

CERTAIN METRICS ON A PRINCIPAL FIBER BUNDLE AND VARIATIONAL PROBLEMS

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Abstract. We define a functional constructed from the scalar curvature of a certain metric on a principal fiber bundle and obtain some equations which correspond to the Einstein field equation, the Yang-Mills equation and the Brans-Dicke type wave equation by variations of this functional.

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§0. Introduction

The first model of the natural unification of gauge fields and gravitation goes back to the five-dimensional model of Kaluza and Klein. Their model extends in a reasonably straightforward way to the case of gauge potentials on principal fiber bundles with arbitrary structure groups. Let P be a principal bundle over a semi-Riemannian manifold (M, g) with a structure group G . If ω is a connection 1-form on P and k_0 is an ad -invariant metric on the Lie algebra of G , then a metric h on P is constructed from g , ω and k_0 . It is called a bundle metric. In this case, the Einstein field equation and the Yang-Mills equation arise from a single variational principle, see [1], for example.

In this paper, we assume that an ad -invariant metric depends on a point of M . Let k be a map from M to the set of all ad -invariant metrics on the Lie algebra of G . In particular we consider the case where k is constructed from a fixed ad -invariant metric and a positive function on M . Physically this scalar field gives scales of the internal spaces. When G is compact and its Lie algebra is simple, a positive definite ad -invariant metric is unique up to multiplication

by a constant [3]. We give a metric h on P similar to a bundle metric using such a map k . Because the projection $\pi : (P, h) \rightarrow (M, g)$ is a semi-Riemannian submersion, geometrical quantities are described by the fundamental tensors defined in [4].

We define a functional constructed from a scalar curvature of (P, h) . By demanding that the integral of this functional be stationary under variations of the metric on M , we obtain the equation correspondence to the Einstein field equation. Similarly, variations of the connection 1-form lead to the equation correspondence to the Yang-Mills equation. Moreover we get the Brans-Dicke type wave equation [2] for a scalar field on M by variations of the function on M .

In Section 1, we will prepare the notation and terminology used in this paper. Section 2 is devoted to compute the fundamental tensors. In Section 3, using the lemmas in Section 2, the curvatures of (P, h) will be calculated. We will define some functional on M from the scalar curvature of h and consider variational problems with respect to the metric, connection and scalar field in Section 4. Finally in Section 5, the results in Section 4 will be applied to cosmology.

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§1. Preliminaries

Throughout this paper, all objects are assumed to be smooth. Let G be a Lie group and \mathcal{G} its Lie algebra. Let P be a principal G -bundle over a manifold M and $\pi : P \rightarrow M$ the projection. We define the vertical space of TP by $\mathcal{V}(P) := \text{Ker}\pi_*$, where TP is the tangent bundle of P and π_* is the differential map of π . The set of connection 1-forms on P is denoted by $\mathcal{C}(P)$. For $\omega \in \mathcal{C}(P)$, we define the horizontal space of TP by $\mathcal{H}(P) := \text{Ker}\omega$. Then we have $TP = \mathcal{V}(P) \oplus \mathcal{H}(P)$ (direct sum). Let \mathcal{V} (resp. \mathcal{H}) be the projection of TP onto $\mathcal{V}(P)$ (resp. $\mathcal{H}(P)$). For a vector field E on M , the horizontal lift of E is denoted by \tilde{E} or E^\sim . For $A \in \mathcal{G}$, the fundamental vector field induced from A is denoted by A^* . For a vector space W , the set of W -valued k -forms on P is denoted by $\Lambda^k(P, W)$. The set of all smooth functions on a manifold M is denoted by $C^\infty(M)$.

For $\varphi \in \Lambda^i(P, \mathcal{G})$ and $\psi \in \Lambda^j(P, \mathcal{G})$, we define $[\varphi, \psi] \in \Lambda^{i+j}(P, \mathcal{G})$ by

$$[\varphi, \psi](X_1, \dots, X_{i+j}) = \frac{1}{i!j!} \sum_{\sigma} (-1)^\sigma [\varphi(X_{\sigma(1)}, \dots, X_{\sigma(i)}), \psi(X_{\sigma(i+1)}, \dots, X_{\sigma(i+j)})],$$

where the sum is over the set of all permutations σ of $1, \dots, i + j$ and $(-1)^\sigma = \pm 1$ is the sign of σ . For $\omega \in \mathcal{C}(P)$ and $\tau \in \Lambda^i(P, \mathcal{G})$, we define $\tau^H \in \Lambda^i(P, \mathcal{G})$ by $\tau^H(X_1, \dots, X_i) = \tau(\mathcal{H}X_1, \dots, \mathcal{H}X_i)$ and the exterior covariant derivative of τ by $D^\omega \tau := (d\tau)^H$. The curvature form $\Omega \in \Lambda^2(P, \mathcal{G})$ is defined by $\Omega := D^\omega \omega$. The equation $\Omega = d\omega + (1/2)[\omega, \omega]$ is called the structure equation. From the structure equation, we see that

$$(1.1) \quad \Omega(X, Y) = -\omega([X, Y]) \quad \text{for horizontal vector fields } X \text{ and } Y$$

and

$$(1.2) \quad d\Omega = [\Omega, \omega].$$

The set of all metrics on a manifold M is denoted by $\mathcal{M}(M)$. For $a \in G$, let $Ad_a : G \rightarrow G$ be the adjoint isomorphism given by $Ad_a(b) = aba^{-1}$ and $ad(a) : \mathcal{G} \rightarrow \mathcal{G}$ the induced isomorphism of \mathcal{G} , that is, $ad(a) = (Ad_a)_*e$. The set of all ad -invariant metrics on \mathcal{G} is denoted by $\mathcal{M}_{ad}(\mathcal{G})$. For $k_0 \in \mathcal{M}_{ad}(\mathcal{G})$, we see that

$$(1.3) \quad k_0([A, B], C) + k_0(B, [A, C]) = 0 \text{ for } A, B, C \in \mathcal{G}.$$

§2. Fundamental tensors

Let P be a principal G -bundle over a manifold M and $\pi : P \rightarrow M$ the projection. We define a metric h on P as follows.

Definition. For $k : M \rightarrow \mathcal{M}_{ad}(\mathcal{G})$, $g \in \mathcal{M}(M)$ and $\omega \in \mathcal{C}(P)$, we define $h \in \mathcal{M}(P)$ by

$$h(E, F) = g(\pi_*E, \pi_*F) + (k \circ \pi)(\omega(E), \omega(F))$$

for any tangent vector E and F of P . When k is a constant map, it is called a bundle metric.

If P is the semi-Riemannian manifold with the metric h as above, then $\pi : (P, h) \rightarrow (M, g)$ is a semi-Riemannian submersion. The tensors T and A are defined for arbitrary vector fields E and F by

$$T_E F := \mathcal{H}\nabla_{\mathcal{V}E}(\mathcal{V}F) + \mathcal{V}\nabla_{\mathcal{V}E}(\mathcal{H}F)$$

and

$$A_E F := \mathcal{V}\nabla_{\mathcal{H}E}(\mathcal{H}F) + \mathcal{H}\nabla_{\mathcal{H}E}(\mathcal{V}F),$$

where ∇ is the covariant derivative of (P, h) . They are called the fundamental tensors in [4] and [5]. For a fixed ad -invariant metric k_0 on \mathcal{G} and a smooth function $K > 0$ on M , we set $k = \varepsilon K^2 k_0$ ($\varepsilon = \pm 1$) in the definition of h and consider only this case in the present paper. We write $\bar{K} := K \circ \pi$ for simplicity. If the Lie algebra of a compact Lie group is simple, then the positive definite ad -invariant metric is unique up to multiplication by a constant [3].

To compute the fundamental tensors and the curvature tensor, we define some differential operators in the usual way. For a function f on a manifold, $\text{grad} f$ is the gradient vector field of f , H^f is the Hessian of f and Δf is the Laplacian of f defined by $\Delta f = -\text{div}(\text{grad} f)$, where div is a divergence. We have the following lemma.

Lemma 2.1. *If U, V are vertical and X, Y are horizontal, then*

$$(2.1) \quad T_U V = -\varepsilon \bar{K} k_0(\omega(U), \omega(V)) \text{grad} \bar{K},$$

$$(2.2) \quad T_U X = \frac{X \bar{K}}{\bar{K}} U = \frac{h(\text{grad} \bar{K}, X)}{\bar{K}} U,$$

$$(2.3) \quad A_X Y = \frac{1}{2} \mathcal{V}[X, Y]$$

and

$$(2.4) \quad A_X U = -\frac{\varepsilon}{2} \bar{K}^2 \overline{\Omega^{\omega(U)}(X)},$$

where $\overline{\Omega^{\omega(U)}(X)}$ is defined by $h(\overline{\Omega^{\omega(U)}(X)}, E) = k_0(\omega(U), \Omega(E, X))$ for any vector field E on P .

Moreover, to compute the curvature of (P, h) , we show the following lemma by using Lemma 2.1.

Lemma 2.2. *If U, V, W are vertical and X, Y, Z are horizontal, then*

$$(2.5) \quad \mathcal{H}((\nabla_V T)_U W) = \frac{1}{2} \bar{K}^3 k_0(\omega(U), \omega(W)) \overline{\Omega^{\omega(V)}(\text{grad} \bar{K})},$$

$$(2.6) \quad \begin{aligned} \mathcal{H}((\nabla_X T)_V W) &= -\varepsilon \bar{K} k_0(\omega(V), \omega(W)) \mathcal{H}(\nabla_X \text{grad} \bar{K}) \\ &\quad + \varepsilon k_0(\omega(V), \omega(W)) h(X, \text{grad} \bar{K}) \text{grad} \bar{K}, \end{aligned}$$

$$(2.7) \quad \omega((\nabla_V A)_X Y) = \frac{1}{4}[\Omega, \omega](X, Y, V) - \frac{1}{2}(\nabla_V \Omega)(X, Y)$$

and

$$(2.8) \quad \omega((\nabla_Z A)_X Y) = -\frac{1}{2} \frac{h(Z, \text{grad} \bar{K})}{\bar{K}} \Omega(X, Y) - \frac{1}{2}(\nabla_Z \Omega)(X, Y),$$

where $(\nabla_E \Omega)$ is defined by $(\nabla_E \Omega)(F_1, F_2) = E\Omega(F_1, F_2) - \Omega(\nabla_E F_1, F_2) - \Omega(F_1, \nabla_E F_2)$.

§3. The Curvature tensors

Let $\hat{\nabla}^x$ be the covariant derivative of $\pi^{-1}(x)$ with respect to the induced metric from h and ∇^* the covariant derivative of (M, g) . Let R (resp. \hat{R}^x , R^*) be the curvature tensor of ∇ (resp. $\hat{\nabla}^x$, ∇^*). However we omit the superscript x for simplicity. Let $R_*(X, Y)Z$ be the horizontal vector field such that $\pi_*(R_*(X, Y)Z) = R^*(\pi_*X, \pi_*Y)\pi_*Z$ at each point of P . We can compute the curvature of (P, h) by using Lemmas 2.1 and 2.2.

Proposition 3.1. *For vertical vector fields U, V, W, F and horizontal vector fields X, Y, Z, H , we obtain*

$$(3.1) \quad h(R(U, V)W, F) = \frac{\varepsilon}{4} \bar{K}^2 k_0([\omega(U), \omega(V)], [\omega(W), \omega(F)]) \\ - \bar{K}^2 k_0(\omega(U), \omega(W)) k_0(\omega(V), \omega(F)) g(\text{grad} K, \text{grad} K) \circ \pi \\ + \bar{K}^2 k_0(\omega(V), \omega(W)) k_0(\omega(U), \omega(F)) g(\text{grad} K, \text{grad} K) \circ \pi,$$

$$(3.2) \quad h(R(U, V)W, X) = -\frac{1}{2} \bar{K}^3 k_0(\omega(U), \omega(W)) k_0(\omega(V), \Omega(\text{grad} \bar{K}, X)) \\ + \frac{1}{2} \bar{K}^3 k_0(\omega(V), \omega(W)) k_0(\omega(U), \Omega(\text{grad} \bar{K}, X)),$$

$$(3.3) \quad h(R(X, V)Y, W) = -\varepsilon \bar{K} k_0(\omega(V), \omega(W)) H^{\bar{K}}(X, Y) \\ + \frac{\varepsilon}{4} \bar{K}^2 k_0([\Omega, \omega](X, Y, V), \omega(W)) - \frac{\varepsilon}{2} \bar{K} k_0((\nabla_V \Omega)(X, Y), \omega(W))$$

$$\begin{aligned}
 & + \frac{1}{4} \bar{K}^2 h(\overline{\Omega^{\omega(V)}(X)}, \overline{\Omega^{\omega(W)}(Y)}), \\
 (3.4) \quad & h(R(X, Y)Z, V) = -\frac{\varepsilon}{2} \bar{K}^2 k_0((\nabla_Z \Omega)(X, Y), \omega(V)) \\
 & + \frac{\varepsilon}{2} \bar{K} k_0(\Omega(Y, Z), \omega(V)) h(X, \text{grad} \bar{K}) + \frac{\varepsilon}{2} \bar{K} k_0(\Omega(Z, X), \omega(V)) h(Y, \text{grad} \bar{K}) \\
 & - \varepsilon \bar{K} k_0(\Omega(X, Y), \omega(V)) h(Z, \text{grad} \bar{K})
 \end{aligned}$$

and

$$\begin{aligned}
 (3.5) \quad & h(R(X, Y)Z, H) = h(R_*(X, Y)Z, H) - \frac{\varepsilon}{2} \bar{K}^2 k_0(\Omega(X, Y), \Omega(Z, H)) \\
 & + \frac{\varepsilon}{2} \bar{K}^2 k_0(\Omega(Y, Z), \Omega(X, H)) + \frac{\varepsilon}{4} \bar{K}^2 k_0(\Omega(Z, X), \Omega(Y, H)).
 \end{aligned}$$

Note that for (3.1), we used (1.3) and

$$\hat{R}(A^*, B^*)C^* = \frac{1}{4} [[A^*, B^*], C^*] = \frac{1}{4} [[A, B], C]^*$$

for fundamental vector fields A^* , B^* and C^* .

By Proposition 3.1, we can compute the sectional curvature. Let Π_{ab} be the nondegenerate space spanned by tangent vectors a , b .

Corollary 3.2. *Let \mathcal{K} , \mathcal{K}_* and $\hat{\mathcal{K}}$ be the sectional curvature of (P, h) , (M, g) and the fibers with the induced metrics from h . If x and y are horizontal vectors at $p \in P$, and v and w are vertical, then*

$$(3.6) \quad \mathcal{K}(\Pi_{vw}) = \hat{\mathcal{K}}(\Pi_{vw}) - \frac{1}{4} \frac{g_{\pi(p)}(\text{grad} K, \text{grad} K)}{K^2(\pi(p))},$$

$$\begin{aligned}
 (3.7) \quad & \mathcal{K}(\Pi_{xv}) = -\frac{g_{\pi(p)}((\nabla_{x_*}^* \text{grad} K), x_*)}{K(\pi(p))g_{\pi(p)}(x_*, x_*)} \\
 & + \frac{\varepsilon K^2(\pi(p))}{4} \frac{g_{\pi(p)}(\overline{\pi_* \Omega^{\omega(v)}(x)}, \overline{\pi_* \Omega^{\omega(v)}(x)})}{g_{\pi(p)}(x_*, x_*)k_0(\omega(v), \omega(v))}
 \end{aligned}$$

and

$$(3.8) \quad \mathcal{K}(\Pi_{xy}) = \mathcal{K}_*(\Pi_{x_*y_*}) - \frac{3}{4} \frac{\varepsilon K^2(\pi(p))k_0(\Omega(x, y), \Omega(x, y))}{g_p(x_*, x_*)g_p(y_*, y_*) - g_p(x_*, y_*)^2},$$

where

$$\hat{\mathcal{K}}(\Pi_{vw}) = \frac{\varepsilon k_0([\omega(v), \omega(w)], [\omega(v), \omega(w)])}{K^2(\pi(p))(k_0(\omega(v), \omega(v))k_0(\omega(w), \omega(w)) - k_0(\omega(v), \omega(w))^2)}$$

and $x_* = \pi_*x$ and $y_* = \pi_*y$.

Next we calculate the Ricci and scalar curvatures of (P, h) by using Proposition 3.1. Especially we will form the functional from scalar curvature h in the next section. Let n (resp. l) be the dimension of M (resp. G). Assume that E_{1*}, \dots, E_{n*} is orthonormal base fields relative to g on a neighborhood $\mathcal{U} \subset M$ and E_1, \dots, E_n their horizontal lifts. Let e_1, \dots, e_l be an orthonormal base on \mathcal{G} relative to the fixed metric k_0 and we set $E_{n+1} := \bar{K}^{-1}e_1^*, \dots, E_{n+l} := \bar{K}^{-1}e_l^*$. Then $E_1, \dots, E_n, E_{n+1}, \dots, E_{n+l}$ is orthonormal base fields on $\pi^{-1}(\mathcal{U}) \subset P$ with respect to h . The indices i, j, \dots (resp. α, β, \dots) range from 1 to n (resp. from $n+1$ to $n+l$) and we set $\varepsilon_i := h(E_i, E_i) = g(E_{i*}, E_{i*})$ and $\varepsilon_\alpha := h(E_\alpha, E_\alpha) = \varepsilon k_0(e_\alpha, e_\alpha)$. Let Ric (resp. Ric^*) be the Ricci tensor of h (resp. g) and $(Ric)_*$ the symmetric 2-form on P such that $(Ric)_*(X, Y) = Ric^*(\pi_*X, \pi_*Y)$ at each point of P .

Proposition 3.3. *If V, W are vertical and X, Y are horizontal, then*

$$(3.9) \quad Ric(V, W) = \varepsilon \bar{K} k_0(\omega(V), \omega(W))((\Delta K) \circ \pi) + \frac{1}{4} \bar{K}^4 \sum_i \varepsilon_i h(\overline{\Omega^{\omega(V)}(E_i)}, \overline{\Omega^{\omega(W)}(E_i)}) + \frac{\varepsilon}{4} \sum_\alpha \varepsilon_\alpha k_0([\omega(V), e_\alpha], [\omega(W), e_\alpha]) - \varepsilon(l-1)k_0(\omega(V), \omega(W))g(\text{grad}K, \text{grad}K) \circ \pi,$$

$$(3.10) \quad Ric(V, X) = \frac{\varepsilon}{2} \bar{K}^2 \sum_i \varepsilon_i k_0((\nabla_{E_i} \Omega)(X, E_i), \omega(V)) - \frac{l+2}{2} \varepsilon \bar{K} k_0(\omega(V), \Omega(\text{grad} \bar{K}, X))$$

and

$$(3.11) \quad Ric(X, Y) = (Ric)_*(X, Y) - \frac{\varepsilon}{2} \bar{K}^2 \sum_i \varepsilon_i k_0(\Omega(X, E_i), \Omega(Y, E_i)) + \varepsilon \sum_\alpha \varepsilon_\alpha \left\{ \frac{1}{4} k_0([\Omega, \omega](X, Y, e_\alpha), e_\alpha) - \frac{1}{2} k_0((\nabla_{e_\alpha} \Omega)(X, Y), e_\alpha) \right\} - \frac{l}{\bar{K}} H^{\bar{K}}(X, Y).$$

By Proposition 3.3, we have

Proposition 3.4. *Let S (resp. S^*) be the scalar curvature of (P, h) (resp. (M, g)). Then*

$$(3.12) \quad S = S^* \circ \pi - \frac{\varepsilon}{4} \bar{K}^2 \sum_{i,j} \varepsilon_i \varepsilon_j k_0(\Omega(E_i, E_j), \Omega(E_i, E_j)) + 2l \frac{(\Delta K) \circ \pi}{\bar{K}} - l(l-1) \frac{g(\text{grad}K, \text{grad}K) \circ \pi}{\bar{K}^2} + \frac{1}{4} \frac{\varepsilon}{\bar{K}^2} \sum_{\alpha, \beta} \varepsilon_\alpha \varepsilon_\beta k_0([e_\alpha, e_\beta], [e_\alpha, e_\beta]).$$

§4. Variational problems

In this section, we consider the variational problems for the integral of a functional constructed from the scalar curvature of (P, h) . Let $\bar{\Lambda}^i(P, \mathcal{G})$ be the space of \mathcal{G} -valued i -forms φ on P such that $R_{a*}\varphi = ad(a^{-1})\varphi$ and $\varphi(X_1, \dots, X_i) = 0$ when one of X_1, \dots, X_i is vertical. For $\tau \in \bar{\Lambda}^i(P, \mathcal{G})$, we have $D^\omega\tau = d\tau + [\omega, \tau]$. The metric g_x on the tangent space at $x \in M$ induces the metric \bar{g}_p on horizontal subspace $\mathcal{H}(P)_p \subset T_pP$ ($p \in \pi^{-1}(x)$) via the isomorphism $\pi_*|_{\mathcal{H}(P)_p} : \mathcal{H}(P)_p \rightarrow T_xM$ (i.e., $\bar{g}_p(X, Y) := g_x(\pi_*X, \pi_*Y)$ for $X, Y \in \mathcal{H}(P)_p$). Let $\tilde{\mu}_p$ be the volume element on $\mathcal{H}(P)_p$ relative to this induced metric and we can define the star operator $\tilde{*}_p : \Lambda^i(\mathcal{H}(P)_p) \rightarrow \Lambda^{n-i}(\mathcal{H}(P)_p)$ ($n = \dim M$). Moreover we define $\tilde{*} : \bar{\Lambda}^i(P, \mathcal{G}) \rightarrow \bar{\Lambda}^{n-i}(P, \mathcal{G})$ by setting (for $\varphi \in \bar{\Lambda}^i(P, \mathcal{G})$) $(\tilde{*}\varphi)_p$ equal to the unique extension of $\tilde{*}_p(\varphi|_{\mathcal{H}(P)_p})$ to a \mathcal{G} -valued $(n-i)$ -form vanishing on vertical vectors. Let $\partial_1, \dots, \partial_n$ be coordinate vector fields on $\mathcal{U} \subset M$. The covariant codifferential $\delta^\omega : \bar{\Lambda}^i(P, \mathcal{G}) \rightarrow \bar{\Lambda}^{i-1}(P, \mathcal{G})$ is defined, for $\varphi \in \bar{\Lambda}^i(P, \mathcal{G})$, by $\delta^\omega(\varphi) := -(-1)^g(-1)^{n(i+1)}\tilde{*}D^\omega(\tilde{*}\varphi)$, where $(-1)^g$ is the sign of determinant of the matrix $(g(\partial_l, \partial_m))$. The self-action of the connection ω relative to the fixed ad -invariant metric k_0 is defined by

$$\mathcal{S}_0(g, \omega) := -\frac{1}{2}(\bar{g}k_0)(\Omega, \Omega) = -\frac{1}{4}g^{hj}g^{im}k_0(\Omega(\tilde{\partial}_h, \tilde{\partial}_i), \Omega(\tilde{\partial}_j, \tilde{\partial}_m)),$$

where $\bar{g}k_0$ is the metric on $\bar{\Lambda}^i(P, \mathcal{G})$ induced from \bar{g} and k_0 . Note that $\mathcal{S}_0(g, \omega)$ is a smooth function on M . The ad -invariant metric k_0 induces the bi-invariant metric \bar{k}_0 on G as follows. For $a \in G$ and $A, B \in T_aG$, we set $\bar{k}_0(A, B) := k_0(L_{a*}^{-1}A, L_{a*}^{-1}B)$, where L_a is the left action on G . Then (G, \bar{k}_0) has the constant scalar curvature

$$c_0 = \frac{1}{4} \sum_{\alpha, \beta} \varepsilon_\alpha \varepsilon_\beta k_0([e_\alpha, e_\beta], [e_\alpha, e_\beta]).$$

Hence the scalar curvature of (P, h) is described by

$$\begin{aligned} S &= S^* \circ \pi + (\varepsilon K^2 \mathcal{S}_0(g, \omega)) \circ \pi + \frac{\varepsilon c_0}{\bar{K}^2} \\ &\quad + 2l \frac{(\Delta K) \circ \pi}{\bar{K}} - l(l-1) \frac{g(\text{grad}K, \text{grad}K) \circ \pi}{\bar{K}^2}. \end{aligned}$$

We define a map $\mathcal{L} : \mathcal{M}(M) \times \mathcal{C}(P) \times C^\infty(M)^+ \longrightarrow C^\infty(M)$ by

$$\begin{aligned} \mathcal{L}(g, \omega, K) &:= \{S^* + \varepsilon K^2 \mathcal{S}_0(g, \omega) + \frac{\varepsilon c_0}{K^2} + 2l \frac{(\Delta K)}{K} - l(l-1) \frac{g(\text{grad}K, \text{grad}K)}{K^2}\} K^l \\ &= S_* K^l, \end{aligned}$$

where

$$S_* := S_*(g, \omega, K) := S^* + \varepsilon K^2 \mathcal{S}_0(g, \omega) + \frac{\varepsilon c_0}{K^2} + 2l \frac{(\Delta K)}{K} - l(l-1) \frac{g(\text{grad}K, \text{grad}K)}{K^2}$$

and $C^\infty(M)^+$ is the set of all positive functions on M . The notation $U \subset\subset M$ means that U is an open subset of M with compact closure. The volume element relative to a metric g is denoted by μ_g . The variational problems for the integral of the scalar curvature of h reduces to those for the integral of $\mathcal{L}(g, \omega, K)$ since $S\mu_h = SK^l \pi^* \mu_g \wedge \mu_{\tilde{k}_0}$, where $\mu_{\tilde{k}_0}$ is the volume element induced from k_0 .

At first, we consider variations of the metric. Let $S^2(M)$ be the set of all symmetric tensors on M . For $g \in \mathcal{M}(M)$, $u \in S^2(M)$, and $t \in \mathbf{R}$, we set $g(t) := g + tu$. For small $t \in \mathbf{R}$, $g(t)$ is in $\mathcal{M}(M)$. Then we denote the curvature tensor, gradient and Laplacian relative to $g(t)$ by $R^*(t)$, $\text{grad}(t)$ and $\Delta(t)$, respectively. We set $g_{ij}(t) := g(t)(\partial_i, \partial_j)$ and define $R^{*i}_{jkl}(t)$ by the components of the curvature tensor of $g(t)$. We write $g_{ij} = g_{ij}(0)$, $R^{*i}_{jkl} = R^{*i}_{jkl}(0)$, etc. The indices are raised and lowered by the initial metric g . For $f \in C^\infty(M)$, we have

$$(4.1) \quad \int_U f g^{ij} R^{*k}_{ijk}{}'(0) \mu_g = \int_U \{ (f_{,k;i}) + (\Delta f) g_{ik} \} u^{ik} \mu_g,$$

$$(4.2) \quad \frac{d}{dt} (\Delta(t)f) |_{t=0} = ((u^{ki})(f_{,k}))_{;i} - \frac{1}{2} ((u^j_j)(f_{,k}))^{;k} - \frac{1}{2} u^j_j \Delta f$$

and

$$(4.3) \quad \frac{d}{dt} g(t)(\text{grad}(t)f, \text{grad}(t)f) |_{t=0} = u^{ij}(f_{,i})(f_{,j}),$$

where a prime denotes the derivative with respect to the parameter t .

Using equations above, we obtain the following theorem.

Theorem 4.1. (Einstein field equation). For all $U \subset\subset M$ and all $u \in S^2(M)$ with support in U , the equation

$$\frac{d}{dt} \int_U \mathcal{L}(g + tu, \omega, K) \mu_{g(t)} = 0 \quad \text{at } t = 0$$

holds if and only if

$$(4.4) \quad R^*_{ij} - \frac{1}{2} S^* g_{ij} = \frac{1}{2} \varepsilon K^2 k_0 (\Omega_{hi}, \Omega_{mj}) g^{hm} + \frac{1}{2} \varepsilon K^2 \mathcal{S}_0(g, \omega) g_{ij} \\ + \frac{1}{2} \frac{\varepsilon c_0}{K^2} g_{ij} + \frac{l}{K} (K_{,i;j} + \Delta K g_{ij}) - \frac{1}{2} l(l-1) \frac{g(\text{grad}K, \text{grad}K)}{K^2} g_{ij},$$

where R^*_{ij} are the components of the Ricci curvature of (M, g) .

Proof. At first, about the first term, we have

$$\begin{aligned} & \frac{d}{dt} \int_U K^l S^*(t) \mu_{g(t)} \Big|_{t=0} \\ &= \int_U K^l (-R^*_{ij} + \frac{1}{2} S^* g_{ij}) u^{ij} \mu_g + \int_U K^l g^{ij} R^{*k}_{ijk}{}'(0) \mu_g. \end{aligned}$$

From

$$\begin{aligned} (K^l)_{,ij} &= K^l \{ l(l-1) \frac{(K_{,i})(K_{,j})}{K^2} + l \frac{K_{,ij}}{K} \}, \\ \Delta(K^l) &= K^l \{ l \frac{\Delta K}{K} - l(l-1) \frac{g(\text{grad}K, \text{grad}K)}{K^2} \} \end{aligned}$$

and (4.1), we have

$$\begin{aligned} \int_U K^l g^{ij} R^{*k}_{ijk}{}'(0) \mu_g &= \int_U K^l \{ l(l-1) \frac{(K_{,i})(K_{,j})}{K^2} + l \frac{K_{,ij}}{K} + l \frac{\Delta K}{K} g_{ij} \\ &\quad - l(l-1) \frac{g(\text{grad}K, \text{grad}K)}{K^2} g_{ij} \} u^{ij} \mu_g. \end{aligned}$$

For the second and third term, by similar calculations in 9.3.3 Theorem in [1], we get

$$\begin{aligned} & \frac{d}{dt} \int_U \varepsilon K^{l+2} \mathcal{S}_0(g(t), \omega) \mu_{g(t)} \Big|_{t=0} \\ &= \int_U \{ K^l (\frac{1}{2} \varepsilon K^2 g^{hm} k_0(\Omega_{hi}, \Omega_{mj}) + \frac{1}{2} \varepsilon K^2 \mathcal{S}_0(g, \omega) g_{ij}) u^{ij} \} \mu_g \end{aligned}$$

and

$$\frac{d}{dt} \int_U K^l \frac{\varepsilon c_0}{K^2} \mu_{g(t)} \Big|_{t=0} = \int_U K^l (\frac{1}{2} \frac{\varepsilon c_0}{K^2} g_{ij}) u^{ij} \mu_g.$$

For the fourth term, from (4.2), it follows that

$$\begin{aligned} & \frac{d}{dt} \int_U K^{l-1} ((\Delta(t))K) \mu_{g(t)} \Big|_{t=0} \\ &= \int_U K^{l-1} \{ ((u^{ki})(K_{,k}))_{,i} - \frac{1}{2} ((u^j)(K_{,k}))^{,k} - \frac{1}{2} u^j_{,j} (\Delta K) \} \mu_g \\ &\quad + \int_U K^{l-1} (\Delta K) (\frac{1}{2} u^i_{,i}) \mu_g \\ &= \int_U K^l \{ -(l-1) \frac{(K_{,i})(K_{,j})}{K^2} + \frac{1}{2} (l-1) \frac{g(\text{grad}K, \text{grad}K)}{K^2} g_{ij} \} u^{ij} \mu_g. \end{aligned}$$

For last term, by (4.3), we have

$$\begin{aligned} & \frac{d}{dt} \int_U K^l \frac{g(t)(\text{grad}(t)K, \text{grad}(t)K)}{K^2} \mu_{g(t)} \Big|_{t=0} \\ &= \int_U K^l \{ -\frac{1}{K^2} (K_{,i})(K_{,j}) + \frac{1}{2} \frac{g(\text{grad}K, \text{grad}K)}{K^2} g_{ij} \} u^{ij} \mu_g. \end{aligned}$$

Piecing these results together, we see that (4.4) holds if and only if g is stationary relative to \mathcal{L} for fixed ω and K . Q.E.D.

Next, we consider variations of the connection. For $\omega \in \mathcal{C}(P)$, $\tau \in \bar{\Lambda}^1(P, \mathcal{G})$, and $t \in \mathbf{R}$, we set $\omega(t) := \omega + t\tau$. Then $\omega(t)$ is in $\mathcal{C}(P)$ for all $t \in \mathbf{R}$. Let $\Omega(t)$ be the curvature form of $\omega(t)$. Let $U \subset\subset M$, and suppose that $\alpha \in \bar{\Lambda}^k(P, \mathcal{G})$, while $\beta \in \bar{\Lambda}^{k+1}(P, \mathcal{G})$. Assume that the projected support of α is contained in U . Then

$$(4.5) \quad \int_U (\bar{g}k_0)(D^\omega \alpha, \beta) \mu_g = \int_U (\bar{g}k_0)(\alpha, \delta^\omega \beta) \mu_g.$$

For the curvature form of $\omega(t) = \omega + t\tau$, from the structure equation, we have

$$(4.6) \quad \frac{d}{dt} \Omega(t) |_{t=0} = d\tau + [\omega, \tau] = D^\omega \tau.$$

Theorem 4.2. (*Yang-Mills equation*). For all $U \subset\subset M$ and all $\tau \in \bar{\Lambda}^1(P, \mathcal{G})$ with projected support in U , the equation

$$\frac{d}{dt} \int_U \mathcal{L}(g, \omega + t\tau, K) \mu_g = 0 \quad \text{at } t = 0$$

holds if and only if

$$(4.7) \quad \delta^\omega (\bar{K}^{l+2} \Omega) = 0,$$

or equivalently

$$(4.7)' \quad \delta^\omega \Omega = \frac{l+2}{\bar{K}} \Omega(\text{grad} \bar{K}, \cdot).$$

Proof. From (4.5) and (4.6), it follows that

$$\begin{aligned} & \frac{d}{dt} \int_U \mathcal{L}(g, \omega + t\tau, K) \mu_g |_{t=0} = - \int_U \varepsilon K^{l+2} (\bar{g}k_0) \left(\frac{d}{dt} \Omega(t) |_{t=0}, \Omega \right) \mu_g \\ & = - \int_U \varepsilon K^{l+2} (\bar{g}k_0) (D^\omega \tau, \Omega) \mu_g = - \int_U (\bar{g}k_0) (\tau, \delta^\omega (\varepsilon \bar{K}^{l+2} \Omega)) \mu_g. \end{aligned}$$

Hence we see that the equation (4.7) holds if and only if ω is stationary relative to \mathcal{L} for fixed g and K . Q.E.D.

Finally, we consider variations of the positive function. We start with the following lemma.

Lemma 4.3. For all $U \subset\subset M$ and all $L \in C^\infty(M)$ with support in U , the equation

$$(4.8) \quad \frac{d}{dt} \int_U K^l S_*(g, \omega, K + tL) \mu_g = 0 \quad \text{at } t = 0$$

holds if and only if

$$\varepsilon K^{l+1} \mathcal{S}_0(g, \omega) - \varepsilon K^{l-3} c_0 - l K^{l-2} (\Delta K) + l(l-1) K^{l-3} g(\text{grad}K, \text{grad}K) = 0.$$

Proof. For $K \in C^\infty(M)^+$ and $L \in C^\infty(M)$, from a straightforward calculation, it follows that

$$\frac{d}{dt} \varepsilon (K + tL)^2 \mathcal{S}_0(g, \omega) |_{t=0} = 2\varepsilon K L \mathcal{S}_0(g, \omega),$$

$$\frac{d}{dt} \frac{c_0}{(K + tL)^2} |_{t=0} = -\frac{2c_0 L}{K^3},$$

$$\frac{d}{dt} \frac{\Delta(K + tL)}{K + tL} |_{t=0} = \frac{\Delta L}{K} - \frac{(\Delta K)}{K^2} L$$

and

$$\frac{d}{dt} \frac{g(\text{grad}(K + tL), \text{grad}(K + tL))}{(K + tL)^2} |_{t=0} = \frac{2g(\text{grad}K, \text{grad}L)}{K^2} - \frac{2g(\text{grad}K, \text{grad}K)L}{K^3}.$$

Moreover by Green's theorem, we obtain

$$\begin{aligned} & \int_U K^{l-1} (\Delta L) \mu_g = \int_U g(\text{grad}(K^{l-1}), \text{grad}L) \mu_g = \int_U \Delta(K^{l-1}) L \mu_g \\ & = \int_U \{(l-1)K^{l-2} \Delta K - (l-1)(l-2)K^{l-3} g(\text{grad}K, \text{grad}K)\} L \mu_g \end{aligned}$$

and

$$\begin{aligned} & \int_U K^{l-2} g(\text{grad}K, \text{grad}L) \mu_g = \int_U g(K^{l-2} \text{grad}K, \text{grad}L) \mu_g \\ & = - \int_U \text{div}(K^{l-2} \text{grad}K) L \mu_g = - \int_U \{g(\text{grad}K^{l-2}, \text{grad}K) + K^{l-2} \text{div}(\text{grad}K)\} L \mu_g \\ & = \int_U \{-(l-2)K^{l-3} g(\text{grad}K, \text{grad}K) + K^{l-2} (\Delta K)\} L \mu_g. \end{aligned}$$

Hence, from these equations, we see that the equation (4.8) holds if and only if

$$\varepsilon K^{l+1} \mathcal{S}_0(g, \omega) - \varepsilon K^{l-3} c_0 - l K^{l-2} (\Delta K) + l(l-1) K^{l-3} g(\text{grad}K, \text{grad}K) = 0.$$

Q.E.D.

Theorem 4.4. (*Brans-Dicke type wave equation*). For all $U \subset\subset M$ and all $L \in C^\infty(M)$ with support in U , the equation

$$\frac{d}{dt} \int_U \mathcal{L}(g, \omega, K + tL) \mu_g = 0 \quad \text{at } t = 0$$

holds if and only if

$$(4.9) \quad \begin{aligned} lK^2 S^* + \varepsilon(l+2)K^4 \mathcal{S}_0(g, \omega) + \varepsilon(l-2)c_0 + 2l(l-1)K(\Delta K) \\ - l(l-1)(l-2)g(\text{grad}K, \text{grad}K) = 0. \end{aligned}$$

Proof. By Lemma 4.3, K is stationary relative to \mathcal{L} for fixed g and ω if and only if

$$\begin{aligned} 0 = & lK^{l-1} \left\{ S^* + \varepsilon K^2 \mathcal{S}_0(g, \omega) + \frac{\varepsilon c_0}{K^2} + 2l \frac{(\Delta K)}{K} - l(l-1) \frac{g(\text{grad}K, \text{grad}K)}{K^2} \right\} \\ & + 2\varepsilon K^{l+1} \mathcal{S}_0(g, \omega) - 2\varepsilon K^{l-3} c_0 + \{2l(l-1) - 2l - 2l(l-1)\} K^{l-2} (\Delta K) \\ & + \{-2l(l-1)(l-2) + 2l(l-1)(l-2) + 2l(l-1)\} K^{l-3} g(\text{grad}K, \text{grad}K) \end{aligned}$$

holds. Then we have

$$\begin{aligned} lK^2 S^* + \varepsilon(l+2)K^4 \mathcal{S}_0(g, \omega) & + \varepsilon(l-2)c_0 + 2l(l-1)K(\Delta K) \\ & - l(l-1)(l-2)g(\text{grad}K, \text{grad}K) = 0. \end{aligned}$$

Q.E.D.

By Theorems 4.1 and 4.4, we have the following corollary.

Corollary 4.5. *If the equation (4.4) and (4.9) hold, then*

$$(4.10) \quad (n+l-2)\{\varepsilon K^4 \mathcal{S}_0(g, \omega) - \varepsilon c_0 - lK \Delta K + l(l-1)g(\text{grad}K, \text{grad}K)\} = 0.$$

If $n+l > 2$, then the equation (4.10) reduces to

$$(4.11) \quad \varepsilon K^4 \mathcal{S}_0(g, \omega) - \varepsilon c_0 - lK \Delta K + l(l-1)g(\text{grad}K, \text{grad}K) = 0.$$

Proof. Contracting the equation (4.4) by g , we have

$$\begin{aligned} \left(1 - \frac{1}{2}n\right)S^* + \varepsilon\left(2 - \frac{n}{2}\right)K^2 \mathcal{S}_0(g, \omega) - \frac{1}{2}n \frac{\varepsilon c_0}{K^2} \\ - \frac{l}{K}(-\Delta K + n\Delta K) + \frac{1}{2}nl(l-1) \frac{g(\text{grad}K, \text{grad}K)}{K^2} = 0. \end{aligned}$$

From this equation and (4.9), it follows that

$$(n+l-2)\{\varepsilon K^4 \mathcal{S}_0(g, \omega) - \varepsilon c_0 - lK \Delta K + l(l-1)g(\text{grad}K, \text{grad}K)\} = 0.$$

Q.E.D.

§5. Cosmology

In this section, we assume that M is a warped product and satisfies the equations in the previous section. By using these equations, we will consider cosmology. Let M_S be an m -dimensional semi-Riemannian manifold. Let $f > 0$ be a smooth function on an interval I in \mathbf{R}_1^1 . Assume that M is the product manifold $I \times M_S$. Let p_I (resp. p_S) be the projection of M onto I (resp. M_S). The metric on M is defined by $g := p_I^* \sigma_I + (f \circ p_I)^2 p_S^* \sigma_S$, where σ_I and σ_S are the metric tensors on I and M_S , respectively. Especially, M is called a Robertson-Walker spacetime, if M_S is a connected 3-dimensional Riemannian manifold of constant curvature $\kappa = -1, 0$ or 1 , see [5], for example.

Let (x^0, x^1, \dots, x^m) be a coordinate system on $\mathcal{U} \subset M = I \times M_S$. We assume that the function K depend only on x^0 . We compute the curvatures of (M, g) . The indices A, B, \dots (resp. i, j, \dots) range from 1 to m (resp. from 0 to m). Then we have

$$g_{00} = -1, g_{AB} = f^2 \sigma_{AB} \text{ and } g^{00} = -1, g^{AB} = \frac{\sigma^{AB}}{f^2}$$

and

$$R_{00}^* = -m \frac{\ddot{f}}{f}, R_{0A}^* = 0 \text{ and } R_{AB}^* = \{f \ddot{f} + (m-1) \dot{f}^2\} \sigma_{AB} + \bar{R}_{AB},$$

where \bar{R}_{AB} are the components of the Ricci curvature of (M_S, σ_S) and $\sigma_{AB} = (\sigma_S)_{AB}$. Putting \bar{S} the scalar curvature of σ_S , the scalar curvature S^* is described by

$$S^* = 2m \frac{\ddot{f}}{f} + m(m-1) \frac{\dot{f}^2}{f^2} + \frac{\bar{S}}{f^2}.$$

Since the function K depends only on x^0 , we get

$$g(\text{grad}K, \text{grad}K) = -\dot{K}^2, \Delta K = \ddot{K} + m \frac{\dot{f}}{f} \dot{K}$$

and

$$K_{,ij} = \begin{cases} \ddot{K} & (i = j = 0) \\ 0 & (i = 0, j = A) \\ -f \dot{f} \ddot{K} \sigma_{AB} & (i = A, j = B) \end{cases}.$$

For the self-action, we have

$$\mathcal{S}_0(g, \omega) = -\frac{1}{4} g^{hj} g^{im} k_0(\Omega_{hi}, \Omega_{jm}) = \frac{1}{2} \frac{1}{f^2} a - \frac{1}{4} \frac{1}{f^4} b,$$

where $a := \sigma^{AB} k_0(\Omega_{0A}, \Omega_{0B})$ and $b := \sigma^{AB} \sigma^{CD} k_0(\Omega_{AC}, \Omega_{BD})$.

From Theorem 4.1, we have

Proposition 5.1. *The following equations hold.*

$$(5.1) \quad m(m-1)\frac{\dot{f}^2}{f^2} + \frac{\bar{S}}{f^2} = \frac{\varepsilon K^2}{2f^2}a + \frac{\varepsilon K^2}{4f^4}b - \frac{\varepsilon c_0}{K^2} \\ - 2ml\frac{\dot{f}\dot{K}}{fK} - l(l-1)\frac{\dot{K}^2}{K^2},$$

$$(5.2) \quad k_0(\Omega_{B0}, \Omega_{CA})\sigma^{BC} = 0$$

and

$$(5.3) \quad (1-m)f\ddot{f}\sigma_{AB} + (1-\frac{m}{2})(m-1)\dot{f}^2\sigma_{AB} + \bar{R}_{AB} - \frac{1}{2}\bar{S}\sigma_{AB} \\ = -\frac{1}{2}\varepsilon K^2 k_0(\Omega_{0A}, \Omega_{0B}) + \frac{1}{2}\varepsilon\frac{K^2}{f^2}k_0(\Omega_{CA}, \Omega_{DB})\sigma^{CD} + \frac{1}{4}\varepsilon K^2 a\sigma_{AB} - \frac{1}{8}\varepsilon\frac{K^2}{f^2}b\sigma_{AB} \\ + \frac{1}{2}\frac{\varepsilon c_0}{K^2}f^2\sigma_{AB} + lf^2\frac{\ddot{K}}{K}\sigma_{AB} + l(m-1)f\dot{f}\frac{\dot{K}}{K}\sigma_{AB} + \frac{1}{2}l(l-1)\frac{\dot{K}^2}{K^2}f^2\sigma_{AB}.$$

Contracting (5.3) by σ_S and from (5.1), we obtain

Corollary 5.2. *It follows that*

$$(5.4) \quad (1-m)m\frac{\ddot{f}}{f} = (\frac{m}{2}-1)\varepsilon\frac{K^2}{f^2}a + \frac{1}{4}\varepsilon\frac{K^2}{f^4}b \\ + \frac{\varepsilon c_0}{K^2} + ml\frac{\dot{f}\dot{K}}{fK} + ml\frac{\ddot{K}}{K} + l(l-1)\frac{\dot{K}^2}{K^2}.$$

From Theorems 4.2, 4.4 and Corollary 4.5, the following equations hold.

Proposition 5.3. *We have*

$$(5.5) \quad \delta^\omega\Omega = -(l+2)\frac{\dot{K} \circ \pi}{\bar{K}}\Omega(\partial_0, \cdot),$$

$$(5.6) \quad 2ml\frac{\ddot{f}}{f} + lm(m-1)\frac{\dot{f}^2}{f^2} + l\frac{\bar{S}}{f^2} + \frac{\varepsilon(l+2)K^2}{2f^2}a - \frac{\varepsilon(l+2)K^2}{4f^4}b \\ + \varepsilon(l-2)\frac{c_0}{K^2} + 2l(l-1)\frac{\ddot{K}}{K} + 2l(l-1)m\frac{\dot{f}\dot{K}}{fK} + l(l-1)(l-2)\frac{\dot{K}^2}{K^2} = 0$$

and

$$(5.7) \quad \frac{\varepsilon K^2}{2f^2}a - \frac{\varepsilon K^2}{4f^4}b - \varepsilon\frac{c_0}{K^2} - l\frac{\ddot{K}}{K} - ml\frac{\dot{f}\dot{K}}{fK} - l(l-1)\frac{\dot{K}^2}{K^2} = 0.$$

The equations (5.4) and (5.7) imply the following corollary.

Corollary 5.4. *It follows that*

$$(5.8) \quad (m-1)\left\{m\frac{\ddot{f}}{f} + \frac{\varepsilon}{2}a + l\frac{\ddot{K}}{K}\right\} = 0.$$

If $m > 1$, then we have

$$(5.9) \quad m\frac{\ddot{f}}{f} + \frac{\varepsilon}{2}a + l\frac{\ddot{K}}{K} = 0.$$

We consider the case of $K(x^0) = F(f(x^0))$, where F is a function on the set of all positive real numbers. Then we have the following equation from the (5.9).

Corollary 5.5. *We have*

$$(5.10) \quad \left(\frac{m}{f} + l\frac{F'}{F}\right)\ddot{f} + \frac{\varepsilon}{2}a + l\frac{F''}{F}f^2 = 0,$$

where $F' = dF/df$.

We assume that σ_S and k_0 are positive definite metrics and $\varepsilon = 1$, that is, M and P are Lorentz manifold. Usually, physicists consider this case in standard models. Then we have $a \geq 0$. Moreover, we assume $F' \leq 0$ and $F'' \geq 0$. The assumption $F' \leq 0$ means that fibers are contracting when the space M_S is expanding. From Corollary 5.5, we get the following corollary.

Corollary 5.6. *If $\ddot{f} \leq 0$ and $\{t \in I \mid \ddot{f}(t) = 0\}$ has no interior points, then we have*

$$-\frac{m}{l} \leq \frac{fF'}{F} \leq 0.$$

For example, when $K = f^\alpha$ ($\alpha \leq 0$) or $K = \exp(\beta f)$ ($\beta < 0$), this inequality reduces to $-(m/l) \leq \alpha \leq 0$, $-\frac{m}{l\beta} \geq f (> 0)$, respectively. When we refer to f as the scale factor, Corollary 5.6 indicates the relation among F , $\dim M_S$, $\dim G$ and the scale of the universe.

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