.1). This completes the proof. \square

2.3. Second approach. We provide here a second approach to Proposition 2.1. To any vector field X we associate the (local) one-parameter group of diffeomorphisms $\varphi: M \to M$ (defined for all sufficiently small s) from the integral curves of the vector field X,

(2.10)
$$\frac{d\varphi_s^i}{ds}(x) = X^i(\varphi_s(x)), \qquad \varphi_0(x) = x.$$

Given a function $u: M \to \mathbb{R}$ we may consider the composite function $u \circ \varphi_s$.

LEMMA 2.6. Let X be a smooth vector field on M and (φ_s) its associated one-parameter group of diffeomorphisms. Then, for any smooth function $u: M \to \mathbb{R}$ it holds

(2.11)
$$\nabla_g \cdot (f(u \circ \varphi_s, \cdot))$$

$$= \nabla_g \cdot (f(u, \cdot)) \circ \varphi_s - s \left(X(\nabla_g \cdot f)(u, \cdot) + g(\nabla_g u, (\mathcal{L}_X f_u)(u, \cdot)) \right) + \mathcal{O}(s^2).$$

Proof. We will first check that for any vector field Y,

(2.12)
$$Y^{j}(\varphi_{s}(x)) - \partial_{i}\varphi_{s}^{j}Y^{i}(x) = s \mathcal{L}_{X}Y^{j} + \mathcal{O}(s^{2}).$$

First, it is obvious that

$$Y^j \circ \varphi_s = Y^j + s X^k \partial_k Y^j + \mathcal{O}(s^2).$$

Second, observe that for all i, j

$$\frac{d}{ds}\partial_i\varphi_s^j = \partial_i\frac{d}{ds}\varphi_s^j = \partial_i(X^j \circ \varphi_s) = (\partial_k X^j) \circ \varphi_s \,\partial_i\varphi_s^k.$$

Also, since $\varphi_0(x) = x$, we have $\partial_i \varphi_s^j|_{s=0} = \delta_i^j$ and so $\frac{d}{ds} \partial_i \varphi_s^j|_{s=0} = \partial_i X^j$. Therefore, Taylor expanding yields us

$$\partial_i \varphi_s^j = \delta_i^j + s \, \partial_i X^j + \mathcal{O}(s^2).$$

In consequence we find

$$Y^{j}(\varphi_{s}(x)) - \partial_{i}\varphi_{s}^{j}Y^{i}(x) = s X^{i}\partial_{i}Y^{j} - s Y^{i}\partial_{i}X^{j} + \mathcal{O}(s^{2})$$
$$= s \mathcal{L}_{X}Y^{j} + \mathcal{O}(s^{2}),$$

which is precisely (2.12).

Setting $v := u \circ \varphi_s$, we start from (2.5) which, for the function v, reads

(2.13)
$$\nabla_g \cdot (f(v,\cdot)) = (\nabla_g \cdot f)(v,\cdot) + dv(f_u(v,\cdot)).$$

Let us compute the last term of the above identity. In local coordinates, we can write

$$(dv)_i = \partial_i v = \partial_i (u \circ \varphi_s)$$
$$= ((\partial_i u) \circ \varphi_s) \partial_i \varphi_s^j,$$

and therefore

$$dv(f_u(v,\cdot)) = ((\partial_j u) \circ \varphi_s) \partial_i \varphi_s^j f_u^i(v,\cdot)$$

$$= ((\partial_j u) \circ \varphi_s) (\partial_i \varphi_s^j f_u^i(v,\cdot) - f_u^j(u,\cdot) \circ \varphi_s) + (\partial_j u f_u^j(u,\cdot)) \circ \varphi_s$$

$$= ((\partial_j u) \circ \varphi_s) (\partial_i \varphi_s^j f_u^i(v,\cdot) - f_u^j(u,\cdot) \circ \varphi_s) + du(f_u(u,\cdot)) \circ \varphi_s.$$

Taking (2.12) into account and noting that

$$((\partial_j u) \circ \varphi_s) \mathcal{L}_X f_u^j(v, \cdot) = \partial_j u \mathcal{L}_X f_u^j(u, \cdot) + \mathcal{O}(s)$$

= $g(\nabla_g u, \mathcal{L}_X f_u(u, \cdot)) + \mathcal{O}(s)$,

we arrive at

$$dv(f_{u}(v,\cdot))$$

$$= -s(\partial_{j}u) \circ \varphi_{s}(\mathcal{L}_{X}f_{u}^{j})(v,\cdot) + du(f_{u}(u,\cdot)) \circ \varphi_{s} + \mathcal{O}(s^{2}).$$

$$= -sg(\nabla_{g}u,(\mathcal{L}_{X}f_{u})(u,\cdot)) + du(f_{u}(u,\cdot)) \circ \varphi_{s} + \mathcal{O}(s^{2})$$

$$= -sg(\nabla_{g}u,(\mathcal{L}_{X}f_{u})(u,\cdot)) + \nabla_{g}\cdot(f(u,\cdot)) \circ \varphi_{s} - (\nabla_{g}\cdot f)(u,\cdot) \circ \varphi_{s} + \mathcal{O}(s^{2}),$$

where we have also used (2.5) with the function u.

Returning now to (2.13), using the identity for $dv(f_u(v,\cdot))$, and re-ordering the terms, we find

$$\nabla_g \cdot (f(v,\cdot)) = \nabla_g \cdot (f(u,\cdot)) \circ \varphi_s + (\nabla_g \cdot f)(v,\cdot) - (\nabla_g \cdot f)(u,\cdot) \circ \varphi_s - s g(\nabla_g u, (\mathcal{L}_X f_u)(u,\cdot)) + \mathcal{O}(s^2).$$

Finally, observing that

$$(\nabla_g \cdot f)(v, x) - (\nabla_g \cdot f)(v, \varphi_s(x)) = -s \frac{d}{ds} (\nabla_g \cdot f)(v, \varphi_s(x))|_{s=0} + \mathcal{O}(s^2)$$

$$= -s X^k \partial_k (\nabla_g \cdot f)(u, \cdot) + \mathcal{O}(s^2)$$

$$= -s X(\nabla_g \cdot f)(u, \cdot) + \mathcal{O}(s^2)$$

concludes the derivation of (2.11). \square

Proof of Proposition 2.1. We provide here a second proof of this proposition. If u is a smooth solution to (1.1) then, observing that $\partial_t v = (\partial_t u) \circ \varphi_s$, we can rewrite (1.1) in the form

$$\partial_t v + \nabla_g \cdot (f(u, \cdot)) \circ \varphi_s = 0.$$

So, using (2.11) we get

$$(2.14) \quad \partial_t v + \nabla_g \cdot \left(f(v, \cdot) \right) = -s \left(X(\nabla_g \cdot f)(v, \cdot) + g(\nabla_g u, (\mathcal{L}_X f_u)(v, \cdot)) \right) + \mathcal{O}(s^2).$$

Now, recalling the conservation law (1.1) and the equation (2.14), we obtain

(2.15)
$$\partial_t \left(\frac{1}{s} (v - u) \right) + \nabla_g \cdot \left(\frac{1}{s} (f(v, \cdot) - f(u, \cdot)) \right) \\ = -X(\nabla_g \cdot f)(v, \cdot) - g(\nabla_g u, (\mathcal{L}_X f_u)(v, \cdot)) + \mathcal{O}(s),$$

in which

$$\lim_{s \to 0} \frac{1}{s} (v - u) = X(u)$$

and

$$\lim_{s \to 0} \frac{1}{s} (f(v, \cdot) - f(u, \cdot)) = \lim_{s \to 0} \frac{d}{ds} (f(v, \cdot))$$
$$= f_u(u, \cdot) X^i \partial_i u = X(u) f_u(u, \cdot).$$

So, letting s tend to zero in (2.15) yields the desired identity (2.1). \square

- 3. Conservation laws on the sphere S^2 . Following [1] we include here a discussion of basic properties of the conservation law (1.1) in the case (so important in the geophysical applications) that the manifold M is the sphere S^2 . We discuss here the choice and properties of the flux in (1.1) and we emphasize differences between the intrinsic and the embedded representations of the sphere. Certain properties will be closely related to the fact that the sphere S^2 can be embedded isometrically in the Euclidian space \mathbb{R}^3 .
- **3.1. Spherical coordinates on** S^2 . One convenient approach consists in parameterizing the smooth manifold S^2 punctured at the North and South poles, by using a single chart in spherical coordinates (φ, θ) , where $\varphi \in [0, 2\pi]$ denotes the longitude and $\theta \in (0, \pi)$ the latitude. Set

$$\Omega := T^1 \times (0, \pi),$$

where $T^1 = [0, 2\pi]$ with the identification of the points $\varphi = 0$ and 2π . By definition, a smooth function defined on Ω is extendable to a smooth periodic function for all $\varphi \in \mathbb{R}$. Indeed, in all what follows, we always consider smooth functions defined on S^2 represented as functions of (φ, θ) outside the poles and extended to the poles by continuity. We imbed S^2 in \mathbb{R}^3 according to

$$x = x(\varphi, \theta) = (\cos \varphi \sin \theta, \sin \varphi \sin \theta, \cos \theta).$$

The Riemannian metric tensor of the punctured sphere, when expressed in spherical coordinates, reads

$$(g_{ij}) = \begin{pmatrix} \sin^2 \theta & 0 \\ 0 & 1 \end{pmatrix}.$$

Recall that the inverse matrix of $g = (g_{ij})$ is denoted by (g^{ij}) .

The components of the gradient $\nabla_g u$ of a function $u: S^2 \to \mathbb{R}$ in the basis $(\partial/\partial\varphi, \partial/\partial\theta)$ of T_xS^2 are

$$(\nabla_g u)^{\varphi} = \frac{1}{\sin^2 \theta} \frac{\partial u}{\partial \varphi}, \qquad (\nabla_g u)^{\theta} = \frac{\partial u}{\partial \theta}.$$

According to (2.4), the divergence operator on the sphere is given by

(3.1)
$$\nabla_g \cdot f := \frac{1}{\sin \theta} \left(\frac{\partial}{\partial \varphi} (f^{\varphi} \sin \theta) + \frac{\partial}{\partial \theta} (f^{\theta} \sin \theta) \right),$$

for every smooth vector field f on S^2 , where f^{φ} and f^{θ} are the coordinates of the vector f in the basis $(\partial/\partial\varphi,\partial/\partial\theta)$. Similarly, the Laplace-Beltrami operator is defined for all $u: S^2 \to \mathbb{R}$ by

(3.2)
$$\Delta_g u := \nabla_g \cdot (\nabla_g u) = \frac{1}{\sin \theta} \left(\frac{\partial}{\partial \varphi} \left(\frac{\partial u}{\sin \theta} \frac{\partial u}{\partial \varphi} \right) + \frac{\partial}{\partial \theta} \left(\frac{\partial u}{\partial \theta} \sin \theta \right) \right).$$

We now turn to the following conservation law on the sphere:

(3.3)
$$\partial_t u + \nabla_q \cdot f(u, \cdot) = 0, \quad (t, x) \in \mathbb{R}_+ \times S^2,$$

where the flux $f: \mathbb{R} \times S^2 \to TS^2$ is a smooth vector field. In view of (3.1), when spherical coordinates are used the conservation law (3.3) takes the form

(3.4)
$$\partial_t(u\sin\theta) + \frac{\partial}{\partial\varphi}(f^{\varphi}(u,\varphi,\theta)\sin\theta) + \frac{\partial}{\partial\theta}(f^{\theta}(u,\varphi,\theta)\sin\theta) = 0.$$

A periodic boundary condition is imposed in the variable $\varphi \in [0, 2\pi]$, that is,

(3.5)
$$u(t, 0, \theta) = u(t, 2\pi, \theta), \quad \theta \in [0, \pi], t \ge 0.$$

On the other hand, no boundary condition may be imposed along $\theta = 0$ and $\theta = \pi$. Observe that the equation (3.4) is *singular* (sin $\theta = 0$) precisely at the poles.

If u is a smooth solution of the equation (3.4), a natural question is whether u satisfies some additional conservation laws. We have already noticed that the Total Variation Diminishing property is available for solutions to conservation laws in \mathbb{R}^d when the flux-function is independent of the spatial variable. We have seen also that, when writing the conservation law in local coordinates, the flux in general depends explicitly on the space variables. It would be meaningless to require that the flux is "independent of x". A natural approach is to restrict attention to flux that are divergence-free, which in spherical coordinates is expressed as

(3.6)
$$\frac{\partial}{\partial \varphi} \left(f^{\varphi}(\overline{u}, \varphi, \theta) \sin \theta \right) + \frac{\partial}{\partial \theta} \left(f^{\theta}(\overline{u}, \varphi, \theta) \sin \theta \right) = 0.$$

Recall that an entropy / entropy flux pair for the conservation law (3.1) is a pair (U, F) in which the function $U : \mathbb{R} \to \mathbb{R}$ is arbitrary and $F(u, x) = \int^u U'(\overline{u}) \frac{\partial}{\partial u} f(\overline{u}, x) d\overline{u}$.

For the sake of completeness we recall:

PROPOSITION 3.1. (Intrinsic description.) Consider the conservation law (3.3) on the sphere S^2 with a divergence-free vector field f = f(u, x), and let (U, F) = (U(u), F(u, x)) be any entropy / entropy flux pair. Then, smooth solutions u = u(t, x) of (3.3)-(3.5) satisfy the additional conservation law

(3.7)
$$\partial_t U(u) + \nabla_g \cdot F(u, \cdot) = 0.$$

In particular

(3.8)
$$\frac{d}{dt} \iint_{\Omega} U(u(t,\cdot)) \, dv_g = 0, \quad t \ge 0.$$

Of course, for weak solutions, (3.7)-(3.8) will hold as *inequalities*, only. In spherical coordinates, equation (3.7) reads

$$\partial_t(U(u)\sin\theta) + \frac{\partial}{\partial\varphi}(F^{\varphi}(u,\varphi,\theta)\sin\theta) + \frac{\partial}{\partial\theta}(F^{\theta}(u,\varphi,\theta)\sin\theta) = 0.$$

Proof. Following [1] we multiply (3.3) by U'(u) and get

$$\partial_t(U(u)) + U'(u)\nabla_a \cdot (f(u(t,x),x)) = 0.$$

Using (2.5) with F instead of f, we find

$$\nabla_g \cdot \left(F(u(t,x),x) \right) = (\nabla_g \cdot F)(u,x) + du(F_u(u,x))$$
$$= \int_0^u U'(v)(\nabla_g \cdot f_u)(v,x) \, dv + U'(u) du(f_u(u,x)).$$

Again by (2.5), since f is divergence-free, the first term above vanishes and

$$U'(u)du(f_u(u,x)) = U'(u)\nabla_a \cdot (f(u,x)),$$

so that

$$\nabla_{q} \cdot (F(u(t,x),x)) = U'(u) \nabla_{q} \cdot (f(u,x)),$$

which proves (3.7).

Furthermore, since S^2 has no boundary, the integral of the divergence of any vector field vanishes, thus yielding (3.8).

It is illustrative to also derive the equation (3.7) using spherical coordinates. To simplify the notation we set

$$h^{i} = f^{i} \sin \theta, \quad i = \varphi, \theta; \qquad \partial_{\alpha} = \frac{\partial}{\partial \alpha}, \quad \alpha = \varphi, \theta.$$

We denote here by $h_u^i, h_{\varphi}^i, h_{\theta}^i$ respectively the derivatives of h^i with respect to the first, the second, and the third variables. Then, the equation (3.7) becomes

$$\partial_t(u\sin\theta) + \partial_\varphi(h^\varphi(u,\varphi,\theta)) + \partial_\theta(h^\theta(u,\varphi,\theta)) = 0.$$

Multiplying it by U'(u), we get

$$0 = \partial_t (U(u) \sin \theta) + U'(u) \partial_\varphi (h^\varphi) + U'(u) \partial_\theta (h^\theta)$$

= $\partial_t (U(u) \sin \theta) + U'(u) (h^\varphi_\theta + h^\varphi_u \partial_\varphi u) + U'(u) (h^\theta_\theta + h^\theta_u \partial_\theta u),$

or equivalently, with $H := F \sin \theta$, so that $H_n^i = F_n^i \sin \theta$,

$$\partial_t (U(u) \sin \theta) + H_u^{\varphi} \partial_{\varphi} u + H_u^{\theta} \partial_{\theta} u + U'(u) (h_{\varphi}^{\varphi} + h_{\theta}^{\theta})$$

= $\partial_t (U(u) \sin \theta) + \partial_{\varphi} H^{\varphi} + \partial_{\theta} H^{\theta} + Q = 0,$

where

$$Q = U'(u)(h_{\varphi}^{\varphi} + h_{\theta}^{\theta}) - H_{\varphi}^{\varphi} - H_{\theta}^{\theta}.$$

Suppose now that the condition (3.6) is satisfied, that is, $h_{\varphi}^{\varphi} + h_{\theta}^{\theta} = 0$ for all $u \in \mathbb{R}$. Then, we have

$$H_{\varphi}^{\varphi} + H_{\theta}^{\theta} = \int_{0}^{u} U'(v) (h_{u,\varphi}^{\varphi} + h_{u,\theta}^{\theta}) dv = \int_{0}^{u} U'(v) (h_{\varphi}^{\varphi} + h_{\theta}^{\theta})_{u} dv = 0,$$

and therefore Q=0, which completes the derivation of (3.7) in spherical coordinates.

3.2. Embedded representation of S^2 . In the embedded approach, the sphere S^2 is regarded as a submanifold of the Euclidian space \mathbb{R}^3 , via some isometric embedding $h: S^2 \to \mathbb{R}^3$. The canonical Riemannian metric g on S^2 is induced by the canonical metric \widehat{g} on the Euclidian space \mathbb{R}^3 , by $g = h^*\widehat{g}$, where, by definition, the pull-back h^* maps the 2-tensor field \widehat{g} on \mathbb{R}^3 to the 2-tensor field g on S^2 , as follows:

$$(3.9) (h^*\widehat{g})(\xi,\eta) := \widehat{g}(h_*\xi,h_*\eta), \xi,\eta \in TS^2.$$

The tangent vector $h_{\star}\xi$ is the *push-forward* of $\xi \in TS^2$ into \mathbb{R}^3 :

$$(3.10) (h_{\star}\xi)(f) := \xi(f \circ h) \text{for every function } f : \mathbb{R}^3 \to \mathbb{R}.$$

From now on we fix some local coordinates $(x^i)_{i=1,2}$ on S^2 (for instance, the spherical coordinates described above) as well as some orthonormal coordinates $(\widehat{x}^{\alpha})_{\alpha=1,2,3}$ on \mathbb{R}^3 . In particular, we have $\widehat{g}_{\alpha\beta} = \delta_{\alpha\beta}$ (the Kronecker symbol).

We have

$$(3.11) (h_{\star}\xi)^{\alpha} = \frac{\partial \widehat{x}^{\alpha}}{\partial x^{i}} \xi^{i},$$

so that we may compute the expression of the metric tensor g in these local coordinates:

$$g_{ij}\xi^{i}\eta^{j} = \widehat{g}_{\alpha\beta} (h_{\star}\xi)^{\alpha} (h_{\star}\eta)^{\beta}$$
$$= \frac{\partial \widehat{x}^{\alpha}}{\partial x^{i}} \frac{\partial \widehat{x}_{\alpha}}{\partial x^{j}} \xi^{i} \eta^{j},$$

where $\widehat{x}_{\alpha} := \widehat{g}_{\alpha\beta}\widehat{x}^{\beta}$. From this, it follows that

$$(3.12) g_{ij} = \frac{\partial \widehat{x}^{\alpha}}{\partial x^{i}} \frac{\partial \widehat{x}_{\alpha}}{\partial x^{j}}.$$

In other words, we have $(g_{ij}) = J^t J$, where $J = (\partial \hat{x}^{\alpha}/\partial x^j)$ is the Jacobian matrix of the embedding h.

In what follows, all vectors are considered to be in \mathbb{R}^3 , unless otherwise stated. A vector tangent to the submanifold S^2 of \mathbb{R}^3 is regarded as an element of \mathbb{R}^3 through the identification

$$T_x S^2 \subset T_x \mathbb{R}^3 \simeq \mathbb{R}^3.$$

Consider now the spherical coordinates introduced earlier. The outward unit vector to S^2 in \mathbb{R}^3 at the point $x = (\varphi, \theta)$ is given by

$$h(\varphi, \theta) = \mathbf{n}(x) := (\sin \theta \cos \varphi, \sin \theta \sin \varphi, \cos \theta).$$

At any given point $x = (\varphi, \theta)$, a normalized basis of the tangent plane to S^2 in \mathbb{R}^3 (now viewed as a two-dimensional subspace of \mathbb{R}^3) is given by the two vectors

$$\mathbf{t}_{1}(x) := \frac{1}{\sin \theta} \frac{\partial \mathbf{n}}{\partial \varphi} = \left(-\sin \varphi, \cos \varphi, 0 \right),$$

$$\mathbf{t}_{2}(x) := -\frac{\partial \mathbf{n}}{\partial \theta} = \left(-\cos \theta \, \cos \varphi, -\cos \theta \, \sin \varphi, \sin \theta \right).$$

Given a vector ξ in TS^2 with (intrinsic) spherical coordinates $(\xi^{\varphi}, \xi^{\theta})$, we associate to it the vector $\tilde{\xi} \in \mathbb{R}^3$ defined by (3.10). The components of $\tilde{\xi}$ in the basis $(\mathbf{t}_1, \mathbf{t}_2, \mathbf{n})$ of \mathbb{R}^3 are denoted by $(\tilde{\xi}^i)_{1 \leq i \leq 3}$. The vector $\tilde{\xi}$ is tangent to the sphere: using (3.11) to compute its coordinates, we find

$$\tilde{\xi} = \frac{\partial \mathbf{n}}{\partial \varphi} \xi^{\varphi} + \frac{\partial \mathbf{n}}{\partial \theta} \xi^{\theta} = \sin \theta \, \xi^{\varphi} \mathbf{t}_{1} - \xi^{\theta} \mathbf{t}_{2},$$

so that in spherical coordinates the intrinsic and embedded coordinates of a tangent vector are related by

(3.13)
$$\tilde{\xi}^1 = \sin \theta \, \xi^{\varphi}, \qquad \tilde{\xi}^2 = -\xi^{\theta}, \qquad \tilde{\xi}^3 = 0.$$

Observe that the normal vector is well-defined for all θ (and φ). However, the North ($\theta = 0$) and South ($\theta = \pi$) poles are singularities for the tangent vectors in spherical coordinates. Namely, the particular value of φ is irrelevant when $\theta = 0$ or

 π ; however, the tangent vectors to S^2 do depend upon φ when $\theta=0$ or π , so that a basis of the tangent space is not uniquely specified at the poles. This is consistent with the well-known fact that any globally smooth, tangent vector field on the sphere must vanish at least once.

In (embedded) spherical coordinates, the flux $f: \mathbb{R} \times \Omega \to \mathbb{R}^3$ has the general form

$$f(u,\varphi,\theta) =: \tilde{f}^1(u,\varphi,\theta) \mathbf{t}_1(\varphi,\theta) + \tilde{f}^2(u,\varphi,\theta) \mathbf{t}_2(\varphi,\theta).$$

Note that f is tangent to the sphere. In view of (3.1), (3.13), we have

$$\tilde{f}^1 = \sin\theta \, f^{\varphi}, \quad \tilde{f}^2 = -f^{\theta},$$

and the divergence operator takes the form

$$\nabla_g \cdot f := \frac{1}{\sin \theta} \left(\frac{\partial \tilde{f}^1}{\partial \varphi} - \frac{\partial}{\partial \theta} (\tilde{f}^2 \sin \theta) \right).$$

In terms of the unknown function $u = u(t, \varphi, \theta)$, (3.3) is a conservation law with non-constant coefficients posed in the rectangular domain $(t, \varphi, \theta) \in \mathbb{R}_+ \times \Omega$:

(3.14)
$$\partial_t (u \sin \theta) + \frac{\partial}{\partial \varphi} (\tilde{f}^1(u, \varphi, \theta)) - \frac{\partial}{\partial \theta} (\tilde{f}^2(u, \varphi, \theta) \sin \theta) = 0.$$

For the sake of clarity we will always state the final results in the intrinsic notation f^{φ} , f^{θ} , but in the course of the calculation it will be convenient to use the embedded notation \tilde{f}^1 , \tilde{f}^2 , which has the advantage to be based on an orthonormal basis of \mathbb{R}^3 .

Let us now define the tubular neighborhood $T \subset \mathbb{R}^3$ of S^2 consisting of all spheres with radius 1/2 < r < 3/2, so that S^2 corresponds to r = 1. Then, any smooth tangent vector field f(u, x), $x \in S^2$, may be represented at each point of the sphere by

(3.15)
$$f(u,x) = \mathbf{n}(x) \times \Phi(u,\mathbf{n}(x)),$$

where $\Phi(u, x)$, $x \in T$ is a smooth vector field in $T \subset \mathbb{R}^3$. We can thus write $\Phi(u, x) = \Phi(u, \mathbf{n}(x))$.

Lemma 3.2. With the notation (3.15) the equation (3.14) takes the form

$$(3.16) \qquad \partial_t \left(u \sin \theta \right) + \frac{\partial}{\partial \varphi} \left(\Phi(u, \mathbf{n}(\varphi, \theta)) \cdot \frac{\partial \mathbf{n}}{\partial \theta} \right) - \frac{\partial}{\partial \theta} \left(\Phi(u, \mathbf{n}(\varphi, \theta)) \cdot \frac{\partial \mathbf{n}}{\partial \varphi} \right) = 0,$$

Proof. Setting

$$A := \begin{pmatrix} \sin \theta \cos \varphi & -\sin \varphi & -\cos \theta \cos \varphi \\ \sin \theta \sin \varphi & \cos \varphi & -\cos \theta \sin \varphi \\ \cos \theta & 0 & \sin \theta \end{pmatrix}$$

we have

$$\mathbf{n} = A \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix}, \quad \mathbf{t}_1 = A \begin{pmatrix} 0 \\ 1 \\ 0 \end{pmatrix}, \quad \mathbf{t}_2 = A \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix},$$

and $A^{-1} = A^t$. Since det(A) = 1 the basis $(\mathbf{t}_1, \mathbf{t}_2, \mathbf{n})$ is positively oriented and

$$\mathbf{n} \times \mathbf{t}_1 = \mathbf{t}_2, \qquad \mathbf{n} \times \mathbf{t}_2 = -\mathbf{t}_1.$$

Using coordinates, we have

$$\tilde{\Phi}^i(u, \mathbf{n}(\varphi, \theta)) := \Phi(u, \mathbf{n}(\varphi, \theta)) \cdot \mathbf{t}_i(\varphi, \theta)$$

and

$$f(u,\varphi,\theta) = \mathbf{n}(\varphi,\theta) \times \left(\tilde{\Phi}^{1}(u,\mathbf{n}(\varphi,\theta)) \mathbf{t}_{1}(\varphi,\theta) + \tilde{\Phi}^{2}(u,\mathbf{n}(\varphi,\theta)) \mathbf{t}_{2}(\varphi,\theta)\right)$$
$$= \tilde{\Phi}^{1}(u,\mathbf{n}(\varphi,\theta)) \mathbf{t}_{2}(\varphi,\theta) - \tilde{\Phi}^{2}(u,\mathbf{n}(\varphi,\theta)) \mathbf{t}_{1}(\varphi,\theta),$$

so that $\tilde{f}^1 = -\tilde{\Phi}^2$, $\tilde{f}^2 = \tilde{\Phi}^1$. But

$$\tilde{\Phi}^1 = \frac{1}{\sin \theta} \Big(\Phi \cdot \frac{\partial \mathbf{n}}{\partial \varphi} \Big), \quad \tilde{\Phi}^2 = - \Big(\Phi \cdot \frac{\partial \mathbf{n}}{\partial \theta} \Big),$$

and so

$$\tilde{f}^1 = \left(\Phi \cdot \frac{\partial \mathbf{n}}{\partial \theta}\right), \quad \tilde{f}^2 = \frac{1}{\sin \theta} \left(\Phi \cdot \frac{\partial \mathbf{n}}{\partial \varphi}\right),$$

or equivalently using the intrinsic notation

$$f^{\varphi} = \frac{1}{\sin \theta} \Big(\Phi \cdot \frac{\partial \mathbf{n}}{\partial \theta} \Big), \quad f^{\theta} = -\frac{1}{\sin \theta} \Big(\Phi \cdot \frac{\partial \mathbf{n}}{\partial \varphi} \Big),$$

which, in view of (3.14), leads to (3.16). \square

Observe that, for smooth solutions, (3.16) takes the equivalent form

(3.17)
$$\partial_t (u \sin \theta) + \partial_u \Phi(u, \mathbf{n}(\varphi, \theta)) \cdot \left(\frac{\partial u}{\partial \varphi} \frac{\partial \mathbf{n}}{\partial \theta} - \frac{\partial u}{\partial \theta} \frac{\partial \mathbf{n}}{\partial \varphi} \right) = 0.$$

We have the analogue of Proposition 3.1:

PROPOSITION 3.3. (Embedded description.) Suppose that the vector field Φ satisfies the divergence-free condition: for all $\overline{u} \in \mathbb{R}$ and $(\varphi, \theta) \in \Omega$,

$$(3.18) \qquad \frac{\partial}{\partial \varphi} \Big(\Phi(\overline{u}, \mathbf{n}(\varphi, \theta)) \cdot \frac{\partial \mathbf{n}}{\partial \theta} (\varphi, \theta) \Big) - \frac{\partial}{\partial \theta} \Big(\Phi(\overline{u}, \mathbf{n}(\varphi, \theta)) \cdot \frac{\partial \mathbf{n}}{\partial \varphi} (\varphi, \theta) \Big) = 0$$

or, equivalently,

$$\frac{\partial \Phi}{\partial \varphi}(\overline{u}) \cdot \frac{\partial \mathbf{n}}{\partial \theta} - \frac{\partial \Phi}{\partial \theta}(\overline{u}) \cdot \frac{\partial \mathbf{n}}{\partial \varphi} = 0.$$

Let $U: \mathbb{R} \to \mathbb{R}$ and $\Psi: \mathbb{R} \times S^2 \to \mathbb{R}^3$ be smooth maps such that

(3.20)
$$\Psi(u, \mathbf{n}(\varphi, \theta)) = \int_0^u U'(\overline{u}) \frac{\partial}{\partial u} \Phi(\overline{u}, \mathbf{n}(\varphi, \theta)) d\overline{u}.$$

Then every smooth solution u of (3.16) satisfies

$$(3.21) \quad \partial_t \left(U(u) \sin \theta \right) + \frac{\partial}{\partial \varphi} \left(\Psi(u, \mathbf{n}(\varphi, \theta)) \cdot \frac{\partial \mathbf{n}}{\partial \theta} \right) - \frac{\partial}{\partial \theta} \left(\Psi(u, \mathbf{n}(\varphi, \theta)) \cdot \frac{\partial \mathbf{n}}{\partial \varphi} \right) = 0.$$

Proof. Multiplying the equation (3.16) by U'(u) we obtain

$$\partial_t (U(u) \sin \theta) - \partial_u \Psi(u, \mathbf{n}(\varphi, \theta)) \cdot \partial_\theta u \frac{\partial \mathbf{n}}{\partial \varphi} + \partial_u \Psi(u, \mathbf{n}(\varphi, \theta)) \cdot \partial_\varphi u \frac{\partial \mathbf{n}}{\partial \theta} = 0,$$

thus

$$\partial_t (U(u) \sin \theta) - \partial_\theta (\Psi(u, \mathbf{n}(\varphi, \theta)) \cdot \frac{\partial \mathbf{n}}{\partial \varphi}) + \partial_\varphi (\Psi(u, \mathbf{n}(\varphi, \theta)) \cdot \frac{\partial \mathbf{n}}{\partial \theta}) + Q = 0,$$

where

$$Q := \Psi_{\theta} \cdot \frac{\partial \mathbf{n}}{\partial \varphi} - \Psi_{\varphi} \cdot \frac{\partial \mathbf{n}}{\partial \theta}$$

$$= \int^{u} U'(v) \, \Phi_{u\theta} \, dv \frac{\partial \mathbf{n}}{\partial \varphi} - \int^{u} U'(v) \Phi_{u\varphi} \, dv \frac{\partial \mathbf{n}}{\partial \theta}$$

$$= \int^{u} U'(v) \Big(\Phi_{\theta} \cdot \frac{\partial \mathbf{n}}{\partial \varphi} - \Phi_{\varphi} \cdot \frac{\partial \mathbf{n}}{\partial \varphi} \Big)_{u} \, dv = 0,$$

by the condition (3.19).

Remark 3.4. The constraint (3.18) can also be written in the form

$$\left(D\Phi \cdot \frac{\partial \mathbf{n}}{\partial \varphi}\right) \cdot \frac{\partial \mathbf{n}}{\partial \theta} - \left(D\Phi \cdot \frac{\partial \mathbf{n}}{\partial \theta}\right) \cdot \frac{\partial \mathbf{n}}{\partial \varphi} = 0.$$

In particular, if the map Φ coincides with the gradient of some potential h, that is, $\Phi = Dh$ for $h: T \to \mathbb{R}$, then the divergence-free condition (3.18) is automatically satisfied.

4. Total variation diminishing estimates on the sphere S^2 . We are now in a position to discuss the total variation of solutions to scalar conservation laws on the sphere. We are going to investigate the conditions discovered in Theorem 2.3. After comparing the intrinsic and embedded description it turned out that the second approach led to simpler (but still not so straightforward) and more intuitive formulas. In view of Section 3, a vector field tangent to the embedded sphere S^2 can be expressed as the cross product between the exterior normal $\mathbf{n}(x)$ and some other vector field which can always be chosen to be tangent to the sphere. So, given arbitrary flux f and vector field X let us introduce (smooth) maps $\Phi : \mathbb{R} \times T \to \mathbb{R}^3$ and $\Psi : T \to \mathbb{R}^3$ (which are not unique) such that

(4.1)
$$f(u, \varphi, \theta) = \mathbf{n}(\varphi, \theta) \times \Phi(u, \mathbf{n}(\varphi, \theta))$$
$$X(\varphi, \theta) = \mathbf{n}(\varphi, \theta) \times \Psi(\mathbf{n}(\varphi, \theta))$$

(where $T \subset \mathbb{R}^3$ is the tubular neighborhood of S^2 defined in Section 3). It should be kept in mind that all the vectors under consideration here belong to \mathbb{R}^3 .

Having in mind the class of divergence-free flux, we introduce:

DEFINITION 4.1. The flux f and vector field X are said to satisfy the **gradient** condition if there exist functions $a: \mathbb{R} \times \mathbb{R}^3 \to \mathbb{R}$ and $b: \mathbb{R}^3 \to \mathbb{R}$ depending solely on $\mathbf{n} = x/|x|$ and smooth in a neighborhood of the 2-sphere such that

(4.2)
$$\Phi(u, x) = D_x a(u, x/|x|), \qquad \Psi(x) = D_x b(x/|x|), \qquad x \in T.$$

From the gradient condition we deduce that, for instance,

$$\Phi = D_x a = D_{\mathbf{n}} a \cdot D_x (x/|x|).$$

Clearly, $D_x(x/|x|) \cdot \mathbf{n}$ vanishes identically, since it is the derivative of the map x/|x| in the direction of the vector \mathbf{n} . Therefore,

$$\Phi \cdot \mathbf{n} = \Psi \cdot \mathbf{n} = 0,$$

that is, the vectors Φ and Ψ are tangent to S^2 . Observe also that as in Remark 3.6, the divergence-free condition is automatically satisfied for fields satisfying the gradient condition.

Here is the main result of this section.

THEOREM 4.2. Let f and X be a flux and a vector field defined on the sphere S^2 respectively, satisfying the gradient condition (4.1)-(4.2) for some maps Φ, Ψ . Then, the flux f is divergence-free and one has

(4.3)
$$\mathcal{L}_X f_u(\overline{u}, \cdot) = 0$$
 if and only if $f_u(\overline{u}, \cdot) = C(\overline{u}, \cdot) X$, $\overline{u} \in \mathbb{R}$

for some scalar-valued flux-function $C = C(\overline{u}, \cdot)$ satisfying $X(C(\overline{u}, \cdot)) = 0$.

When (4.3) holds, and when $(\nabla_g \cdot f)(u, \cdot) = 0$, the conservation law (3.14) decouples into a one-parameter family of independent, one-dimensional conservation laws.

Hence, in the case of flux which are "gradients", we are able to identify the class of flux satisfying the condition of Section 2. For each \overline{u} , the vectors $f_u(\overline{u},\cdot)$ and X are parallel and $C(\overline{u},\cdot)$ is constant along the integral curves of the vector field X. The curved geometry of the sphere restrict the class of flux enjoying the TVD property. This should be compared with the case of equations in the Euclidian space \mathbb{R}^d where no such condition is implied on the flux.

Throughout the rest of this section, a value $u \in \mathbb{R}$ is fixed and so we omit it. Recall from Section 3 that the components $(\tilde{f}^1, \tilde{f}^2)$ of f in the basis $(\mathbf{t}_1, \mathbf{t}_2)$ of the tangent plane to S^2 at a point $x(\varphi, \theta)$ read:

$$\tilde{f}^1 = -\Phi(\mathbf{n}) \cdot \mathbf{n}_{\theta}, \quad \tilde{f}^2 = \frac{1}{\sin \theta} \Phi(\mathbf{n}) \cdot \mathbf{n}_{\varphi},$$

where we have used subscripts to indicate differentiation. Using the equation (3.13) which connects these components with the intrinsic components $(f^{\varphi}, f^{\theta})$ of f, we find

$$f^{\varphi} = \frac{1}{\sin \theta} \Phi(\mathbf{n}) \cdot \mathbf{n}_{\theta}$$
$$f^{\theta} = -\frac{1}{\sin \theta} \Phi(\mathbf{n}) \cdot \mathbf{n}_{\varphi}.$$

A similar formula holds for X:

$$X^{\varphi} = \frac{1}{\sin \theta} \Psi(\mathbf{n}) \cdot \mathbf{n}_{\theta}$$
$$X^{\theta} = -\frac{1}{\sin \theta} \Psi(\mathbf{n}) \cdot \mathbf{n}_{\varphi}.$$

Observe now the following elementary identities:

(4.4)
$$\mathbf{n}_{\varphi\varphi} = -\sin^2 \theta \, \mathbf{n} - \sin \theta \cos \theta \, \mathbf{n}_{\theta},$$
$$\mathbf{n}_{\varphi\theta} = \frac{\cos \theta}{\sin \theta} \mathbf{n}_{\varphi}$$
$$\mathbf{n}_{\theta\theta} = -\mathbf{n}.$$

We are now in a position to reformulate the condition $\mathcal{L}_X f_u = 0$.

LEMMA 4.3. The condition $\mathcal{L}_X f_u = 0$ is equivalent to the requirement that $\Psi = \Psi(\mathbf{n})$ satisfies the following system of partial differential equations with unknown X:

$$(\Phi_{u} \cdot \mathbf{n}_{\theta})(\Psi_{\varphi} \cdot \mathbf{n}_{\theta}) - (\Phi_{u} \cdot \mathbf{n}_{\varphi})(\Psi_{\theta} \cdot \mathbf{n}_{\theta})$$

$$= -(\Psi \cdot \mathbf{n}_{\varphi})((\Phi_{u\theta} \cdot \mathbf{n}_{\theta}) - (\Phi_{u} \cdot \mathbf{n}))$$

$$-(\Psi \cdot \mathbf{n})(\Phi_{u} \cdot \mathbf{n}_{\varphi}) + (\Psi \cdot \mathbf{n}_{\theta})(\Phi_{u\varphi} \cdot \mathbf{n}_{\theta}),$$

$$(\Phi_{u} \cdot \mathbf{n}_{\theta})(\Psi_{\varphi} \cdot \mathbf{n}_{\varphi}) - (\Phi_{u} \cdot \mathbf{n}_{\varphi})(\Psi_{\theta} \cdot \mathbf{n}_{\varphi})$$

$$= (\Psi \cdot \mathbf{n}_{\theta})((\Phi_{u\varphi} \cdot \mathbf{n}_{\varphi}) - \sin^{2}\theta(\Phi_{u} \cdot \mathbf{n}))$$

$$+ \sin^{2}\theta(\Psi \cdot \mathbf{n})(\Phi_{u} \cdot \mathbf{n}_{\theta}) - (\Psi \cdot \mathbf{n}_{\varphi})(\Phi_{u\theta} \cdot \mathbf{n}_{\varphi}).$$

Proof. The condition on the first component, $(\mathcal{L}_X f_u)^{\varphi}$, of the vector $\mathcal{L}_X f_u$ reads

$$(X^{\varphi}\partial_{\varphi} + X^{\theta}\partial_{\theta})f_{u}^{\varphi} - (f_{u}^{\varphi}\partial_{\varphi} + f_{u}^{\theta}\partial_{\theta})X^{\varphi} = 0,$$

that is

$$\frac{1}{\sin \theta} (\Psi \cdot \mathbf{n}_{\theta}) \partial_{\varphi} (\Phi_{u} \cdot \mathbf{n}_{\theta}) - (\Psi \cdot \mathbf{n}_{\varphi}) \partial_{\theta} \left(\frac{1}{\sin \theta} (\Phi_{u} \cdot \mathbf{n}_{\theta}) \right) \\
- \frac{1}{\sin \theta} (\Phi \cdot \mathbf{n}_{\theta}) \partial_{\varphi} (\Psi \cdot \mathbf{n}_{\theta}) + (\Phi \cdot \mathbf{n}_{\varphi}) \partial_{\theta} \left(\frac{1}{\sin \theta} (\Psi \cdot \mathbf{n}_{\theta}) \right) = 0.$$

This equation takes the form

$$(\Psi \cdot \mathbf{n}_{\theta})(\Phi_{u\varphi} \cdot \mathbf{n}_{\theta}) - (\Psi \cdot \mathbf{n}_{\varphi})(\Phi_{u\varphi} \cdot \mathbf{n}_{\theta}) + \frac{\cos \theta}{\sin \theta}(\Psi \cdot \mathbf{n}_{\varphi})(\Phi_{u} \cdot \mathbf{n}_{\theta})$$
$$- (\Phi_{u} \cdot \mathbf{n}_{\theta})(\Psi_{\varphi} \cdot \mathbf{n}_{\theta}) + (\Phi_{u} \cdot \mathbf{n}_{\varphi})(\Psi_{\theta} \cdot \mathbf{n}_{\theta}) - \frac{\cos \theta}{\sin \theta}(\Phi_{u} \cdot \mathbf{n}_{\varphi})(\Psi \cdot \mathbf{n}_{\theta})$$
$$+ (\Psi \cdot \mathbf{n}_{\theta})(\Phi_{u} \cdot \mathbf{n}_{\varphi\theta}) - (\Psi \cdot \mathbf{n}_{\varphi})(\Phi_{u} \cdot \mathbf{n}_{\theta\theta})$$
$$- (\Phi_{u} \cdot \mathbf{n}_{\theta})(\Psi \cdot \mathbf{n}_{\theta\varphi}) + (\Phi_{u} \cdot \mathbf{n}_{\varphi})(\Psi \cdot \mathbf{n}_{\theta\theta}) = 0.$$

In view of (4.4), all terms above which contain the factor $(\cos \theta / \sin \theta)$ cancel out. Using that $\mathbf{n}_{\theta\theta} = -\mathbf{n}$ and re-arranging the terms leads to the first equation.

The second equation is obtained similarly, by developing

$$(\mathcal{L}_X f_u)^{\theta} = (X^{\varphi} \partial_{\varphi} + X^{\theta} \partial_{\theta}) f_u^{\theta} - (f_u^{\varphi} \partial_{\varphi} + f_u^{\theta} \partial_{\theta}) X^{\theta}$$

and using (4.4). \square

We now show:

LEMMA 4.4. A point $\overline{x} \in S^2$ being fixed, there exist non-trivial flux f and vector fields X satisfying the gradient condition, and possibly defined in a small neighborhood of \overline{x} only, such that $\mathcal{L}_X f_u = 0$. These vector fields are characterized explicitly by the condition

$$X \times f_u = \widetilde{C} \mathbf{n}$$

for some scalar $\widetilde{C} = \widetilde{C}(u)$ (independent of x). Among these solutions, the only flux f and vector field X that are globally defined and smooth on the whole sphere S^2 are given by

$$f_u(u,\cdot) = C(u,\cdot) X$$

where $C = C(\cdot)$ is a (scalar-valued) function such that $X(C(u, \cdot)) = 0$.

Vector fields that are defined locally on S^2 may still be useful to control the total variation in a domain of dependence, but they do not provide information on the large-time behavior of solutions.

Proof. First, let us show that for $i = \varphi, \theta$ we have $a_i = \Phi \cdot \mathbf{n}_i$ and $b_i = \Psi \cdot \mathbf{n}_i$ (subscripts indicate differentiation). A straightforward calculation shows that

$$D_x a = D_{\mathbf{n}} a \cdot D_x \left(\frac{x}{|x|} \right) = D_{\mathbf{n}} a \cdot \left(\frac{\operatorname{Id} - \mathbf{n} \otimes \mathbf{n}}{|x|} \right).$$

In particular, restricting the formula to the sphere we find $D_x a = D_{\mathbf{n}} a \cdot (\mathrm{Id} - \mathbf{n} \otimes \mathbf{n})$, and therefore

$$a_i = D_{\mathbf{n}} a \cdot \mathbf{n}_i = D_{\mathbf{n}} a \cdot (\mathrm{Id} - \mathbf{n} \otimes \mathbf{n}) + D_{\mathbf{n}} a \cdot (\mathbf{n} \otimes \mathbf{n}) \cdot \mathbf{n}_i = D_x a \cdot \mathbf{n}_i = \Phi \cdot \mathbf{n}_i$$

because $(\mathbf{n} \otimes \mathbf{n}) \cdot \mathbf{n}_i = 0$. In consequence, we now simply write Da, Db. Recalling that by assumption we have $Da \cdot \mathbf{n} = Db \cdot \mathbf{n} = 0$, we can rewrite (4.5) in terms of b:

$$(4.6) \qquad a_{u\theta}(\partial_{\varphi}Db \cdot \mathbf{n}_{\theta}) - a_{u\varphi}(\partial_{\theta}Db \cdot \mathbf{n}_{\theta}) = -b_{\varphi}(\partial_{\theta}Da_{u} \cdot \mathbf{n}_{\theta}) + b_{\theta}(\partial_{\varphi}Da_{u} \cdot \mathbf{n}_{\theta}),$$

$$a_{u\theta}(\partial_{\varphi}Db \cdot \mathbf{n}_{\varphi}) - a_{u\varphi}(\partial_{\theta}Db \cdot \mathbf{n}_{\varphi}) = b_{\theta}(\partial_{\varphi}Da_{u} \cdot \mathbf{n}_{\varphi}) - b_{\varphi}(\partial_{\theta}Da_{u} \cdot \mathbf{n}_{\varphi}).$$

This is now a system in the scalar variable b. Furthermore, using (4.4) we find

$$a_{u\theta\theta} = \partial_{\theta}(Da_{u} \cdot \mathbf{n}_{\theta}) = \partial_{\theta}Da_{u} \cdot \mathbf{n}_{\theta} - Da_{u} \cdot \mathbf{n} = \partial_{\theta}Da_{u} \cdot \mathbf{n}_{\theta},$$

$$a_{u\varphi\varphi} = \partial_{\varphi}Da_{u} \cdot \mathbf{n}_{\varphi} + Da_{u} \cdot \mathbf{n}_{\varphi\varphi} = \partial_{\varphi}Da_{u} \cdot \mathbf{n}_{\varphi} - \sin\theta\cos\theta \, a_{u\theta},$$

$$a_{u\theta\varphi} = \partial_{\varphi}Da_{u} \cdot \mathbf{n}_{\theta} + Da_{u} \cdot \mathbf{n}_{\varphi\theta} = \partial_{\varphi}Da_{u} \cdot \mathbf{n}_{\theta} + \frac{\cos\theta}{\sin\theta}Da_{u} \cdot \mathbf{n}_{\varphi} = \partial_{\varphi}Da_{u} \cdot \mathbf{n}_{\theta} + \frac{\cos\theta}{\sin\theta}a_{u\varphi},$$

so that

$$\partial_{\theta} D a_{u} \cdot \mathbf{n}_{\theta} = a_{u\theta\theta},$$

$$\partial_{\varphi} D a_{u} \cdot \mathbf{n}_{\varphi} = a_{u\varphi\varphi} + \sin\theta \cos\theta \, a_{u\theta},$$

$$\partial_{\varphi} D a_{u} \cdot \mathbf{n}_{\theta} = a_{u\theta\varphi} - \frac{\cos\theta}{\sin\theta} a_{u\varphi}.$$

Similarly, we obtain

$$\partial_{\theta} Db \cdot \mathbf{n}_{\theta} = b_{\theta\theta}$$

$$\partial_{\varphi} Db \cdot \mathbf{n}_{\varphi} = b_{\varphi\varphi} + \sin\theta \cos\theta \, b_{\theta}$$

$$\partial_{\varphi} Db \cdot \mathbf{n}_{\theta} = b_{\theta\varphi} - \frac{\cos\theta}{\sin\theta} b_{\varphi}$$

and, therefore, the first equation of the system (4.6) becomes

$$b_{\varphi} a_{u\theta\theta} + b_{\varphi\theta} a_{u\theta} - a_{u\varphi} b_{\theta\theta} - b_{\theta} a_{u\varphi\theta} = \frac{\cos \theta}{\sin \theta} (a_{u\theta} b_{\varphi} - b_{\theta} a_{u\varphi}).$$

This is equivalent to saying

$$(a_{u\theta} b_{\varphi} - b_{\theta} a_{u\varphi})_{\theta} = \frac{\cos \theta}{\sin \theta} (a_{u\theta} b_{\varphi} - b_{\theta} a_{u\varphi}),$$

which, after integration, leads to

$$a_{u\theta} b_{\varphi} - b_{\theta} a_{u\varphi} = \sin \theta e^{G(\varphi)}$$

for some function $G = G(\varphi)$.

As for the second equation in (4.6), we have

$$a_{u\theta} b_{\varphi\varphi} + b_{\varphi} a_{u\theta\varphi} - b_{\theta} a_{u\varphi\varphi} - a_{u\varphi} b_{\varphi\theta} + \sin\theta \cos\theta a_{u\theta} b_{\theta} - \sin\theta \cos\theta b_{\theta} a_{u\theta}$$
$$= \frac{\cos\theta}{\sin\theta} (a_{u\varphi} b_{\varphi} - b_{\varphi} a_{u\varphi}),$$

so that

$$\left(a_{u\theta}\,b_{\varphi} - b_{\theta}\,a_{u\varphi}\right)_{\varphi} = 0$$

and $a_{u\theta} b_{\varphi} - b_{\theta} a_{u\varphi}$ is a function of θ alone. Therefore, $G = G(\varphi)$ is a constant and we find

$$a_{u\theta} b_{\varphi} - b_{\theta} a_{u\varphi} = C \sin \theta$$

for some scalar C = C(u).

The condition above can be written in terms of the original components of f_u and X, yielding

$$f_u^{\theta} X^{\varphi} - f_u^{\varphi} X^{\theta} = \frac{C}{\sin \theta}$$

and, for the embedded components,

$$\tilde{f}_u^2 \tilde{X}^1 - \tilde{f}_u^1 \tilde{X}^2 = C.$$

Since X and f_u are both tangent to the sphere, the last equation means that there exists a constant C such that

$$X \times f_u = C \mathbf{n}$$

for all $x \in S^2$, so

$$|X \times f_u| = C, \quad x \in S^2.$$

This formula describes vector fields that satisfy the constraint but may not be globally smooth.

Imposing now that X is smooth, we deduce that it must vanish at least once on the sphere, as must the flux f_u . By continuity, we necessarily have C = 0 so that f_u and X are colinear. This means that there exists a function C(u, x) such that $f_u = CX$ for all $x \in S^2$. Finally, substituting f_u for CX in $\mathcal{L}_X f_u = 0$ yields X(C) = 0. This completes the proof of Lemma 4.4. \square

Proof of Theorem 4.2. The first claim of the theorem is an immediate consequence of Lemma 4.4.

The equations are written along any integral curve of the vector field X. To see this, fix a point $x \in S^2$ and introduce a parametrization $s \mapsto \varphi(s)$ of the (unique, integral) curve passing through x and such that the velocity vector $\frac{d}{ds}\varphi(s)$ equals the vector X for every s. Then define the function

$$\tilde{u}(t,s) := u(t,\varphi(s)), \qquad t \ge 0, \quad s \in \mathbb{R}.$$

Note that since X(C) = 0 we have $C(\tilde{u}(t, s), \varphi(s)) =: \tilde{C}(\tilde{u})$ for all s. Therefore, using (4.3) and (2.5), we deduce that

$$\partial_s \widetilde{C}(\widetilde{u}(t,s)) = \widetilde{C}_u(\widetilde{u}(t,s))X(\varphi(s))(u) = du(\widetilde{C}_u(u(t,\varphi(s)))X(\varphi(s)))$$
$$= \nabla_q \cdot (f(u(t,\varphi(s)),\varphi(s))) = -\partial_t u(t,\varphi(s)) = -\partial_t \widetilde{u}(t,s),$$

thus

$$\partial_t u(t,s) + \partial_s \widetilde{C}(u(t,s)) = 0.$$

Note that this is a *one-dimensional* conservation law in which the flux \tilde{C} does not explicitly depend on the spatial coordinate. By Theorem 2.3, these conservation laws admit total variation bounds of the form (2.3), provided that f is divergence-free. \square

5. Finite volume method. The rest of this paper is devoted to establishing the convergence of the finite volume method on a d-dimensional Riemannian manifold (M, g). We will construct approximate solutions to the Cauchy problem associated with the conservation law (1.1) and the initial condition $(u_0 \in L^1(M) \cap L^{\infty}(M))$

$$(5.1) u(0,x) = u_0(x), \quad x \in M,$$

using a finite volume method based on monotone numerical flux-functions, and we will establish that such approximations converge strongly to the unique entropy solution of the problem. We need not assume that the flux f is divergence free (see (1.2)), and on the map $\nabla_g \cdot f$ we require the following growth condition

$$(5.2) |\nabla_g \cdot f(u, x)| \le C_1 + C_2 |u|, (u, x) \in \mathbb{R} \times M,$$

for some constants $C_1, C_2 > 0$.

Let us introduce a triangulation \mathcal{T}^h of M, that is, a set of curved polyhedra $K\subset M$ whose vertices are joined by geodesic lines. We assume that $M=\bigcup_{K\in\mathcal{T}^h}\overline{K}$ and that for every distinct elements $K_1,K_2\in\mathcal{T}^h$, the set $K_1\cap K_2$ is either a common face of K_1,K_2 or else a submanifold with dimension at most (d-2). To each element K we associate the set ∂K of its faces e, and we denote by K_e the unique element distinct from K sharing the face e with K. The outward unit normal to an element K at some point K0 is denoted by $\mathbf{n}_{e,K}(K)$ 1 is \mathbf{n} 2. Finally, K3 and K4 and K5 and K6 and K7 is a submanifold of K8. When K9 is a submanifold of K9, we denote by K9 the volume element on K9 induced by the metric K9.

We impose that

$$h := \sup_{\substack{x,y \in K \\ K \in \mathcal{T}^h}} d_g(x,y) \to 0,$$

where $d_g(\cdot, \cdot)$ is the distance function associated to the metric g. In particular, elements may become flat in the limit as h tends to zero. Interestingly enough, this condition on h is slightly more general than the corresponding condition in [2], where the exterior diameter of the elements was considered. We denote the time increment of the discretization by $\tau = \tau(h)$, and set $t_n := \tau n$ for $n = 0, 1, \ldots$ We also assume that, as h tends to zero,

(5.4)
$$\tau \to 0, \qquad h^2/\tau \to 0.$$

To each element K and each edge $e \in \partial K$ we associate a family of locally Lipschitz continuous, numerical flux-functions $f_{e,K} : \mathbb{R} \times \mathbb{R} \to \mathbb{R}$ satisfying the following properties.

• Consistency property:

(5.5)
$$f_{e,K}(u,u) = \frac{1}{|e|} \int_{e} g(f(u,x), \mathbf{n}_{e,K}(x)) dv_{e}, \quad u \in \mathbb{R}.$$

• Conservation property:

$$(5.6) f_{e,K}(u,v) = -f_{e,K_e}(v,u), u,v \in \mathbb{R}.$$

• Monotonicity property:

(5.7)
$$\frac{\partial}{\partial u} f_{e,K} \ge 0, \quad \frac{\partial}{\partial v} f_{e,K} \le 0.$$

Note that, in the Euclidian case and when the flux depends only on the conservative variable, the consistency condition (5.5) reads $f_{e,K}(\overline{u},\overline{u}) = f(\overline{u}) \cdot \mathbf{n}_{e,K}$, which is the usual consistency condition (cf. [2]). For the sake of stability we impose the following *CFL condition*, for every element $K \in \mathcal{T}$:

(5.8)
$$\frac{\tau p_K}{|K|} \max_{K \in \mathcal{T}^h, e \in \partial K} \sup_{u, v \in \mathbb{R}, u \neq v} \left| \frac{f_{e,K}(u, v) - f_{e,K}(v, v)}{u - v} \right| < 1,$$
$$\frac{\tau p_K}{|K|} \operatorname{Lip}(f) < 1,$$

where

$$\operatorname{Lip}(f) := \sup_{\substack{x \in M \\ \overline{u} \in \mathbb{R}}} \left| \frac{df}{du}(\overline{u}, x) \right|_g$$

and $|\cdot|_g$ represents the norm induced on each tangent space T_xM by the metric g. As in the Euclidean case, the finite volume method can be motivated by formally averaging the conservation law (1.1) over an element K and using Stokes' theorem:

$$0 = \frac{d}{dt} \frac{1}{|K|} \int_{K} u \, dv_{M} + \frac{1}{|K|} \int_{K} \nabla_{g} \cdot f(u, x) \, dv_{M}$$

$$= \frac{d}{dt} \frac{1}{|K|} \int_{K} u \, dv_{M} + \frac{1}{|K|} \int_{\partial K} g(f(u, x), \mathbf{n}(x)) \, dv_{\partial K}$$

$$= \frac{d}{dt} \frac{1}{|K|} \int_{K} u \, dv_{M} + \frac{1}{|K|} \sum_{g \in \partial K} \int_{e} g(f(u, x), \mathbf{n}_{e, K}(x)) \, dv_{e}.$$

We then introduce the approximation

$$u_K^n \approx \frac{1}{|K|} \int_K u \, dv_M, \quad f_{e,K}(u_K^n, u_{K_e}^n) \approx \frac{1}{|e|} \int_{\mathbb{R}} g(f(u, x), \mathbf{n}_{e,K}(x)) \, dv_e$$

and discretize the time derivative with a two-point scheme.

Based on the above the finite volume approximations are defined by

(5.9)
$$u_K^{n+1} := u_K^n - \sum_{e \in \partial K} \frac{\tau|e|}{|K|} f_{e,K}(u_K^n, u_{K_e}^n).$$

The initial condition (5.1) gives us

(5.10)
$$u_K^0 := \frac{1}{|K|} \int_K u_0(x) \, dv_M.$$

Finally we define the function $u^h: \mathbb{R}_+ \times M \to \mathbb{R}$ by

(5.11)
$$u^h(t,x) = u_K^n, \quad x \in K, \quad t_n \le t \le t_{n+1}.$$

We can now state our main convergence result.

Theorem 5.1. Consider the Cauchy problem (1.1)-(5.1) associated with a conservation law on the Riemannian manifold (M,g) and some initial data $u_0 \in L^1(M) \cap L^{\infty}(M)$. Suppose that the flux satisfies the growth condition (5.2). Let T^h be a triangulation and $\tau = \tau(h)$ be a time increment satisfying the conditions (5.3)-(5.4), and let $f_{e,K}$ be a family of numerical flux-functions satisfying the consistency, conservation, and monotonicity conditions (5.5)-(5.8). Let $u^h : \mathbb{R}_+ \times M \to \mathbb{R}$ be the sequence of approximate solutions constructed by the finite volume method (5.9)-(5.11). Then, for all T > 0, the sequence u^h is uniformly bounded in $L^{\infty}([0,T),L^1(M) \cap L^{\infty}(M))$ and converges almost everywhere towards the unique entropy solution $u \in L^{\infty}(\mathbb{R}_+,L^1(M) \cap L^{\infty}(M))$ of the Cauchy problem (1.1)-(5.1).

The proof of this theorem is the subject of the next section.

Remark 5.2. Without loss of generality, we may assume that $\mathrm{Lip}(f)>0$. Indeed, in the trivial case $\mathrm{Lip}(f)=0$, f does not depend on u and the conservation law (1.1) has the explicit solution $u(t,x)=u_0(x)-t\nabla_g\cdot f(x)$. Thus, for any triangulation, if we set $f_{e,K}:=\frac{1}{|e|}\int_e g(f(x),\mathbf{n}(x))\,dv_e$, we immediately find from (5.9) that $u_K^n=u_K^0-n\tau\nabla_g\cdot f(x)$ for all K. This gives $u^h(t,x)=u^h(0,x)-t\nabla_g\cdot f(x)$, and so $\lim_{h\to 0}u^h(t,x)=u_0(x)-t\nabla_g\cdot f(x)$, which is the exact solution.

- **6. Proof of convergence.** This section contains the proof of Theorem 5.1. We follow the strategy of proof proposed by Cockburn, Coquel, and LeFloch in the Euclidian case [2]. As we will see, the derivation of the required estimates for a conservation law on a manifold is significantly more involved.
- 6.1. Discrete maximum principle, entropy dissipation and entropy inequalities. We begin with the discrete maximum principle and the L^1 contraction principle.

Lemma 6.1. The finite volume approximations satisfy

(6.1)
$$\max_{K \in \mathcal{T}^h} |u_K^n| \le \left(\max_{K \in \mathcal{T}^h} |u_K^0| + C_1 t_n \right) e^{C_2 t_n},$$

and, given another approximation sequence v^h (defined in the obvious way from a different initial data $v_0 \in L^1(M) \cap L^{\infty}(M)$),

(6.2)
$$||u^h(t+\tau) - v^h(t+\tau)||_{L^1(M)} \le ||u^h(t) - v^h(t)||_{L^1(M)}.$$

Proof. Let us first point out the following discrete version of Gronwall's inequality: if w_n is a non-negative sequence satisfying $w_0 \leq a$ and

(6.3)
$$w_n \le a + b \sum_{j=1}^{n-1} w_j, \qquad n \ge 1,$$

for some non-negative constants a, b, then it follows that

$$(6.4) w_n \le a e^{bn}.$$

Namely, let us define ϕ_n by $\phi_0 = a$ and

$$\phi_n := a + b \sum_{j=1}^{n-1} \phi_j.$$

We have $\phi_n - \phi_{n-1} = b\phi_{n-1}$, or $\phi_n = (1+b)\phi_{n-1}$. By induction we get $\phi_n = a(1+b)^n$, and the desired estimate now follows from $1+b \le e^b$ and $w_n \le \phi_n$.

Let us turn to the proof of (6.1). Using (5.9) and the consistency condition (5.5), we easily write

$$\begin{split} u_K^{n+1} &= \left(1 + \frac{\tau}{|K|} \sum_{e \in \partial K} \frac{f_{e,K}(u_K^n, u_{K_e}^n) - f_{e,K}(u_K^n, u_K^n)}{u_{K_e}^n - u_K^n} |e| \right) u_K^n \\ &- \frac{\tau}{|K|} \sum_{e \in \partial K} \frac{f_{e,K}(u_K^n, u_{K_e}^n) - f_{e,K}(u_K^n, u_K^n)}{u_{K_e}^n - u_K^n} |e| \, u_{K_e}^n - \frac{\tau}{|K|} \int_K \nabla_g \cdot f(u_K^n, x) \, dv_M. \end{split}$$

The monotonicity condition (5.7) implies

$$-\frac{\tau}{|K|} \sum_{e \in \partial K} \frac{f_{e,K}(u_K^n, u_{K_e}^n) - f_{e,K}(u_K^n, u_K^n)}{u_{K_e}^n - u_K^n} \ge 0,$$

so that, using the CFL condition (5.8),

$$-\frac{\tau|e|}{|K|} \sum_{e \in \partial K} \frac{f_{e,K}(u_K^n, u_{K_e}^n) - f_{e,K}(u_K^n, u_K^n)}{u_{K_e}^n - u_K^n} \le \frac{\tau p_K}{|K|} \operatorname{Lip}(f) \le 1.$$

Thus, we see that u_K^{n+1} is a convex combination of u_K^n and $(u_{K_e}^n)_{e \in \partial K}$, up to a divergence term in (6.5). Therefore, we can write

$$u_K^{n+1} \le \max(u_K^n, \max_{e \in \partial K} u_{K_e}^n) - \frac{\tau}{|K|} \int_K \nabla_g \cdot f(u_K^n, x) \, dv_M,$$

$$u_K^{n+1} \ge \min(u_K^n, \min_{e \in \partial K} u_{K_e}^n) - \frac{\tau}{|K|} \int_K \nabla_g \cdot f(u_K^n, x) \, dv_M,$$

and, by induction,

$$|u_K^n| \le \max_{K \in \mathcal{T}^h} |u_K^0| + \sum_{j=0}^{n-1} \tau \max_{K \in \mathcal{T}^h} \frac{1}{|K|} \int_K |\nabla_g \cdot f(u_K^j, x)| \, dv_M$$

$$\le \max_{K \in \mathcal{T}^h} |u_K^0| + \sum_{j=0}^{n-1} \tau \max_{K \in \mathcal{T}^h} \left(C_1 + C_2 |u_K^j| \right)$$

$$\le \max_{K \in \mathcal{T}^h} |u_K^0| + n\tau C_1 + \sum_{j=0}^{n-1} \tau C_2 \max_{K \in \mathcal{T}^h} |u_K^j|$$

for all n > 0. Using Gronwall's inequality (6.3)-(6.4) we arrive at the desired inequality (6.1).

We now turn to the proof of the L^1 contraction property (6.2). It will follow immediately once we check that

(6.6)
$$\sum_{K \in \mathcal{T}^h} |u_K^{n+1} - v_K^{n+1}||K| \le \sum_{K \in \mathcal{T}^h} |u_K^n - v_K^n||K|.$$

Introduce the discrete solution operator H,

$$u^{n+1} \equiv (u_K^{n+1})_{K \in \mathcal{T}^h} = (H_K(u^n))_{K \in \mathcal{T}^h} =: H(u^n).$$

Using the CFL and monotonicity conditions (5.8) and (5.7), it is easy to check that H is a monotone operator:

$$(u_K^n \le v_K^n, K \in \mathcal{T}^h) \Rightarrow (H_K(u_K^n) \le H_K(v_K^n), K \in \mathcal{T}^h).$$

Setting $a \wedge b = \max(a,b)$ and $a \vee b = \min(a,b)$, we have $H_K(u_K^n \vee v_K^n) = u_K^{n+1} \vee v_K^{n+1}$ and $H_K(u_K^n \wedge v_K^n) = u_K^{n+1} \wedge v_K^{n+1}$. Next, using the definition (5.9) we find

$$\begin{split} u_K^{n+1} \vee v_K^{n+1} - u_K^{n+1} \wedge v_K^{n+1} &= u_K^n \vee v_K^n - u_K^n \wedge v_K^n \\ &- \sum_{e \in \partial K} \frac{\tau|e|}{|K|} \big(f_{e,K}(u_K^n \vee v_K^n, (u_K^n \vee v_K^n)_e) - f_{e,K}(u_K^n \wedge v_K^n, (u_K^n \wedge v_K^n)_e) \big). \end{split}$$

Observe now that

$$(u_K^n \vee v_K^n)_e \leq u_{K_e}^n \vee v_{K_e}^n \text{ and } (u_K^n \wedge v_K^n)_e \geq u_{K_e}^n \wedge v_{K_e}^n.$$

This, together with $\partial f_{e,K}/\partial v \leq 0$, implies that

$$f_{e,K}(u_K^n \vee v_K^n, (u_K^n \vee v_K^n)_e) - f_{e,K}(u_K^n \wedge v_K^n, (u_K^n \wedge v_K^n)_e)$$

$$\geq f_{e,K}(u_K^n \vee v_K^n, u_{K_e}^n \vee v_{K_e}^n) - f_{e,K}(u_K^n \wedge v_K^n, u_{K_e}^n \wedge v_{K_e}^n),$$

so that

$$\begin{split} u_K^{n+1} \vee v_K^{n+1} - u_K^{n+1} \wedge v_K^{n+1} &\leq u_K^n \vee v_K^n - u_K^n \wedge v_K^n \\ &- \sum_{e \in \partial K} \frac{\tau|e|}{|K|} \big(f_{e,K}(u_K^n \vee v_K^n, u_{K_e}^n \vee v_{K_e}^n) - f_{e,K}(u_K^n \wedge v_K^n, u_{K_e}^n \wedge v_{K_e}^n) \big). \end{split}$$

The inequality (6.6) now follows by multiplying by |K| and summing in $K \in \mathcal{T}^h$, since the conservation property (5.6) allows us to cancel all of the numerical flux terms. The proof of Lemma 6.1 is complete. \square

The discrete entropy inequalities will be formulated in terms of the following convex decomposition of u_K^n . For each K, e, define (cf. [2], formula (3.6))

(6.7)
$$\tilde{u}_{K,e}^{n+1} := u_K^n - \frac{\tau p_K}{|K|} \left(f_{e,K}(u_K^n, u_{K_e}^n) - f_{e,K}(u_K^n, u_K^n) \right)$$

and set

(6.8)
$$u_{K,e}^{n+1} := \tilde{u}_{K,e}^{n+1} - \frac{\tau}{|K|} \int_K \nabla_g \cdot f(u_K^n, x) \, dv_M.$$

In view of (5.9) and the consistency property (5.5), we have

(6.9)
$$u_K^{n+1} = \frac{1}{p_K} \sum_{e \in \partial K} |e| u_{K,e}^{n+1}.$$

The following lemma establishes the existence of numerical flux terms and a first entropy inequality relating $\tilde{u}_{K,e}^{n+1}$ and u_K^n . We omit the proof and refer to [2] and the references therein.

LEMMA 6.2. Let (U, F) = (U(u), F(u, x)) be a convex entropy pair. Then there exists a family of numerical flux functions $F_{e,K} : \mathbb{R}^2 \to \mathbb{R}$ satisfying the following conditions:

• $F_{e,K}$ is consistent with the entropy flux F:

$$F_{e,K}(u,u) = \frac{1}{|e|} \int_{e} g(F(u,x), \mathbf{n}_{e,K}(x)) dv_e, \qquad u \in \mathbb{R}.$$

• Conservation property:

$$F_{e,K}(u,v) = -F_{e,K_e}(v,u), \qquad u,v \in \mathbb{R}.$$

• Discrete entropy inequality:

$$(6.10) U(\tilde{u}_{K,e}^{n+1}) - U(u_K^n) + \frac{\tau p_K}{|K|} \Big(F_{e,K}(u_K^n, u_{K_e}^n) - F_{e,K}(u_K^n, u_K^n) \Big) \le 0.$$

We can now write the discrete entropy inequality in terms of $u_{K,e}^{n+1}$ and u_K^n (see (6.10)):

(6.11)
$$U(u_{K,e}^{n+1}) - U(u_K^n) + \frac{\tau p_K}{|K|} \Big(F_{e,K}(u_K^n, u_{K_e}^n) - F_{e,K}(u_K^n, u_K^n) \Big) \le R_{K,e}^{n+1},$$

$$R_{K,e}^{n+1} := U(u_{K,e}^{n+1}) - U(\tilde{u}_{K,e}^{n+1}).$$

The following entropy dissipation estimate is the key to the proof of convergence, as was observed in the Euclidian case by Cockburn, Coquel, and LeFloch [2].

PROPOSITION 6.3. Let $U : \mathbb{R} \to \mathbb{R}$ be a strictly convex, smooth function. Then, for all n we have

$$(6.12) \sum_{K \in \mathcal{T}^{h}} U(u_{K}^{n+1})|K| + \frac{\alpha}{2} \sum_{\substack{K \in \mathcal{T}^{h} \\ e \in \partial K}} \frac{|e||K|}{p_{K}} |u_{K,e}^{n+1} - u_{K}^{n+1}|^{2}$$

$$\leq \sum_{\substack{K \in \mathcal{T}^{h} \\ e \in \partial K}} U(u_{K}^{n})|K| + \sum_{\substack{K \in \mathcal{T}^{h} \\ e \in \partial K}} \tau|e|F_{e,K}(u_{K}^{n}, u_{K}^{n}) + \sum_{\substack{K \in \mathcal{T}^{h} \\ e \in \partial K}} \frac{|e||K|}{p_{K}} R_{K,e}^{n+1},$$

where α denotes the modulus of convexity of U.

Proof. First, we observe that from an elementary lemma on convex functions (Lemma 3.5 of [2]) and (6.9) one easily obtains

(6.13)
$$\sum_{K \in \mathcal{T}^h} |K| U(u_K^{n+1}) + \frac{\alpha}{2} \sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \frac{|e||K|}{p_K} |u_{K,e}^{n+1} - u_K^{n+1}|^2 \le \sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \frac{|e||K|}{p_K} U(u_{K,e}^{n+1}).$$

Next, we sum up the inequality (6.11) with respect to all K and e, and divide by p_K to obtain

$$\sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \frac{|e||K|}{p_K} U(u_{K,e}^{n+1}) - \sum_{K \in \mathcal{T}^h} U(u_K^n)|K| - \sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \tau |e| F_{e,K}(u_K^n, u_K^n) - \sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \frac{|e||K|}{p_K} R_{K,e}^{n+1} \le 0.$$

Note that, again, all of the flux terms $F_{e,K}(u_K^n, u_{K_e}^n)$ cancel out due to conservation. Combining (6.14) with (6.13) gives (6.12). The proof of Proposition 6.3 is complete.

As a consequence of the entropy dissipation estimate (6.12), one can deduce the following bound:

(6.15)
$$\sum_{n=1}^{N} \sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \frac{|e||K|}{p_K} |u_{K,e}^n - u_K^n|^2 \le ||u_0||_{L^2(M)}^2 + C(T),$$

where C(T) is a constant depending only on T and N is an integer such that $T - N\tau < \tau$. For instance, it suffices to take $U(u) = u^2/2$ in (6.12).

We now derive a local entropy inequality which does not involve the numerical flux and is a consequence of the inequality (6.11). The proof is completely similar to that of Lemma 3.6 in [2] and we omit it. This inequality will then allow us to prove a global discrete entropy inequality.

LEMMA 6.4. Let (U, F) be a convex entropy pair. Then for all elements K and all faces $e \in \partial K$, we have

$$(6.16) \frac{\frac{|K|}{p_{K}}U(u_{K,e}^{n+1}) - \frac{|K|}{p_{K}}U(u_{K}^{n}) + \frac{|K_{e}|}{p_{K_{e}}}U(u_{K_{e},e}^{n+1}) - \frac{|K_{e}|}{p_{K_{e}}}U(u_{K_{e}}^{n})}{+ \frac{\tau}{|e|} \int_{e} g\left(F(u_{K_{e}}^{n}, x) - F(u_{K}^{n}, x), \mathbf{n}_{e,K}(x)\right) dv_{e} \leq \frac{|K|}{p_{K}} R_{K,e}^{n+1} + \frac{|K_{e}|}{p_{K_{e}}} R_{K_{e},e}^{n+1},$$

where $R_{K,e}^{n+1}$ is defined in (6.11).

PROPOSITION 6.5. Let (U, F) be a convex entropy pair and let $\phi = \phi(t, x) \in C_c([0, T) \times M)$ be a test function. For each element K and each face e of K, set

(6.17)
$$\phi_e^n = \frac{1}{\tau |e|} \int_{t_n}^{t_{n+1}} \int_e \phi(t, x) \, dv_e dt, \qquad \hat{\phi}_K^n = \sum_{e \in \partial K} \frac{|e|}{p_K} \phi_e^n,$$

(6.18)
$$\widehat{\partial_t \phi_K}^n = \frac{1}{\tau} (\widehat{\phi}_K^n - \widehat{\phi}_K^{n-1}).$$

Then, u_K^n satisfies the following inequality:

$$(6.19) \qquad -\sum_{n=0}^{N} \sum_{K \in \mathcal{T}^{h}} \int_{t_{n}}^{t_{n+1}} \int_{K} \left(U(u_{K}^{n}) \widehat{\partial_{t} \phi_{K}}^{n} + g\left(F(u_{K}^{n}, x), \nabla_{g} \phi(t, x)\right) \right) dv_{M} dt$$

$$+ \sum_{K \in \mathcal{T}^{h}} \int_{K} U(u_{K}^{0}) \widehat{\phi}_{K}^{0} dv_{M} \leq E^{h} + Q^{h} + R^{h},$$

where the error terms E^h , Q^h and R^h are given by

(6.20)
$$E^{h} = \sum_{n=0}^{N} \sum_{\substack{K \in \mathcal{T}^{h} \\ e \in \partial K}} \frac{|e||K|}{p_{K}} U(u_{K,e}^{n+1})(\hat{\phi}_{K}^{n} - \phi_{e}^{n}),$$

$$Q^{h} := \sum_{n=0}^{N} \sum_{K \in \mathcal{T}^{h}} Q_{K}^{n},$$

$$Q_{K}^{n} := \sum_{e \in \partial K} \tau \int_{e} g(F(u_{K}^{n}, x), \mathbf{n}_{e,K}(x)) dv_{e} \phi_{e}^{n}$$

$$- \int_{t_{n}}^{t_{n+1}} \int_{K} g(F(u_{K}^{n}, x), \nabla_{g} \phi(t, x)) dv_{K} dt,$$

(6.22)
$$R^{h} = \sum_{n=0}^{N} \sum_{\substack{K \in \mathcal{T}^{h} \\ e \in \partial K}} \frac{|K||e|}{p_{K}} \phi_{e}^{n} R_{K,e}^{n+1}$$
$$R_{K,e}^{n+1} = U(u_{K,e}^{n+1}) - U(\tilde{u}_{K,e}^{n+1}).$$

Proof. Multiply the inequality (6.16) by $\phi_e^n|e|$ and sum over all $K \in \mathcal{T}^h$, $e \in \partial K$. Note that the first four terms in (6.16), as well as the last, give the same result whether they involve K or K_e (since the sum is taken over the same set of elements). We thus obtain

$$2\sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \frac{|K||e|}{p_K} \left(U(u_{K,e}^{n+1}) - U(u_K^n) \right) \phi_e^n$$

$$+ \sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \tau |e| \phi_e^n \left(\frac{1}{|e|} \int_e g\left(F(u_{K_e}^n, x) - F(u_K^n, x), \mathbf{n}_{e,K}(x) \right) dv_e \right) \le 2 \sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \frac{|K||e|}{p_K} R_{K,e}^{n+1} \phi_e^n.$$

Also, due to the conservation and consistency properties of $F_{e,K}$ we have

$$\sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \frac{1}{|e|} \int_e g\big(F(u_K^n, x), \mathbf{n}_{e,K}(x)\big) dv_e = -\sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \frac{1}{|e|} \int_e g\big(F(u_{K_e}^n, x), \mathbf{n}_{e,K}(x)\big) dv_e,$$

which implies

(6.23)
$$\sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \frac{|K||e|}{p_K} \left(U(u_{K,e}^{n+1}) - U(u_K^n) \right) \phi_e^n$$

$$- \sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \tau |e| \phi_e^n \frac{1}{|e|} \int_e g\left(F(u_K^n, x), \mathbf{n}_{e,K}(x) \right) dv_e \le \sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \frac{|K||e|}{p_K} R_{K,e}^{n+1} \phi_e^n.$$

Observe now that, setting Q_K^n as in (6.21), we deduce from the definition of ϕ_e^n (see (6.17)) that

$$(6.24) \sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \tau |e| \phi_e^n \frac{1}{|e|} \int_e g(F(u_K^n, x), \mathbf{n}_{e,K}(x)) dv_e$$

$$= \sum_{K \in \mathcal{T}^h} \int_{t_n}^{t_{n+1}} \int_K g(F(u_K^n, x), \nabla \phi(t, x)) dv_M dt + \sum_{K \in \mathcal{T}^h} Q_K^n.$$

From the convexity of U and Jensen's inequality, we see that

(6.25)
$$\sum_{K \in \mathcal{T}^h} |K| U(u_K^{n+1}) \hat{\phi}_K^n \le \sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \frac{|K||e|}{p_K} U(u_{K,e}^{n+1}) \hat{\phi}_K^n.$$

Thus, from (6.23), (6.24) and (6.25) and after summation in n, we obtain

$$\begin{split} & \sum_{n=0}^{N} \sum_{K \in \mathcal{T}^{h}} |K| \left(U(u_{K}^{n+1}) - U(u_{K}^{n}) \right) \hat{\phi}_{K}^{n} - \sum_{n=0}^{N} \sum_{K \in \mathcal{T}^{h}} \int_{t_{n}}^{t_{n+1}} \int_{K} g\left(F(u_{K}^{n}, x), \nabla \phi(t, x) \right) dv_{M} dt \\ & \leq E^{h} + \sum_{n=0}^{N} \sum_{K \in \mathcal{T}^{h}} Q_{K}^{n} + \sum_{n=0}^{N} \sum_{K \in \mathcal{T}^{h}} \frac{|K||e|}{p_{K}} \phi_{e}^{n} R_{K, e}^{n+1}, \end{split}$$

with E^h given by (6.20). Summing by parts in the first term, one obtains the desired entropy inequality (6.19). The proof of Proposition 6.5 is complete. \square

6.2. Measure-valued solutions and convergence analysis. We now prove the strong convergence of the approximate solutions u^h given by the finite volume method towards the entropy solution to the Cauchy problem under consideration. To this end, we need to rely on the framework of measure-valued solutions to conservation laws developed in [4] in the Euclidian case and extended to manifolds by Ben-Artzi and LeFloch [1]. The key question is to justify the passage to the limit as $h \to 0$ in the inequality (6.19) and use the entropy dissipation estimate (6.15) to control the error terms.

It is well-known that to the sequence u^h (which is uniformly bounded in $L^{\infty}([0,T)\times M)$ for every T>0) we can associate a subsequence and a Young measure $\nu:[0,T)\times M\to \operatorname{Prob}(\mathbb{R})$, which is a family of probability measures in \mathbb{R} parametrized by $(t,x)\in[0,T)\times M$. The Young measure allows us to determine all weak-* limits of composite functions $a(u^h)$, for arbitrary real continuous functions a, according to the following property:

(6.26)
$$a(u^h) \stackrel{*}{\rightharpoonup} \langle \nu, a \rangle$$
 when $h \to 0$

where we use the notation

$$\langle \nu, a \rangle := \int_{\mathbb{R}} a(\lambda) \, d\nu(\lambda).$$

In view of the above property, the passage to the limit in the left hand side of (6.19) is (almost) immediate. The key uniqueness theorem [4, 1] tells us that once we know that ν is a measure-valued solution to the conservation law, we can prove that the support of each probability measure $\nu_{t,x}$ actually reduces to a single point u(t,x), if the same is true for t=0, that is, $\nu_{t,x}$ is the Dirac measure $\delta_{u(t,x)}$. It is then standard to deduce that the convergence in (6.26) is actually strong, and that, in particular, u^h converges strongly to u which in turn is the unique entropy solution of the Cauchy problem under consideration.

LEMMA 6.6. Let $\nu:[0,T)\times M\to \operatorname{Prob}(\mathbb{R})$ be a Young measure associated with the sequence u^h . Then, for every convex entropy / entropy-flux pair (U,F) and for every non-negative, smooth function $\phi:[0,T)\times M\to \mathbb{R}_+$ with compact support, one

has (6.27) $-\int_{0}^{T} \int_{M} \langle \nu_{t,x}, U(\cdot) \rangle \partial_{t} \phi(t,x) + g(\langle \nu_{t,x}, F(\cdot, x) \rangle, \nabla \phi(t,x)) dv_{M} dt + \int_{M} U(u_{0}(x)) \phi(0,x) dv_{M} dt$ $\leq \int_{0}^{T} \int_{M} \langle \nu_{t,x}, \nabla_{g} \cdot F(\cdot, x) \rangle \phi(t,x) dv_{M} dt - \int_{0}^{T} \int_{M} \langle \nu_{t,x}, U'(\cdot) \nabla_{g} \cdot f(\cdot, x) \rangle \phi(t,x) dv_{M} dt.$

Proof. First, we claim that the discretizations (6.17), (6.18) satisfy the following estimates, whose proofs are elementary:

$$\sup_{\substack{t \in [t_n, t_{n+1}] \\ x \in K}} \left| \partial_t \phi(t, x) - \widehat{\partial_t \phi_k}^n \right| \le (\tau + h_K) \|\phi\|_{C^2([t_n, t_{n+1}] \times K)},$$

and, for each face e,

$$|\phi_e^n - \hat{\phi}_K^n| \le (\tau + h_K) \|\phi\|_{C^1([t_n, t_{n+1}] \times K)}.$$

In view of these estimates and (6.26), the left hand side of (6.19) converges to the left-hand side of (6.27).

We now consider the error terms on the right-hand side of (6.19). First, note that E^h equals the corresponding one in [2], so we refer the reader to that paper for a proof that E^h converges to zero. Note that the use of estimate (6.15) is essential in that proof.

Next, let us prove that

(6.28)
$$\lim_{h \to 0} R^h = -\int_0^T \int_M \langle \nu_{t,x}, U'(\cdot) \nabla_g \cdot f(\cdot, x) \rangle \phi(t, x) \, d\nu_M dt.$$

Recall that \mathbb{R}^h is given by (6.22). Introducing the averages

$$\phi_K^n := \frac{1}{|K|\tau} \int_t^{t_{n+1}} \int_K \phi(t, x) \, dv_M dt,$$

we claim that it suffices to prove that

(6.29)
$$\lim_{h \to 0} \left(R^h + \sum_n \sum_{K \in \mathcal{T}^h} U'(u_K^n) \tau \int_K \nabla_g \cdot f(u_K^n, x) \, dv_M \, \phi_K^n \right) = 0.$$

Indeed, we have

$$\begin{split} &-\sum_{n}\sum_{K\in\mathcal{T}^{h}}U'(u_{K}^{n})\tau\int_{K}\nabla_{g}\cdot f(u_{K}^{n},x)\,dv_{M}\,\phi_{K}^{n}\\ &=-\sum_{n}\sum_{K\in\mathcal{T}^{h}}U'(u_{K}^{n})\int_{t_{n}}^{t_{n+1}}\int_{K}\nabla_{g}\cdot f(u_{K}^{n},x)\phi(t,x)\,dv_{M}dt\\ &-\sum_{n}\sum_{K\in\mathcal{T}^{h}}U'(u_{K}^{n})\Big(\int_{t_{n}}^{t_{n+1}}\int_{K}\nabla_{g}\cdot f(u_{K}^{n},x)\,dv_{M}\frac{1}{|K|}\int_{K}\phi(t,x)\,dv_{M}dt\\ &-\int_{t_{n}}^{t_{n+1}}\int_{K}\nabla_{g}\cdot f(u_{K}^{n},x)\phi(t,x)\,dv_{M}dt\Big), \end{split}$$

so that taking the limit as $h \to 0$ we find, for the first term,

$$\lim_{h \to 0} - \int_0^T \int_M U'(u^h) \nabla_g \cdot f(u^h, x) \phi(t, x) \, dv_M dt$$
$$= - \int_0^T \int_M \langle \nu_{t, x}, U'(\cdot) \nabla_g \cdot f(\cdot, x) \rangle \phi(t, x) \, dv_M dt,$$

while the second term easily converges to zero.

We now prove (6.29). First, observe that from the definition of ϕ_K^n , it is easy to see that we may as well consider the term R^h with ϕ_K^n instead of ϕ_e^n . We thus have

$$\begin{split} R^h + \sum_{n=0}^{N} \sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \frac{|K||e|}{p_K} U'(u_K^n) \frac{\tau}{|K|} \int_K \nabla_g \cdot f(u_K^n, x) \, dv_M \, \phi_K^n \\ &= R^h + \sum_{n=0}^{N} \sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \frac{|K||e|}{p_K} U'(u_K^{n+1}) \frac{\tau}{|K|} \int_K \nabla_g \cdot f(u_K^{n+1}, x) \, dv_M \, \phi_K^{n+1} \\ &+ \sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \frac{|K||e|}{p_K} U'(u_K^0) \frac{\tau}{|K|} \int_K \nabla_g \cdot f(u_K^0, x) \, dv_M \, \phi_K^0 \\ &= \sum_{n=0}^{N} \sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \frac{|K||e|}{p_K} \phi_K^n \Big(U(u_{K,e}^{n+1}) - U(\tilde{u}_{K,e}^{n+1}) + U'(u_K^{n+1}) \frac{\tau}{|K|} \int_K \nabla_g \cdot f(u_K^{n+1}, x) \, dv_M \Big) \\ &+ \sum_{n=0}^{N} \sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \frac{|K||e|}{p_K} U'(u_K^{n+1}) \frac{\tau}{|K|} \int_K \nabla_g \cdot f(u_K^{n+1}, x) \, dv_M (\phi_K^{n+1} - \phi_K^n) \\ &- \sum_{K \in \mathcal{T}^h} U'(u_K^0) \tau \int_K \nabla_g \cdot f(u_K^0, x) \, dv_M \phi_K^0 = A + B + C. \end{split}$$

Note that all sums in n may be written from 0 to N because of the compact support of ϕ .

Consider the first term, A. From the bound (5.2) on the divergence of f and from the maximum principle (6.1) we conclude that

$$\frac{\tau}{|K|} \int_K \nabla_g \cdot f(u_K^{n+1}, x) \, dv_M = \mathcal{O}(\tau).$$

Thus, Taylor developing U around the point $\tilde{u}_{K,e}^{n+1}$, and using (6.8), the definition of $u_{K,e}^{n+1}$, we find

$$|A| \leq \sum_{n} \sum_{\substack{K \in \mathcal{T}^{h} \\ e \in \partial K}} \frac{|K||e|}{p_{K}} \left| U'(u_{K,e}^{n+1})\mathcal{O}(\tau) + \mathcal{O}(\tau^{2}) - U'(u_{K}^{n+1})\mathcal{O}(\tau) \right| \phi_{K}^{n}$$

$$\leq \mathcal{O}(\tau) \sup U'' \sum_{n} \sum_{\substack{K \in \mathcal{T}^{h} \\ e \in \partial K}} \frac{|K||e|}{p_{K}} |u_{K,e}^{n+1} - u_{K}^{n+1}| \phi_{K}^{n}.$$

Next, using the estimate (6.15) and the Cauchy-Schwartz inequality, we find that the term inside the sum above is bounded, so that $A = \mathcal{O}(\tau)$.

Consider now the second term, B. The function ϕ is regular, so that $\phi_K^{n+1} - \phi_K^n =$ $\mathcal{O}(\tau)$. Therefore, we have

$$|B| \le \sum_{n} \sum_{K \in \mathcal{T}^h} |K| \mathcal{O}(\tau^2) = \mathcal{O}(\tau).$$

Finally, using again the bound (5.2), we see that the last term C tends to zero with

au. This concludes the proof of (6.28). Next, consider the term $Q^h = \sum_{n,K} Q_K^n$ given by (6.21). Using Green's formula, we can write

$$\begin{split} Q_K^n &= \sum_{e \in \partial K} \frac{1}{|e|} \int_e g\big(F(u_K^n, x), \mathbf{n}_{e,K}(x)\big) \, dv_e \int_{t_n}^{t_{n+1}} \int_e \phi(t, x) \, dv_e dt \\ &- \int_{t_n}^{t_{n+1}} \int_{\partial K} g\big(F(u_K^n, x) \phi(t, x), \mathbf{n}(x)\big) \, dv_{\partial K} dt + \int_{t_n}^{t_{n+1}} \int_K \nabla_g \cdot F(u_K^n, x) \phi(t, x) \, dv_M dt. \end{split}$$

We wish to obtain

(6.30)
$$\lim_{h \to 0} Q^h = \int_0^T \int_M \langle \nu_{t,x}, \nabla_g \cdot F(\cdot, x) \rangle \phi(t, x) \, d\nu_M dt,$$

so that according to the above equality we must prove that

(6.31)
$$\lim_{h \to 0} \sum_{n} \sum_{\substack{K \in \mathcal{T}^h \\ e \in \partial K}} \left(\frac{1}{|e|} \int_{e} g(F(u_K^n, x), \mathbf{n}_{e,K}(x)) dv_e \int_{t_n}^{t_{n+1}} \int_{e} \phi(t, x) dv_e dt - \int_{t_n}^{t_{n+1}} \int_{e} g(F(u_K^n, x)\phi(t, x), \mathbf{n}_{e,K}(x)) dv_{\partial K} dt \right) = 0.$$

Let x_e be any point on the edge e. Note that

$$\sum_{n} \sum_{\substack{K \in \mathcal{T}^{h} \\ e \in \partial K}} \int_{t_{n}}^{t_{n+1}} \int_{e} g\left(F(u_{K}^{n}, x), \mathbf{n}_{e,K}(x)\right) \left(\frac{1}{|e|} \int_{e} \phi(t, x) \, dv_{e} - \phi(t, x)\right) dv_{e} dt$$

$$= \sum_{n} \sum_{\substack{K \in \mathcal{T}^{h} \\ e \in \partial K}} \int_{t_{n}}^{t_{n+1}} \int_{e} \left(g\left(F(u_{K}^{n}, x), \mathbf{n}_{e,K}(x)\right) - g\left(F(u_{K}^{n}, x_{e}), \mathbf{n}_{e,K}(x_{e})\right)\right) \cdot \left(\frac{1}{|e|} \int_{e} \phi(t, x) \, dv_{e} - \phi(t, x)\right) dv_{e} dt,$$

so that taking absolute values and in view of the regularity of F and ϕ , we can bound the last term by

$$CT\sum_{K\in\mathcal{T}^h}p_Kh^2.$$

using the CFL condition (5.8), we can finally write

$$\begin{aligned} \left| Q^h - \int_0^T \int_M \nabla_g \cdot F(u^h, x) \phi(t, x) \, dv_M dt \right| &\leq \frac{CT}{\operatorname{Lip}(f)} \sum_{K \in \mathcal{T}^h} p_K h^2 \\ &\leq C \sum_{K \in \mathcal{T}^h} \frac{|K|}{\tau} h^2 = \frac{h^2}{\tau} C|M|, \end{aligned}$$

so that taking the limit as $h \to 0$ and using (5.4) yields

$$Q^h \to \int_0^T \int_M \langle \nu_{t,x}, \nabla_g \cdot F(\cdot, x) \rangle \phi(t, x) \, dv_M dt,$$

which is (6.30). This completes the proof of Lemma 6.6. \square

Remark 6.7. From the proof of the previous lemma, we see that if the divergencefree condition (1.2) is satisfied, then the error term Q^h converges to zero with h and R^h vanishes altogether. In that case, we obtain, instead of (6.27), the usual weak formulation

$$-\int_{0}^{T} \int_{M} \langle \nu_{t,x}, U(\cdot) \rangle \partial_{t} \phi(t,x) + g(\langle \nu_{t,x}, F(\cdot, x) \rangle, \nabla \phi(t,x)) dv_{M} dt$$
$$+ \int_{M} U(u_{0}(x)) \phi(0,x) dv_{M} \leq 0.$$

Proof of Theorem 5.1. According to the inequality (6.27), we have for all convex entropy pairs (U, F),

$$\partial_t \langle \nu, U(\cdot) \rangle + \nabla_q \cdot (\langle \nu, F(\cdot, x) \rangle) \le \langle \nu, (\nabla_q \cdot F)(\cdot, x) \rangle - \langle \nu, U'(\cdot)(\nabla_q \cdot f)(\cdot, x) \rangle$$

in the sense of distributions in $[0,T) \times M$. Since, for t=0, the Young measure ν is the Dirac mass δ_{u_0} (because u_0 is a bounded function), we know from [1] that there exists a unique function $u \in L^{\infty}([0,T) \times M)$ such that the measure ν remains the Dirac mass δ_u for all time $0 \le t \le T$. Moreover, this implies that the approximations u^h converge strongly to u on compact sets at least. This concludes the proof. \square

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