

## Heat Equation on Phase Space and the Classical Limit of Quantum Mechanical Expectation Values

A. Grossmann

Centre de Physique Théorique, C.N.R.S., F-13274 Marseille Cedex 2, France

R. Seiler

Freie Universität Berlin, FB 20, Physik, D-1000 Berlin 33

**Abstract.** The expectation value of a quantum mechanical operator, taken in coherent states and suitably rescaled, is the solution of an initial value problem for the heat equation on phase space, in which  $\hbar$  plays the role of time, and the classical observable is the distribution of temperature at  $\hbar=0$ .

### Introduction

A recent paper by Hepp [1] is devoted to the classical limit of (rescaled) expectation values in coherent states and to their time evolution. Here we sharpen some results of [1] by relating the classical limit to an initial value problem in  $\hbar$ . This is done with the help of a quantization formula derived in [2].

### Notations

Denote by  $E$  a  $2\nu$ -dimensional real vector space with a symplectic form  $\sigma$ . (Phase space for  $\nu < \infty$  degrees of freedom.) Elements of  $E$  will be denoted by  $a, b, v, \dots$ . Fix on  $E$  a  $\sigma$ -allowed complex structure  $J$ , i.e. a linear map satisfying  $J^2 = -1$ ,  $\sigma(Ja, Jv) = \sigma(a, v)$  and  $\sigma(a, Ja) > 0$  for  $a \neq 0$ . Introduce the orthogonal form  $s(a, v) = \sigma(a, Jv)$ , and the (phase space) Gaussian  $\Omega(v) = e^{-\pi s(v, v)}$ . Normalize the invariant measure  $dv$  on  $E$  by the requirement  $\int \Omega(v) dv = 1$ . This is equivalent to the requirement  $F^2 = 1$  where  $F$  is the symplectic Fourier transform:

$$Ff(v) = \tilde{f}(v) = \int e^{2i\pi\sigma(v, v')} f(v') dv'.$$

In the Hilbert space  $L^2(E; dv)$  consider the family of functions  $\Omega^a$ :

$$\Omega^a(v) = e^{-2i\pi\sigma(a, v)} \Omega(v + a).$$

Denote by  $\mathcal{H}$  the closed linear span of the family  $\Omega^a$ , with the scalar product inherited from  $L^2(E; dv)$ . For any  $\Phi \in \mathcal{H}$  one has  $(\Omega^a, \Phi) = k\Phi(-a)$ , with

$$k = (\Omega, \Omega) = 2^{-\nu}.$$

Also, for every  $\Phi \in \mathcal{H}$  one has  $F\Phi = M\Phi$ , where  $M$  is the parity operator:

$$(M\Phi)(v) = \Phi(-v).$$

Define  $(W(a)\Phi)(v) = e^{-2i\pi\sigma(a,v)}\Phi(v+a)$ . The Weyl operators  $W(a)$  act irreducibly in  $\mathcal{H}$ .

A convenient way of writing (even very unbounded) linear operators  $A$  in  $\mathcal{H}$  is to consider the associated kernel:  $A(a, b) = (\Omega^a, A\Omega^b)$ . One proves then

$$(A\Phi)(a) = (1/k^2) \int A(-a, -b)\Phi(b)db.$$

### Weyl Quantization

It consists in associating, to a function  $f_c$  on phase space, the operator  $Q(f_c)$  defined formally by

$$Q(f_c) = \int \tilde{f}_c(v)W(-v/2)dv = \int f_c(v/2)W(v)Mdv. \tag{1}$$

It has been shown in [2] that the two expressions coincide. In order to avoid a discussion of the convergence of the operator-valued integrals, we replace (1) by the kernel

$$(\Omega^a, Q(f_c)\Omega^b) = \int \tilde{f}_c(v)\Omega^{(a,b)}(-v/2)dv = \int f_c(v/2)\Omega^{(a,-b)}(v)dv \tag{2}$$

where

$$\Omega^{(a,b)}(v) = (\Omega^a, W(v)\Omega^b) = ke^{2i\pi\sigma(b,a)}e^{-2i\pi\sigma(a+b,v)}\Omega(v-a+b). \tag{3}$$

### Heat Equation on Phase Space

Define the Laplacian,  $\Delta$ , on  $E$ , by  $\Delta = -FsF$ , where  $s$  is the operator of multiplication by  $s(v, v)$ . Consider on  $E$  the heat equation:

$$\partial f / \partial \hbar = (\pi/4)\Delta f. \tag{4}$$

Let  $f_c$  be a function in the uniqueness and correctness class for (4); this is only a very mild requirement. Define  $f(\hbar, v)$  as the solution of (4), with initial data  $f_c(v)$ . In other words,  $f(\hbar, v)$  is the distribution of “temperature” at “time”  $\hbar$ , resulting from an initial distribution  $f(0, v) = f_c(v)$ .

**Theorem.** *One has*

$$f(\hbar, v) = (1/k) (\Omega^{\hbar^{-\frac{1}{2}}v}, Q(f_c(\hbar^{\frac{1}{2}}\cdot))\Omega^{\hbar^{-\frac{1}{2}}v}) \tag{5}$$

where  $f_c(\hbar^{\frac{1}{2}}\cdot)$  is the function  $v \rightarrow f_c(\hbar^{\frac{1}{2}}v)$ .  $\square$

Equation (5) describes very intuitively the way in which a (suitably rescaled) matrix element tends to a classical function. We shall apply it in a forthcoming paper to the study of time evolution.

In order to prove (5), specialize (2) to

$$(\Omega^a, Q(f)\Omega^a) = k \int f(v/2)\Omega(v-2a)dv$$

and notice that  $G_\lambda(v) = \lambda^{-\nu} \Omega(\lambda^{-\frac{1}{2}} v)$  is the elementary solution of the heat equation  $\partial G / \partial \lambda = \pi \Delta G$ .

It is possible to derive analogous equations for off-diagonal matrix elements, and equations in which the initial data are given by  $\tilde{f}_c$ .

*Acknowledgement.* This work was partially supported by contract SE 2871 of the Deutsche Forschungsgemeinschaft.

## References

1. Hepp, K.: Commun. math. Phys. **35**, 265 (1974)
2. Grossmann, A.: Commun. math. Phys. **48**, 191—194 (1976)

Communicated by H. Araki

Received November 10, 1975

