

## A CERTAIN SPACE-TIME METRIC AND SMOOTH GENERAL CONNECTIONS

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### Introduction.

For a manifold  $M$  with a general connection  $\Gamma$  we say a connected subset  $A$  is a black hole, if it has a neighborhood  $U$  such that if any one going on along a geodesic enters  $U$ , then he will be finally swallowed in  $A$ . The present author gave a way in [8] by which we can construct a general connection  $\Gamma$  for any Riemannian manifold  $(M, g)$  and any point  $p$  of  $M$  such that  $\Gamma$  has  $p$  as a black hole and has the same system of geodesics as the one of  $(M, g)$  outside of a neighborhood.

In the theory of general relativity, the Eddington-Finkelstein metric  $g$  is given by

$$(1) \quad d\tau^2 = -\left(1 - \frac{2m}{r}\right) dt^2 + 2dt dr + r^2(d\theta^2 + \sin^2\theta d\varphi^2),$$

where  $(r, \theta, \varphi)$  are the polar coordinates of the space  $R^3$  with the coordinates  $(x_1, x_2, x_3)$  as

$$r = \sqrt{\sum x_i^2}, \quad x_1 = r \sin \theta \cos \varphi, \quad x_2 = r \sin \theta \sin \varphi, \quad x_3 = r \cos \theta.$$

As is well known, the curve  $r=0$  in the space-time is a black hole as is mentioned above, even though the metric (1) loses the meaning along this curve, (1) is locally equivalent to the Schwarzschild metric

$$(2) \quad d\tau^2 = -\frac{r-2m}{r} dt^2 + \frac{r}{r-2m} dr^2 + r^2(d\theta^2 + \sin^2\theta d\varphi^2),$$

through the change of time  $t$  in (2) to  $t - r - \log|r - 2m|^{2m}$ . (2) loses its meaning where  $r=0$  and  $r=2m$  but (1) is everywhere regular except  $r=0$ .

Now, we denote the affine connection made by the Christoffel symbols from the space-time metric (1) by  $\Gamma_g$ . Taking a tensor field  $P$  of type  $(1, 1)$ , consider the general connection  $\Gamma = P\Gamma_g$ . Then, any geodesic of  $\Gamma_g$  is also a geodesic with respect to  $\Gamma$ . Conversely any geodesic of  $\Gamma$  is also a geodesic with respect to  $\Gamma_g$ , where  $P$  is an isomorphism on the tangent space of  $R \times (R^3 - \{0\})$ . We consider a problem: Taking  $P$  suitably, is it possible  $\Gamma = P\Gamma_g$  to extend smoothly

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over  $R \times R^3 = R^4$  with the canonical coordinates  $(x_0, x_1, x_2, x_3)$ ? Let  $\Gamma = (P_j^i, \Gamma_{jk}^i)$ , where  $P_j^i$  and  $\Gamma_{jk}^i$  are the components of  $\Gamma$  with respect to the coordinates

$$t = u_1, \quad r = u_2, \quad \theta = u_3, \quad \varphi = u_4.$$

We have the Christoffel symbols  $\left\{ \begin{smallmatrix} i \\ jk \end{smallmatrix} \right\}$  made by (1) as follows:

$$\left( \left\{ \begin{smallmatrix} 1 \\ jk \end{smallmatrix} \right\} \right) = \begin{pmatrix} m/u_2 u_2 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & -u_2 & 0 \\ 0 & 0 & 0 & -u_2 \sin^2 u_3 \end{pmatrix},$$

$$\left( \left\{ \begin{smallmatrix} 2 \\ jk \end{smallmatrix} \right\} \right) = \begin{pmatrix} mB/u_2 u_2 & -m/u_2 u_2 & 0 & 0 \\ -m/u_2 u_2 & 0 & 0 & 0 \\ 0 & 0 & 2m - u_2 & 0 \\ 0 & 0 & 0 & (2m - u_2) \sin^2 u_3 \end{pmatrix},$$

$$\left( \left\{ \begin{smallmatrix} 3 \\ jk \end{smallmatrix} \right\} \right) = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & 1/u_2 & 0 \\ 0 & 1/u_2 & 0 & 0 \\ 0 & 0 & 0 & -\sin u_3 \cos u_3 \end{pmatrix},$$

and

$$\left( \left\{ \begin{smallmatrix} 4 \\ jk \end{smallmatrix} \right\} \right) = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 1/u_2 \\ 0 & 0 & 0 & \cot u_3 \\ 0 & 1/u_2 & \cot u_3 & 0 \end{pmatrix},$$

where  $B = 1 - 2m/r$ . Since we have by definition  $\Gamma_{jk}^i = \Sigma P_h^i \left\{ \begin{smallmatrix} h \\ jk \end{smallmatrix} \right\}$  and from the condition that  $\Gamma$  is extended smoothly to  $R^4$ ,  $P_j^i$  must be of the forms as

$$(3) \quad P_j^i = F_j^i u_2 u_2 + 2m F_{2j}^i u_2, \quad P_2^i = F_2^i u_2 u_2, \quad P_3^i = F_3^i u_2, \quad P_4^i = F_4^i u_2 \sin u_3,$$

where,  $F_j^i$  are continuous near  $r=0$ . Hence we have

$$(4) \quad (\Gamma_{jk}^i) = \begin{pmatrix} m(F_1^i + F_2^i) - mF_2^i & 0 & 0 \\ -mF_2^i & 0 & F_3^i & F_4^i \sin u_3 \\ 0 & F_3^i & -(F_1^i + F_2^i)(u_2)^3 & F_4^i u_2 \cos u_3 \\ 0 & F_4^i \sin u_3 & F_4^i u_2 \cos u_3 & * \end{pmatrix}$$

where  $*$  is  $-(F_1^i + F_2^i)(u_2)^3 \sin^2 u_3 - F_3^i u_2 \sin u_3 \cos u_3$ . This expression tells us that if we compute the components of  $\Gamma$  in the canonical coordinates  $(x_0, x_1, x_2, x_3)$  of  $R \times R^3$ , it is possible to make it continuous but impossible to make it smooth. Since we have the expression of  $g$  in the coordinates  $(x_0, x_1, x_2, x_3)$  as

$$d\tau^2 = -\left(1 - \frac{2m}{r}\right) dt^2 + \frac{2}{r} \sum_{i=1}^3 x_i dt dx_i + \sum_{i=1}^3 dx_i dx_i - \left(\sum_{i=1}^3 \frac{x_i}{r} dx_i\right)^2$$

and the coefficients of the quadratic form  $rd\tau^2$  are continuous but some of them are not differentiable at the points where  $r=0$ . This fact may be the reason which implies the above situation on the general connection.

**§1. A certain space-time metric.**

In this section, we shall give a space-time metric on  $R \times (R^3 - \{0\})$  with the curve  $r=0$  as a black hole and make smooth general connections on  $R^4$  having the same system of geodesics with the one of this pseudo-Riemannian metric in  $R \times (R^3 - \{0\})$ .

First we consider a space-time metric  $g$  given by

$$(1.1) \quad d\sigma^2 = -\left(1 - \frac{4m^2}{r^2}\right)dt^2 + \frac{2}{r} dt dr + r^2(d\theta^2 + \sin^2\theta d\varphi^2)$$

in the same coordinates  $(t, r, \theta, \varphi)$  in Introduction and setting  $d\sigma^2 = \sum_{i,j} g_{ij} du_i du_j$ , where  $t=u_1, r=u_2, \theta=u_3$  and  $\varphi=u_4$ . Then we have the Christoffel symbols  $\left\{ \begin{smallmatrix} i \\ jk \end{smallmatrix} \right\}$  made by (1.1) as follows.

$$\left( \left\{ \begin{smallmatrix} 1 \\ jk \end{smallmatrix} \right\} \right) = \begin{pmatrix} 4m^2/u_2 u_2 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & -u_2 u_2 & 0 \\ 0 & 0 & 0 & -u_2 u_2 \sin^2 u_3 \end{pmatrix},$$

$$\left( \left\{ \begin{smallmatrix} 2 \\ jk \end{smallmatrix} \right\} \right) = \begin{pmatrix} 4m^2 B/u_2 & -4m^2/u_2 u_2 & 0 & 0 \\ -4m^2/u_2 u_2 & -1/u_2 & 0 & 0 \\ 0 & 0 & -B(u_2)^3 & 0 \\ 0 & 0 & 0 & -B(u_2)^3 \sin^2 u_3 \end{pmatrix},$$

$$\left( \left\{ \begin{smallmatrix} 3 \\ jk \end{smallmatrix} \right\} \right) = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & 1/u_2 & 0 \\ 0 & 1/u_2 & 0 & 0 \\ 0 & 0 & 0 & -\cos u_3 \sin u_3 \end{pmatrix},$$

$$\left( \left\{ \begin{smallmatrix} 4 \\ jk \end{smallmatrix} \right\} \right) = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 1/u_2 \\ 0 & 0 & 0 & \cot u_3 \\ 0 & 1/u_2 & \cot u_3 & 0 \end{pmatrix},$$

where  $B=1-4m^2/r^2$ . Hence the equation of a geodesic with respect to this space-time metric are

$$(1.2) \quad \left\{ \begin{array}{l} \frac{d^2 t}{dp^2} + \frac{4m^2}{r^2} \left( \frac{dt}{dp} \right)^2 - r^2 \left( \frac{d\theta}{dp} \right)^2 - r^2 \sin^2 \theta \left( \frac{d\varphi}{dp} \right)^2 = 0, \\ \frac{d^2 r}{dp^2} + \frac{4m^2 B}{r} \left( \frac{dt}{dp} \right)^2 - \frac{8m^2}{r^2} \frac{dt}{dp} \frac{dr}{dp} - \frac{1}{r} \left( \frac{dr}{dp} \right)^2 \\ \quad - Br^3 \left( \frac{d\theta}{dp} \right)^2 - Br^3 \sin^2 \theta \left( \frac{d\varphi}{dp} \right)^2 = 0, \\ \frac{d^2 \theta}{dp^2} + \frac{2}{r} \frac{dr}{dp} \frac{d\theta}{dp} - \cos \theta \sin \theta \left( \frac{d\varphi}{dp} \right)^2 = 0, \\ \frac{d^2 \varphi}{dp^2} + \frac{2}{r} \frac{dr}{dp} \frac{d\varphi}{dp} + 2 \cot \theta \frac{d\theta}{dp} \frac{d\varphi}{dp} = 0, \end{array} \right.$$

where  $p$  is the canonical parameter of the geodesic as

$$(1.3) \quad \frac{d\sigma^2}{dp^2} = - \left( 1 - \frac{4m^2}{r^2} \right) \left( \frac{dt}{dp} \right)^2 + \frac{2}{r} \frac{dt}{dp} \frac{dr}{dp} + r^2 \left\{ \left( \frac{d\theta}{dp} \right)^2 + \sin^2 \theta \left( \frac{d\varphi}{dp} \right)^2 \right\}$$

$$= c = \begin{cases} -1 \\ 0 \\ 1 \end{cases}$$

according to the sign of  $\sum_{i,j} g_{i,j} (du_i/dp)(du_j/dp)$ .

Next, consider a geodesic which pass through a given point  $q_0 = (t_0, r_0, \theta_0, \varphi_0)$  and  $(dq/dp)_0 = (\xi_0, \eta_0, \lambda_0, \mu_0)$ . Then we may put

$$\theta_0 = \frac{\pi}{2} \quad \text{and} \quad \lambda_0 = 0$$

without loss of generality, because the metric (1.1) is spherical symmetric with respect to  $(x_1, x_2, x_3)$ . Noticing that the third of (1.2) is satisfied with  $\theta \equiv \pi/2$ , we put  $\theta = \pi/2$  in (1.2) and (1.3), we obtain the following equations:

$$(1.2') \quad \left\{ \begin{array}{l} \frac{d^2 t}{dp^2} + \frac{4m^2}{r^2} \left( \frac{dt}{dp} \right)^2 - r^2 \left( \frac{d\varphi}{dp} \right)^2 = 0, \\ \frac{d^2 r}{dp^2} + \frac{4m^2 B}{r} \left( \frac{dt}{dp} \right)^2 - \frac{8m^2}{r^2} \frac{dt}{dp} \frac{dr}{dp} - \frac{1}{r} \left( \frac{dr}{dp} \right)^2 - Br^3 \left( \frac{d\varphi}{dp} \right)^2 = 0, \\ \frac{d^2 \varphi}{dp^2} + \frac{2}{r} \frac{dr}{dp} \frac{d\varphi}{dp} = 0 \end{array} \right.$$

and

$$(1.3') \quad -B \left( \frac{dt}{dp} \right)^2 + \frac{2}{r} \frac{dt}{dp} \frac{dr}{dp} + r^2 \left( \frac{d\varphi}{dp} \right)^2 = c,$$

From the third of (1.2') we see that  $r^2(d\varphi/dp)$  is constant along the geodesic and so we put this constant as

$$(1.4) \quad r^2 \frac{d\varphi}{dp} = r_0^2 \mu_0 = J.$$

Using this fact, the first and second of (1.2') become

$$\begin{cases} \frac{d^2t}{dp^2} = -\frac{4m^2}{r^2} \left(\frac{dt}{dp}\right)^2 + \frac{J^2}{r^2}, \\ \frac{d^2r}{dp^2} + \frac{4m^2B}{r} \left(\frac{dt}{dp}\right)^2 - \frac{8m^2}{r^2} \frac{dt}{dp} \frac{dr}{dp} - \frac{1}{r} \left(\frac{dr}{dp}\right)^2 - \frac{B}{r} J^2 = 0, \end{cases}$$

from which we obtain

$$\frac{d^2r}{dp^2} - \frac{8m^2}{r^2} \frac{dt}{dp} \frac{dr}{dp} - \frac{1}{r} \left(\frac{dr}{dp}\right)^2 - Br \frac{d^2t}{dp^2} = 0$$

by cancelling  $J$  and hence

$$\frac{d}{dp} \left( \frac{dr}{dp} - Br \frac{dt}{dp} \right) = \frac{1}{r} \frac{dr}{dp} \left( \frac{dr}{dp} - Br \frac{dt}{dp} \right).$$

Therefore we see that  $(1/r)(dr/dp - Br(dt/dp))$  is also constant along the geodesic and we put this constant as

$$(1.5) \quad \frac{1}{r} \left( \frac{dr}{dp} - Br \frac{dt}{dp} \right) = \frac{\eta_0}{r_0} - \left( 1 - \frac{4m^2}{r_0^2} \right) \xi_0 = A.$$

Finally using (1.4) and (1.5) for (1.3') we have

$$-\left( B \frac{dt}{dp} - \frac{1}{r} \frac{dr}{dp} \right)^2 + \frac{1}{r^2} \left( \frac{dr}{dp} \right)^2 + \frac{B}{r^2} J^2 = cB$$

and so

$$(1.6) \quad -A^2 + \frac{1}{r^2} \left( \frac{dr}{dp} \right)^2 = -B \left( \frac{J^2}{r^2} - c \right) \quad \text{or} \quad \left( \frac{d \log r}{dp} \right)^2 = A^2 - B \left( \frac{J^2}{r^2} - c \right).$$

From (1.6) we see the following fact. When  $c = -1$  or  $0$ , if  $r \leq 2m$  ( $B \leq 0$ ), which implies

$$\left| \frac{d \log r}{dp} \right| \geq |A|.$$

Let  $t_1$  be the moment such that the geodesic passes through the hypersurface  $r = 2m$  at the point  $(t_1, 2m, \pi/2, \varphi_1)$  then we have from (1.5)

$$A = \frac{1}{2m} \eta_1, \quad \eta_1 = \left. \frac{dr}{dp} \right|_{t=t_1}.$$

Therefore the geodesic enters the hypersurface  $r = 2m$  with  $\eta_1 < 0$ , then the decreasing ratio of  $\log r$  is greater than  $|A|$ .

**THEOREM 1.** *The space-time metric (1.1) has the curve  $r = 0$  in  $R \times R^3$  as a black hole for the system of visible geodesics, i. e.  $c = -1$  or  $0$  in (1.3).*

§ 2. Smooth general connections with the same system of geodesics of (1.1).

In the canonical coordinates  $(x_0, x_1, x_2, x_3)$  of  $R \times R^3$  with  $x_0=t$ , (1.1) can be represented as

$$(2.1) \quad d\sigma^2 = -\left(1 - \frac{4m^2}{r^2}\right) dt^2 + \frac{2}{r^2} \sum_{i=1}^3 x_i dt dx_i + \sum_{i=1}^3 dx_i dx_i - \frac{1}{r^2} \left(\sum_{i=1}^3 x_i dx_i\right)^2$$

and setting the right hand side of (2.1) as  $\sum_{\alpha, \beta=0}^3 g_{\alpha\beta} dx_\alpha dx_\beta$ , we have

$$(2.2) \quad g_{00} = -\left(1 - \frac{4m^2}{r^2}\right) = -B, \quad g_{0j} = g_{j0} = \frac{x_j}{r^2}, \quad g_{ij} = g_{ji} = \delta_{ij} - \frac{x_i x_j}{r^2},$$

from which  $(g^{\alpha\beta}) = (g_{\alpha\beta})^{-1}$  are given as

$$(2.3) \quad g^{00} = 0, \quad g^{0i} = g^{i0} = x_i, \quad g^{ij} = \delta_{ij} + \left(B - \frac{1}{r^2}\right) x_i x_j.$$

Making use of (2.2) and (2.3), the Christoffel symbols  $\left\{ \begin{smallmatrix} \alpha \\ \beta\gamma \end{smallmatrix} \right\}$  of (1.1) in the canonical coordinates  $(x_\alpha)$  are given by the formulas as follows

$$(2.4) \quad \left\{ \begin{smallmatrix} 0 \\ \beta\gamma \end{smallmatrix} \right\} = \frac{1}{r^2} \begin{pmatrix} 4m^2 & 0 \\ 0 & -(r^2 \delta_{ij} - x_i x_j) \end{pmatrix},$$

$$\left\{ \begin{smallmatrix} h \\ \beta\gamma \end{smallmatrix} \right\} = \frac{x_h}{r^4} \begin{pmatrix} 4m^2(r^2 - 4m^2) & -4m^2 x_j \\ -4m^2 x_i & -x_i x_j - (r^2 - 4m^2 - 1)(r^2 \delta_{ij} - x_i x_j) \end{pmatrix}.$$

Now, take a tensor field  $P$  of type (1,1) with local components  $P_\beta^\alpha$  and let  $\Gamma$  be the general connection  $P\Gamma_g$ , where  $\Gamma_g$  is the affine connection with the components  $\left\{ \begin{smallmatrix} \alpha \\ \beta\gamma \end{smallmatrix} \right\}$ . Since for  $\Gamma = (P_\beta^\alpha, \Gamma_{\beta\gamma}^\alpha)$  we have

$$(2.5) \quad \Gamma_{\beta\gamma}^\alpha = \sum_{\tau=0}^3 P_\tau^\alpha \left\{ \begin{smallmatrix} \tau \\ \beta\gamma \end{smallmatrix} \right\}$$

and so in order to be determined  $\Gamma$  so that it is smooth near  $r=0$  and has the same system of geodesics as the one of (2.1) in  $R \times (R^3 - \{0\})$ , it is necessary and sufficient to put

$$(2.6) \quad P_0^\alpha = F_0^\alpha r^2, \quad P_i^\alpha = F_i^\alpha r^4,$$

where  $F_\beta^\alpha$  are smooth near  $r=0$ , and

$$(2.7) \quad |F_\beta^\alpha| \neq 0 \quad \text{where } r \neq 0.$$

Then,  $\Gamma$  is regular where  $r \neq 0$ . Thus we obtain

$$(2.8) \quad (\Gamma_{\beta\gamma}^\alpha) = F_0^\alpha \begin{pmatrix} 4m^2 & 0 \\ 0 & -(r^2 \delta_{ij} - x_i x_j) \end{pmatrix}$$

$$+\sum_h F_h^\alpha x_h \begin{pmatrix} 4m^2(r^2-4m^2) & -4m^2 x_j \\ -4m^2 x_i & -x_i x_j - (r^2-4m^2-1)(r^2 \delta_{ij} - x_i x_j) \end{pmatrix}$$

There are many freedoms of the choice of  $P$  for the purpose mentioned above for (2.1). We request that

$$(2.9) \quad \partial_{\beta,0}^\alpha = 0,$$

where “,” denotes the covariant differentiation with respect to  $\Gamma$ . If we put for  $F$  the condition

$$(2.10) \quad \frac{\partial F_\beta^\alpha}{\partial t} = 0,$$

then we have for the covariant components of  $\Gamma$ :

$$A_{\beta\gamma}^\alpha := \Gamma_{\beta\gamma}^\alpha - \frac{\partial P_\beta^\alpha}{\partial x_\gamma}$$

the equalities

$$A_{\beta 0}^\alpha = \Gamma_{\beta 0}^\alpha,$$

and hence

$$\partial_{\beta,0}^\alpha = \sum_\rho \Gamma_{\rho 0}^\alpha P_{\beta^\rho}^\rho - \sum_\rho P_\rho^\alpha A_{\beta 0}^\rho = \sum_\rho \Gamma_{\rho 0}^\alpha P_{\beta^\rho}^\rho - \sum_\rho P_\rho^\alpha \Gamma_{\beta 0}^\rho = 0.$$

From (2.6) and (2.8), we have

$$(P_\beta^\alpha) = r^2(F_\beta^\alpha, r^2 F_j^\alpha), \quad (\Gamma_{\beta 0}^\alpha) = 4m^2(F_\beta^\alpha + (r^2 + 4m^2)V^\alpha, -V^\alpha x_j),$$

where  $V^\alpha := \sum_h F_h^\alpha x_h$ . Hence, setting  $W_\alpha = \sum F_h^\alpha x_h$ , and using (2.7), the condition  $\partial_{0,0}^\alpha = 0$  is equivalent to

$$(2.11) \quad \begin{cases} (r^2 - 4m^2)V^0 = 0, \\ r^2 \{F_j^0 + (r^2 - 4m^2)V^j\} = \{(r^2 - 4m^2)F_0^0 - W_0\} x_j \end{cases}$$

and the condition  $\partial_{j,0}^\alpha = 0$  is equivalent to

$$(2.12) \quad \begin{cases} -V^0 x_j = r^2 F_j^0, \\ -V^h x_j = \{(r^2 - 4m^2)F_j^0 - W_j\} x_h. \end{cases}$$

From the first of (2.11) and (2.12) we obtain

$$(2.13) \quad \begin{cases} F_j^0 = 0, \\ \sum_h F_h^i x_h = \lambda x_i, \quad \sum_h x_h F_j^h = \lambda x_j. \end{cases}$$

where  $\lambda$  is an auxiliary function. Then using (2.13) for the second of (2.11) we can put

$$(2.14) \quad F_0^i = \mu x_i,$$

where  $\mu$  is an auxiliary function, and hence we obtain

$$(2.15) \quad 2r^2\mu = (r^2 - 4m^2)(F_0^0 - \lambda r^2).$$

Thus, we see that

LEMMA 2.1. *Supposing that  $P$  does not depend on  $t$ , then  $\delta_{\beta,0}^\alpha = 0$  is equivalent to (2.13), (2.14) and (2.15).*

Now, considering (2.13), we take a special one such that

$$(2.16) \quad F_0^0 = r^2 F \quad \text{and} \quad F_j^i = \lambda \delta_j^i,$$

then from (2.15) we obtain

$$\lambda = -\frac{2\mu}{r^2 - 4m^2} + F$$

and so if we put

$$(2.17) \quad \mu = (r^2 - 4m^2)G,$$

then we obtain from the above equality

$$(2.18) \quad \lambda = F - 2G.$$

Thus, we obtain a special  $P = (P_\beta^\alpha)$  implying  $\delta_{\beta,0}^\alpha = 0$  given by

$$(2.19) \quad (P_\beta^\alpha) = r^2 \begin{pmatrix} r^2 F & 0 \\ (r^2 - 4m^2)G x_i & r^2(F - 2G)\delta_j^i \end{pmatrix},$$

where  $F$  and  $G$  are smooth,  $F \neq 0$  and  $F - 2G \neq 0$ .

THEOREM 2. *The general connection  $\Gamma = P\Gamma_g$  with  $P$  given by (2.19) is smooth on  $R \times R^3$ , has the same system of geodesics as the one of the space-time metric (2.1) where  $r \neq 0$ , and satisfies the conditions:*

$$\delta_{\beta,0}^\alpha = 0 \quad \text{and} \quad \tilde{g}_{\alpha\beta,0} = 0,$$

where  $\tilde{g}_{\alpha\beta} = r^2 g_{\alpha\beta}$ .

*Proof.* Except the last condition  $\tilde{g}_{\alpha\beta,0} = 0$ , the rest ones are evident from the above argument. In fact, we have

$$(\tilde{g}_{\alpha\beta}) = \begin{pmatrix} -(r^2 - 4m^2) & x_j \\ x_i & r^2 \delta_{ij} - x_i x_j \end{pmatrix}$$

and

$$\tilde{g}_{\alpha\beta,0} := \sum_{\rho,\sigma} P_\alpha^\rho P_\beta^\sigma \frac{\partial \tilde{g}_{\rho\sigma}}{\partial x_0} - \sum_{\rho,\sigma} \tilde{g}_{\rho\sigma} A_{\alpha 0}^\rho P_\beta^\sigma - \sum_{\rho,\sigma} \tilde{g}_{\rho\sigma} P_\alpha^\rho A_{\beta 0}^\sigma,$$

into which substituting (2.19) and

$$(A_{\beta 0}^{\alpha})=(\Gamma_{\beta 0}^{\alpha})=4m^2\begin{pmatrix} r^2F & 0 \\ (r^2-4m^2)(F-G)x_i & -(F-2G)x_ix_j \end{pmatrix}$$

we can easily obtain  $\tilde{g}_{\alpha\beta,0}=0$ . Q. E. D.

Finally, we give the components  $\Gamma_{\beta\gamma}^{\alpha}$  of  $\Gamma=PF_g$  in Theorem 2, which are

$$(2.20) \quad (\Gamma_{\beta\gamma}^{\alpha})=F_0^{\alpha}\begin{pmatrix} 4m^2 & 0 \\ 0 & -(r^2\delta_{ij}-x_ix_j) \end{pmatrix} + \sum_h F_h x_h \begin{pmatrix} 4m^2(r^2-4m^2) & -4m^2x_j \\ -4m^2x_i & -x_ix_j-(r^2-4m^2-1)(r^2\delta_{ij}-x_ix_j) \end{pmatrix}$$

where

$$F_0^0=r^2F, \quad F_0^i=(r^2-4m^2)Gx_i, \quad F_j^0=0, \quad F_j^i=(F-2G)\delta_j^i.$$

**§ 3. The curvature form for a special general connection in Theorem 2.**

In this section, we shall give the curvature form for the special general connection in Theorem 2 given by

$$(3.1) \quad F \equiv 1 \quad \text{and} \quad G \equiv 0.$$

Then we have

$$(3.2) \quad P_{\beta}^{\alpha}=r^4\delta_{\beta}^{\alpha}$$

and (2.20) becomes

$$(3.3) \quad (\Gamma_{\beta\gamma}^{\alpha})=r^2\begin{pmatrix} 4m^2 & 0 \\ 0 & -r^2\delta_{jh}+x_jx_h \end{pmatrix}, \\ (\Gamma_{\beta\gamma}^i)=x_i\begin{pmatrix} 4m^2(r^2-4m^2) & -4m^2x_h \\ -4m^2x_j & -x_jx_h+(r^2-4m^2-1)(x_jx_h-r^2\delta_{jh}) \end{pmatrix}.$$

The connection  $\Gamma=PF_g$  in Theorem 2 is smooth on  $R \times R^s$  and hence we can obtain the curvature along  $r=0$  by taking its limit from the outside of the curve. Where  $r \neq 0$ ,  $\Gamma$  is regular, i.e.  $\det(P_{\beta}^{\alpha}) \neq 0$ , therefore the curvature form  $\Omega_{\beta}^{\alpha}$  can be computed from the one of  $\Gamma_g$  by the formula ([2], § 7)

$$(3.4) \quad \Omega_{\beta}^{\alpha}=\sum P_{\gamma}^{\alpha}P_{\rho}^{\gamma}\prime\Omega_{\rho}^{\alpha}P_{\beta}^{\rho}+\sum P_{\rho}^{\alpha}\prime DP_{\rho}^{\alpha}\wedge\prime DP_{\beta}^{\rho},$$

where  $\prime D$  denote the covariant differentiation with respect to  $\Gamma_g$ . Hence (3.4) becomes in this case as

$$(3.5) \quad \Omega_{\beta}^{\alpha}=r^{12}\prime\Omega_{\beta}^{\alpha}.$$

Since we obtain the connection forms of  $\Gamma_g$  from (2.4) as follows:

$$(3.6) \quad \prime\omega_0^0=\frac{4m^2}{r^2}dt, \quad \prime\omega_j^0=-dx_j+x_jd \log r, \quad \prime\omega_0^i=\frac{4m^2}{r^2}x_i(Bdt-d \log r),$$

$$' \omega_j^i = - \left( B - \frac{1}{r^2} \right) x_i dx_j + x_i x_j \left\{ - \frac{4m^2}{r^4} dt + \left( B - \frac{2}{r^2} \right) d \log r \right\},$$

the curvature forms of  $\Gamma_g$

$$' \Omega_{\beta}^{\alpha} := d' \omega_{\beta}^{\alpha} + \sum_{\rho} ' \omega_{\rho}^{\alpha} \wedge ' \omega_{\beta}^{\rho}$$

are given by a little long computation as follows:

$$\begin{aligned} ' \Omega_0^0 &= \frac{8m^2}{r^2} dt \wedge d \log r, \\ ' \Omega_j^0 &= - \frac{4m^2}{r^2} dt \wedge dx_j + \frac{4m^2}{r^2} x_j dt \wedge d \log r + dx_j \wedge d \log r, \\ ' \Omega_i^0 &= - \frac{4m^2 B}{r^2} dt \wedge dx_i + \frac{12m^2}{r^2} B x_i dt \wedge d \log r - \frac{4m^2}{r^2} dx_i \wedge d \log r, \\ ' \Omega_j^i &= - \left( B - \frac{1}{r^2} \right) dx_i \wedge dx_j + \frac{4m^2 + 1}{r^2} x_i dx_j \wedge d \log r \\ &\quad - \frac{12m^2}{r^2} x_i x_j dt \wedge d \log r + \frac{4m^2}{r^4} x_j dt \wedge dx_i + \left( B - \frac{2}{r^2} \right) x_j dx_i \wedge d \log r. \end{aligned}$$

We see that the curvature forms vanish on the curve  $r=0$ .

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