

## A NOTE ON PLANAR BROWNIAN MOTION

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The joint density of the total winding and the radius of a planar Brownian motion is calculated, by solving the associated backward equation. An explicit expression for the distribution of the hitting time of an angular barrier is shown; in particular, the behavior of the tail of the distribution is determined.

**1. Introduction.** Consider a planar Brownian motion  $Z_t = X_t + iY_t$  starting at some point  $z_0 \neq 0$ . Let  $R_t = |Z_t|$  and  $\theta_t$  be the total winding of the a.s. continuous path  $\{Z_s; s \leq t\}$  about zero. The continuous process  $(R_t, \theta_t)$  can be thought of as a diffusion on the Riemann surface of the logarithm (since  $Z$  does not hit 0 a.s.), identified with the half upper plane divided into vertical strips of width  $2\pi$ . Each one of these strips is sent into  $\mathbb{R}^2 \setminus 0$  via the map  $(r, \theta) \rightarrow (r \cos \theta, r \sin \theta)$ , which introduces local coordinates. As discussed in Theorem 1,  $(R_t, \theta_t)$  is governed by  $\frac{1}{2}\Delta$ , which yields (in the given coordinates), that the density

$$(1.1) \quad p(t, r, \theta; \rho, \alpha) = \mathbf{P}\{R_t \in dr, \theta_t \in d\theta | R_0 = \rho, \theta_0 = \alpha\} 1/(r dr d\theta)$$

satisfies the equation

$$(1.2) \quad \partial_t u = \frac{1}{2} \left( \partial_r^2 u + \frac{1}{r} \partial_r u + \frac{1}{r^2} \partial_\theta^2 u \right), \quad t, r > 0, \theta \in (-\infty, \infty).$$

As a first result (Theorem 1) we obtain  $p$  by finding a fundamental solution of (1.2) by standard methods [with the aid of Gradshteyn and Ryzhik's tables (1980)]. Once this density is known, an application of the reflection principle to  $\theta_t$  permits computation of the joint distribution of  $(R_t, \theta_t)$  and  $T_\beta$  [the exit time of  $\theta_t$  from the region  $(0, \beta)$ ]. This is done in Theorem 2. The asymptotic behavior of  $\mathbf{P}\{T_\beta > t\}$  as  $t$  goes to infinity is determined as a corollary; in particular, the result of Spitzer (1958) asserting that  $\mathbf{E}T_\beta$  is finite if and only if  $\beta < \pi/2$  follows directly from it.

While we were writing this note, we learned of an article of Yor (1980), where he shows, using martingale methods, that  $\mathbf{E}\{\exp(iv(\theta_t - \theta_0)) | R_t = r, R_0 = \rho\} = (I_{|v|}(\rho r/t))/(I_0(\rho r/t))$ . [In fact, this formula has a longer history; see Yor (1980) and Pitman and Yor (1986) for details and further references]. Since the distribution of  $R_t$  is known, the joint density  $p$  can be deduced. We think, however, that it has some interest to consider the extended equation, since it is very simple and may be useful in other cases.

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**2. Results and proofs.** Our results will be expressed in terms of the modified Bessel functions  $I_\nu$ . Several representations of them are known. See Gradshteyn and Ryzhik (1980) for a brief account, or Watson (1966) for details. We write down some of them, just for easy reference.

*Representations of the Bessel function  $I_\nu$ .* We shall only consider  $I_\nu(x)$  for  $x, \nu \in \mathbb{R}, x > 0$ . Then,

$$(2.1) \quad I_\nu(x) = \frac{1}{2\pi i} \int_{\infty - \pi i}^{\infty + \pi i} \exp(x \cosh \omega - \nu \omega) d\omega.$$

Taking as the path of integration three sides of the rectangle with corners at  $\infty - \pi i, -\pi i, \pi i$  and  $\infty + \pi i$ , the previous expression yields

$$(2.2) \quad I_\nu(x) = \frac{1}{\pi} \int_0^\pi e^{x \cos \theta} \cos \nu \theta d\theta - \frac{\sin \nu \pi}{\pi} \int_0^\infty e^{-x \cosh t - \nu t} dt.$$

Finally, the series expansion is

$$(2.3) \quad I_\nu(x) = \sum_{k=0}^\infty \frac{(x/2)^{\nu+2k}}{k! \Gamma(\nu+k+1)}.$$

Now, we can state:

**THEOREM 1.** *The fundamental density defined in (1.1) is given by*

$$(2.4) \quad p(t, r, \theta; \rho, \alpha) = \frac{1}{\pi t} \exp\left(-\frac{r^2 + \rho^2}{2t}\right) \int_0^\infty \cos \nu(\theta - \alpha) I_\nu\left(\frac{\rho r}{t}\right) d\nu.$$

**PROOF.** The backward equation for  $p(t, r, \theta; \rho, \alpha)$  is

$$(2.5) \quad \partial_t u = \frac{1}{2} \left( \partial_\rho^2 u + \frac{1}{\rho} \partial_\rho u + \frac{1}{\rho^2} \partial_\alpha^2 u \right).$$

Indeed, the proof that the backward equation for the density of  $Z_t$  is the heat equation is based on a local argument about the initial point [see, for instance, Feller (1971)] and yields, in polar coordinates, (2.5). But the argument still applies if one considers the total angle instead of identifying points with equal angle modulo  $2\pi$ . The periodic boundary conditions that arise in this last case disappear and we are left with equation (2.5), for  $\rho > 0$ , and  $\alpha \in (-\infty, +\infty)$ .

By symmetry,  $p$  is an even function of the difference  $\omega = \theta - \alpha$  (that will be denoted again by  $p$ ). We shall consider (2.5) in this variable and  $\alpha \in [0, 2\pi)$ , so it corresponds with the usual polar coordinate of the initial point of  $Z_t$ . The corresponding forward equation is (1.2), so  $p$  also satisfies

$$(2.6) \quad \partial_t u = \frac{1}{2} \left( \partial_r^2 u + \frac{1}{r} \partial_r u + \frac{1}{r^2} \partial_\omega^2 u \right).$$

Let us call  $\mathcal{M} = (0, \infty) \times (-\infty, \infty)$ . Recall that the volume element that corresponds to  $\mathcal{M}$  in our case is  $r dr d\omega$ . We shall find a fundamental solution of

this last equation, that is, a function  $u(t, r, \omega; \rho) \in C^\infty[(0, \infty) \times \mathcal{M}]$  satisfying it and such that for any given open set  $\mathbf{U} \subset \mathcal{M}$ ,

$$(2.7) \quad \lim_{t \downarrow 0} \int_{\mathbf{U}} u(t, r, \omega; \rho) r dr d\omega = \mathbb{1}_{\mathbf{U}}(\rho, 0),$$

where  $\mathbb{1}_{\mathbf{U}}$  denotes the indicator function of the set  $\mathbf{U}$ . Now, denote  $\hat{u}(t, r, v; \rho) = \int_{-\infty}^{\infty} e^{iv\omega} u(t, r, \omega; \rho) d\omega$ . From (2.5) and (2.6),

$$(2.8) \quad \partial \hat{u}_t = \frac{1}{2} \left( \partial_r^2 \hat{u} + \frac{1}{r} \partial_r \hat{u} - \frac{v^2}{r^2} \hat{u} \right) = \frac{1}{2} \left( \partial_\rho^2 \hat{u} + \frac{1}{\rho} \partial_\rho \hat{u} - \frac{v^2}{\rho^2} \hat{u} \right).$$

Performing the separation of variables  $\hat{u} = T(t)F(r, \rho, v)$ , one gets

$$(2.9a) \quad 2T' = \lambda_0 T,$$

$$(2.9b) \quad \partial_r^2 F + \frac{1}{r} \partial_r F - \frac{v^2}{r^2} F = \lambda_0 F,$$

$$(2.9c) \quad \partial_\rho^2 F + \frac{1}{\rho} \partial_\rho F - \frac{v^2}{\rho^2} F = \lambda_0 F,$$

where  $\lambda_0$  is the separation constant. Observe that  $\int_0^{2\pi} \varphi(t, r, \omega; \rho) d\omega = \int_0^{2\pi} \varphi(t, r, \omega; \rho) d\omega$ , where

$$(2.10) \quad \varphi(t, r, \omega; \rho) = \frac{1}{2\pi t} \exp\left(-\frac{r^2 + \rho^2}{2t} + \frac{\rho r}{t} \cos \omega\right),$$

that is, the density  $P\{R_t \in dr, \tilde{\theta}_t \in d\omega | R_0 = \rho, \theta_0 = 0\} 1/(r dr d\omega) \tilde{\theta}_t$  being now the usual polar angle of  $Z_t$ . Then,  $\hat{p}$  remains bounded as  $t \rightarrow \infty$  (for fixed positive  $r, \rho$ ) which justifies taking  $\lambda_0 = -\lambda \leq 0$ . In this case, (2.9b) and (2.9c) are Bessel equations with solution of the form  $A J_v(\lambda^{1/2} r) J_v(\lambda^{1/2} \rho)$ , where  $A$  may depend on  $v$  and  $\lambda$ . (Recall that  $p$  should be regular as  $r, \rho \rightarrow 0$ .) Putting all this together, one obtains that

$$(2.11) \quad \int_0^\infty \exp\left(\frac{-\lambda t}{2}\right) A(\lambda, v) J_v(\lambda^{1/2} r) J_v(\lambda^{1/2} \rho) d\lambda$$

is a solution of (2.8), where  $A$  has to be determined from (2.7). But it is known [see Gradshteyn and Ryzhik (1980), formula 6.615] that

$$(2.12) \quad \int_0^\infty \exp\left(\frac{-\lambda t}{2}\right) J_v(\lambda^{1/2} r) J_v(\lambda^{1/2} \rho) d\lambda = \frac{2}{t} I_v\left(\frac{\rho r}{t}\right) \exp\left(-\frac{\rho^2 + r^2}{2t}\right).$$

Also, from the asymptotic expansion of  $I_v$  [see Watson (1966)], one knows that  $(2\pi\rho r/t)^{1/2} I_v(\rho r/t) e^{-\rho r/t} \rightarrow_{t \rightarrow 0} 1$ , which yields that  $(r/t) I_v(\rho r/t) \times \exp -(\rho^2 + r^2)/2t \rightarrow_{t \rightarrow 0} \delta(r - \rho)$ . Taking  $A(\lambda, v) = \frac{1}{2}$ , one finds (after inverting the Fourier transform) that the right-hand side (r.h.s.) of (2.4) is a formal solution of the forward equation (1.2) satisfying (2.7). It is easily checked that it is, in fact, a solution. Indeed, from (2.1) it follows that the r.h.s. of (2.12) is  $C^\infty$  as a function of  $t$  and  $r$ ; it is clear from the left-hand side

(l.h.s.) that it solves in (2.8). Taking now as the path of integration in (2.1) three sides of the rectangle with corners at  $\infty - \pi i$ ,  $a - \pi i$ ,  $a + \pi i$  and  $\infty + \pi i$ , for some positive  $a$ , it follows at once that  $I_\nu(x) \in L^1(d\nu)$ , for fixed positive  $x$ . Fourier inversion is then justified. Moreover,  $\nu^2 I_\nu(x)$  also belongs to  $L^1(d\nu)$ , which permits us to conclude that we have a true solution. In order to see that it is, in fact, the density  $p$ , it is enough to see that it is positive and that it is the smallest elementary solution of (2.6). [See McKean, (1969).]

Now, the positivity follows from the formula

$$\begin{aligned}
 & \int_0^\infty I_\nu(x) \cos \nu \omega \, d\nu \\
 (2.13) \quad &= \frac{1}{2} e^{r \cos \omega} \mathbb{1}_{(-\pi, \pi)}(\omega) \\
 & \quad - \frac{1}{2\pi} \int_0^\infty e^{-r \cosh t} \left[ \frac{\pi + \omega}{(\pi + \omega)^2 + t^2} + \frac{\pi - \omega}{(\pi - \omega)^2 + t^2} \right] dt,
 \end{aligned}$$

which is obtained after some calculations from (2.2).

Also, it is known [see Gradshteyn and Ryzhik (1980), formula 8.511-5] that

$$\begin{aligned}
 (2.14) \quad & \frac{1}{\pi t} \exp\left(-\frac{r^2 + \rho^2}{2}\right) \sum_{k \in \mathbb{Z}} \int_0^\infty \cos [v(\omega + 2k\pi)] I_\nu\left(\frac{r\rho}{t}\right) d\nu \\
 &= \frac{1}{2\pi t} \exp\left(-\frac{r^2 + \rho^2}{2t} + \left(\frac{\rho r}{t}\right) \cos \omega\right).
 \end{aligned}$$

Observe that the right-hand side is the density (2.10) (of the polar coordinates of  $Z_t$ ), which is the smallest solution of (2.6) with periodic boundary condition at  $[0, 2\pi]$  and pole at  $(\rho, 0)$ . Then the r.h.s. of (2.4) is the smallest elementary solution of (2.6), as desired.  $\square$

REMARKS. From (2.13) we have for  $p$  a formula in terms of elementary functions (recall that  $\omega = \theta - \alpha$ ):

$$\begin{aligned}
 (2.15) \quad p(t, r, \omega; \rho) &= \frac{1}{\pi t} \exp\left(-\frac{r^2 + \rho^2}{2t}\right) \\
 & \times \left\{ \frac{1}{2} e^{r \cos \omega} \mathbb{1}_{(-\pi, \pi)}(\omega) \right. \\
 & \quad \left. - \frac{1}{2\pi} \int_0^\infty e^{-r \cosh t} \left[ \frac{\pi + \omega}{(\pi + \omega)^2 + t^2} + \frac{\pi - \omega}{(\pi - \omega)^2 + t^2} \right] dt \right\}.
 \end{aligned}$$

Also, the characteristic function  $\mathbf{E}(e^{i\theta_t} | \theta_0 = 0, R_0 = \rho)$  can be calculated from (2.4):

$$\begin{aligned}
 (2.16) \quad & \mathbf{E}\{e^{i\theta_t} | \theta_0 = 0, R_0 = \rho\} \\
 &= \rho \frac{\sqrt{2\pi}}{4\sqrt{t}} \exp\left(-\frac{\rho^2}{4t}\right) \left[ I_{(|\nu|-1)/2}\left(\frac{\rho^2}{4t}\right) + I_{(|\nu|+1)/2}\left(\frac{\rho^2}{4t}\right) \right].
 \end{aligned}$$

This formula was deduced by Spitzer (1958) and from it, he concluded at once that  $2\theta_t/\log t$  converges in distribution to a Cauchy random variable as  $t$  goes to  $\infty$ . A simpler proof of Spitzer's asymptotic law can be found in Section 3 of Pitman and Yor (1986). Extensions, related developments and more references on the subject are also contained in that article and in Pitman and Yor (1989).

Consider now  $0 < \alpha < \beta$  and  $T_\beta = \inf\{t > 0: \theta_t \notin (0, \beta)\}$ , where  $\theta_0 = \alpha$ . We want to calculate

$$u(t, r, \theta; \rho, \alpha) = \mathbf{P}\{T_\beta > t, R_t \in dr, \theta_t \in d\theta | R_0 = \rho, \theta_0 = \alpha\} \frac{1}{(r dr d\theta)}.$$

But, from the rotational symmetry of  $Z_t$ ,  $u$  can be obtained by applying successive reflections in both barriers. This gives that

$$(2.17) \quad u(t, r, \theta; \rho, \alpha) = \sum_{k \in \mathbb{Z}} p(t, r, 2k\beta + \theta; \rho, \alpha) - p(t, r, 2k\beta - \theta; \rho, \alpha).$$

From Theorem 1, and after some manipulations of this last expression, one obtains:

**THEOREM 2.**

$$(2.18) \quad u(t, r, \theta; \rho, \alpha) = \frac{1}{\beta t} \exp\left(-\frac{r^2 + \rho^2}{2t}\right) \sum_{k \in \mathbb{Z}} I_{|k\pi/\beta|}\left(\frac{\rho r}{t}\right) \sin\left(\frac{k\pi\alpha}{\beta}\right) \sin\left(\frac{k\pi\theta}{\beta}\right).$$

**PROOF.** The series in (2.18) is clearly convergent. From (2.17), the r.h.s. of (2.18) satisfies (1.2), condition (2.7) and the boundary condition  $u(t, r, 0; \rho, \alpha) = u(t, r, \beta; \rho, \alpha) = 0$  for all  $r, t > 0$ , so it must be the desired density.

**COROLLARY.** Consider the case  $\alpha = \beta/2$ . Then,

$$(2.19) \quad \mathbf{P}\{T_\beta > t\} = 4 \frac{e^{-\rho^2/4t}}{\pi^{1/2} \Gamma(\pi/2\beta + 1/2)} \left(\frac{\rho^2}{8t}\right)^{\pi/2\beta} + o(t^{-\pi/2\beta}) \quad \text{as } t \rightarrow \infty.$$

**PROOF.** Integration of (2.18) [see Gradshteyn and Ryzhik (1980), formula 6.618-4] yields

$$(2.20) \quad \begin{aligned} & \mathbf{P}\{T_\beta > t | R_0 = \rho, \theta_0 = \alpha\} \\ &= \frac{\rho e^{-\rho^2/4t}}{(2\pi t)^{1/2}} \sum_{k \in \mathbb{Z}} \frac{\sin[(2k + 1)\pi\alpha/\beta]}{2k + 1} \left[ \mathbf{I}_{|2k + 1|\pi/2\beta - 1/2}(\rho^2/4t) \right. \\ & \quad \left. + \mathbf{I}_{|2k + 1|\pi/2\beta + 1/2}(\rho^2/4t) \right]. \end{aligned}$$

Taking  $\theta_0 = \beta/2$  and replacing each Bessel function by its series expansions (2.3), one obtains a double series that is clearly uniformly convergent for  $\rho^2/8t < 1$ , and (2.19) follows.

**Acknowledgments.** This note is the result of developing some ideas of E. Cabaña. He also discussed them with L. Shepp, who kindly agreed that I quote related communication between them. In particular, I know he did some of the calculations involved here.

#### REFERENCES

- FELLER, W. (1971). *An Introduction to Probability Theory and Its Applications* 2. Wiley, New York.
- GRADSHTEYN, I. S. and RYZHIK, I. M. (1980). *Table of Integrals, Series and Products*, 4th ed. Academic, New York.
- McKEAN, H. P. (1969). *Stochastic Integrals*. Academic, New York.
- PITMAN, J. W. and YOR, M. (1986). Asymptotic laws of planar Brownian motion. *Ann. Probab.* **14** 733–779.
- PITMAN, J. W. and YOR, M. (1989). Further asymptotic laws of planar brownian motion. *Ann. Probab.* **17** 965–1011.
- SHEPP, L. (1981). Private communication.
- SPITZER, F. (1958). Some theorems concerning two dimensional Brownian motion. *Trans. Amer. Math. Soc.* **87** 187–197.
- WATSON, G. N. (1966). *A Treatise on the Theory of Bessel Functions*. Cambridge Univ. Press.
- YOR, M. (1980). Loi de l'indice du lacet Brownien et distribution de Hartman-Watson. *Z. Wahrsch. Verw. Gebiete* **53** 71–95.

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