

Appendix D

Some flat metrics with Cauchy horizons.

In this Appendix we present a family of flat $n + 1$ dimensional space-times which have at least two non-isometric extensions across a smooth Cauchy horizon; this example is a generalization of Misner's model for the Taub-NUT space-times [84]. The space-times considered here do not have compact spacelike hypersurfaces, as opposed to Misner's example which, in $n+1$ dimensions, can be given $\mathbb{R} \times S^1 \times T^{n-1}$ topology, where T^ℓ is an ℓ dimensional torus. In view of the renewed interest in flat Lorentzian manifolds [130] [92] [83] it would be useful to understand the global structure of all flat Lorentzian manifolds, at least in 4 dimensions.

For $n \geq 2$ let (t, x, y) , $y = y^1$ if $n = 2$ or $y = (y^1, \dots, y^{n-1})$ otherwise, be coordinates in $n + 1$ dimensional Minkowski space,

$$ds^2 = -dt^2 + dx^2 + dy^2, \quad (\text{D.0.1})$$

with $dy^2 = (dy^1)^2$ if $n = 2$ and $dy^2 = d\bar{y}^2 = (dy^1)^2 + \dots + (dy^{n-1})^2$ otherwise. Let Λ_0 be a fixed Lorentz boost in the $t - x$ plane,

$$\Lambda_0 = \begin{bmatrix} \gamma_0 & \gamma_0\beta & 0 \\ \gamma_0\beta & \gamma_0 & 0 \\ 0 & 0 & id_{\mathbb{R}^{n-1}} \end{bmatrix},$$

$$\gamma_0 = \frac{1}{\sqrt{1 - \beta_0^2}}, \quad |\beta_0| < 1,$$

where $id_{\mathbb{R}^{n-1}}$ is the identity matrix in \mathbb{R}^{n-1} . Let \mathcal{H}_τ denote a hyperboloid in \mathbb{R}^{n+1} ,

$$\mathcal{H}_\tau = \{\tau = t^2 - x^2 - \vec{y}^2\},$$

with \mathcal{H}_0 — the light cone at the origin. Define \mathcal{M}_{Λ_0} as the quotient of the interior $I^+(0)$ of the solid future light cone $J^+(0)$ from the origin by the group generated by Λ_0 . In $I^+(0)$ we can introduce coordinates (τ, β, β_1) as follows:

$$\tau = t^2 - x^2 - y^2 - z^2 > 0,$$

$$t = ch \beta_1 ch \beta + \frac{(\tau - 1)e^{-\beta}}{2ch \beta_1}, \quad x = ch \beta_1 sh \beta - \frac{(\tau - 1)e^{-\beta}}{2ch \beta_1}, \quad (\text{D.0.2})$$

$$y = sh \beta_1 \text{ if } n = 2, \quad \vec{y} = sh \beta_1 \vec{\omega}, \quad \vec{\omega} \in S^{n-1}(1) \text{ if } n > 2, \quad (\text{D.0.3})$$

where $S^k(1)$ is the k -dimensional unit sphere. In terms of (D.0.2) the metric (D.0.1) takes the form

$$ds^2 = \gamma_{\mu\nu} dx^\mu dx^\nu = (\tau + sh^2 \beta_1) d\beta^2 + \frac{1 + \tau sh^2 \beta_1}{ch^2 \beta_1} d\beta_1^2 - d\tau d\beta - \frac{sh \beta_1}{ch \beta_1} d\tau d\beta_1 + \frac{2(\tau - 1) sh \beta_1}{ch \beta_1} d\beta_1 d\beta, \quad (\text{D.0.4})$$

and a term $sh^2 \beta_1 d\vec{\omega}^2$ has to be added to (D.0.4) if $n > 2$. One finds

$$\det \gamma_{\mu\nu} = -\frac{1}{4} ch^2 \beta_1 sh^{2(n-2)} \beta_1 \det h_{AB}^0, \quad (\text{D.0.5})$$

h_{AB}^0 — the round metric on a $n - 1$ dimensional sphere ($\det h_{AB}^0 = 1$ if $n = 2$). Equation (D.0.4) and the definition of \mathcal{M}_{Λ_0} show that (τ, β, β_1) can be used as coordinates on \mathcal{M}_{Λ_0} if β is identified with $\beta + \beta_0$. By (D.0.5) it follows that (D.0.4) gives an analytic extension of the flat metric on \mathcal{M}_{Λ_0} to the larger manifold M ,

$$\tau \in (-\infty, \infty), \quad \beta_1 \in (-\infty, \infty) \text{ for } n = 2, \quad \sinh \beta_1 \vec{\omega} \in \mathbb{R}^{n-1} \text{ for } n > 2, \quad (\text{D.0.6})$$

$$\beta \in [0, \beta_0]_{\text{mod } \beta_0}.$$

Another extension $(\tilde{M}, \tilde{\gamma})$ is obtained by defining coordinates $(\tau, \tilde{\beta}, \beta_1)$, with $\tilde{\beta} \in [0, \beta_0] \bmod \beta_0$ and τ, β_1 — as in (D.0.6), by

$$t = ch\beta_1 ch\tilde{\beta} + \frac{(\tau - 1)e^{\tilde{\beta}}}{2ch\beta_1}, \quad x = ch\beta_1 sh\tilde{\beta} + \frac{(\tau - 1)e^{\tilde{\beta}}}{2ch\beta_1}, \quad (\text{D.0.7})$$

y as in (D.0.3), which yields

$$ds^2 = \tilde{\gamma}_{\mu\nu} dx^\mu dx^\nu = (\tau + sh^2\beta_1) d\tilde{\beta}^2 + \frac{(1 + \tau sh^2\beta_1)}{ch^2\beta_1} d\beta_1^2 + d\tau d\tilde{\beta} \quad (\text{D.0.8})$$

$$- \frac{sh\beta_1}{ch\beta_1} d\tau d\beta_1 - \frac{2(\tau - 1)sh\beta_1}{ch\beta_1} d\tilde{\beta} d\beta_1, \quad (\text{D.0.9})$$

$$\det \tilde{\gamma}_{\mu\nu} = - \frac{ch^2\beta_1}{4} sh^{2(n-2)} \beta_1 \det h_{AB}^0.$$

There exists an isometry $\Phi : \{p \in M : \tau(p) > 0\} \rightarrow \tilde{M}$, where Φ is obtained by calculating $(\tau, \tilde{\beta}, \beta_1)$ as a function of (τ, β, β_1) from (D.0.2) and (D.0.7). From the equation

$$e^{\tilde{\beta}} = \frac{e^\beta ch^2\beta_1}{\tau + sh^2\beta_1},$$

it follows that $\tilde{\beta}$ blows up at $\tau = \beta_1 = 0$, thus Φ cannot be extended beyond $\tau = 0$. It should be noted that the map $\Psi : M \rightarrow \tilde{M}$ defined by

$$\tau \rightarrow \tau, \quad \beta \rightarrow -\tilde{\beta}, \quad \beta_1 \rightarrow \beta_1,$$

is a globally defined isometry, thus (M, γ) and $(\tilde{M}, \tilde{\gamma})$ are isometric. (M, γ) and $(\tilde{M}, \tilde{\gamma})$ are, however, *inequivalent* space-times from a Cauchy problem point of view for the following reason: let $i : \Sigma \rightarrow M$, $\tilde{i} : \tilde{\Sigma} \rightarrow \tilde{M}$ be embeddings such that M and \tilde{M} are developments of the Cauchy data set (Σ, g, K) , then there exists no isometry $\Xi : M \rightarrow \tilde{M}$ such that $\Xi \circ i = \tilde{i}$ (cf. [36]).

It should be noted that the space-times described above, restricted to the t - x plane, coincide with the so-called ‘‘Misner model’’ for the Taub-NUT space-time.