TAIWANESE JOURNAL OF MATHEMATICS Vol. 15, No. 2, pp. 523-541, April 2011 This paper is available online at http://www.tjm.nsysu.edu.tw/

CLASSIFICATION THEOREMS FOR SPACE-LIKE SURFACES IN 4-DIMENSIONAL INDEFINITE SPACE FORMS WITH INDEX 2

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Abstract. Surfaces in 4D Riemannian space forms have been investigated extensively. In contrast, only few results are known for surfaces in 4D neutral indefinite space forms $R_2^4(c)$. Thus, in this paper we study space-like surfaces in $R_2^4(c)$ satisfying certain simple geometric properties. In particular, we classify space-like surfaces in \mathbb{E}_2^4 with constant mean and Gauss curvatures and null normal curvature. We also classify Wintgen ideal surfaces in $R_2^4(c)$ whose Gauss and normal curvatures satisfy $K^D = 2K$.

1. INTRODUCTION

Let \mathbb{E}_t^m denote the pseudo-Euclidean *m*-space equipped with pseudo-Euclidean metric of index *t* given by

(1.1)
$$g_t = -\sum_{i=1}^t dx_i^2 + \sum_{j=t+1}^n dx_j^2,$$

where (x_1, \ldots, x_m) is a rectangular coordinate system of \mathbb{E}_t^m . We put

(1.2)
$$S_s^k(c) = \left\{ x \in \mathbb{E}_s^{k+1} : \langle x, x \rangle = c^{-1} > 0 \right\},$$

(1.3)
$$H_s^k(c) = \left\{ x \in \mathbb{E}_{s+1}^{k+1} : \langle x, x \rangle = c^{-1} < 0 \right\},$$

where \langle , \rangle is the associated inner product. Then $S_s^k(c)$ and $H_s^k(c)$ are pseudo-Riemannian manifolds of constant curvature c and with index s, which are known as *pseudo-Riemannian k-sphere* and the *pseudo-hyperbolic k-space*, respectively. The pseudo-Riemannian manifolds $\mathbb{E}_s^k, S_s^k(c)$ and $H_s^k(-c)$ are called *indefinite space forms*, denoted by R_s^k .

Received September 4, 2009, accepted September 7, 2009.

Communicated by J. C. Yao.

²⁰⁰⁰ Mathematics Subject Classification: Primary 53C40; Secondary 53C50.

Key words and phrases: Gauss curvature, Normal curvature, Wintgen ideal surface, Space-like surface.

Surfaces in 4-dimensional Riemannian space forms have been investigated very extensively (see, for instance, [1, 2, 3]). In contrast, only few results are known for surfaces in 4-dimensional neutral indefinite space forms $R_2^4(c)$ of constant curvature c and index 2. Thus, we study in this paper space-like surfaces in $R_2^4(c)$ satisfying some simple geometric properties.

In Section 2 of this paper we provide basic definitions and formulas. In Section 3 we completely classify space-like surfaces in \mathbb{E}_2^4 with constant mean and Gauss curvatures and null normal curvature. In Section 4, we present a result of Sasaki and the precise expression of a minimal immersion ψ_B of the hyperbolic plane $H^2(-\frac{1}{3})$ of curvature $-\frac{1}{3}$ into the unit pseudo-hyperbolic 4-space $H_2^4(-1)$ discovered by the first author in [4]. It is known that the immersion ψ_B is a Wintgen ideal surfaces in $H_2^4(-1)$ whose Gauss and normal curvatures satisfy $K^D = 2K$. In the last section, we classify Wintgen ideal surfaces in $R_2^4(c)$ whose Gauss and normal curvatures satisfy $L^2(-\frac{1}{3}) \to H_2^4(-1)$.

2. Preliminaries

A vector v is called *space-like* (resp., *time-like*) if $\langle v, v \rangle > 0$ (resp., $\langle v, v \rangle < 0$). A surface M in a pseudo-Riemannian manifold is called *space-like* if each nonzero tangent vector is space-like.

Let $R_2^4(c)$ denote an indefinite space form of constant curvature c and with index 2. The curvature tensor \tilde{R} of $R_2^4(c)$ is given by

(2.1)
$$\tilde{R}(X,Y)Z = c\{\langle Y,Z \rangle X - \langle X,Z \rangle Y\}$$

for vectors X, Y, Z tangent to $R_2^4(c)$. Let $\psi : M \to R_2^4(c)$ be an isometric immersion of a space-like surface M into $R_2^4(c)$. Denote by ∇ and $\tilde{\nabla}$ the Levi-Civita connections on M and $R_2^4(c)$, respectively. For vector fields X, Y tangent to Mand ξ normal to M, the formulas of Gauss and Weingarten are given respectively by (cf. [1, 2, 10]):

(2.2)
$$\tilde{\nabla}_X Y = \nabla_X Y + h(X, Y),$$

(2.3)
$$\tilde{\nabla}_X \xi = -A_\xi X + D_X \xi,$$

where $\nabla_X Y$ and $A_{\xi} X$ are the tangential components and h(X, Y) and $D_X \xi$ the normal components of $\tilde{\nabla}_X Y$ and $\tilde{\nabla}_X \xi$, respectively.

The shape operator A and the second fundamental form h are related by

(2.4)
$$\langle h(X,Y),\xi\rangle = \langle A_{\xi}X,Y\rangle$$

The mean curvature vector H of M in $H_2^4(-1)$ is defined by $H = \frac{1}{2} \operatorname{trace} h$.

The equations of Gauss, Codazzi and Ricci are given respectively by

(2.5)
$$R(X,Y)Z = c\{\langle Y, Z \rangle X - \langle X, Z \rangle Y\} + A_{h(Y,Z)}X - A_{h(X,Z)}Y,$$

(2.6)
$$(\bar{\nabla}_X h)(Y, Z) = (\bar{\nabla}_Y h)(X, Z),$$

(2.7)
$$\langle R^D(X,Y)\xi,\eta\rangle = \langle [A_\xi,A_\eta]X,Y\rangle$$

for vector fields X, Y, Z tangent to M, and ξ, η normal to M, where $\overline{\nabla}h$ is defined by

$$(\bar{\nabla}_X h)(Y, Z) = D_X h(Y, Z) - h(\nabla_X Y, Z) - h(Y, \nabla_X Z),$$

and \mathbb{R}^D is the curvature tensor associated with the normal connection D, i.e.,

(2.8)
$$R^{D}(X,Y)\xi = D_{X}D_{Y}\xi - D_{Y}D_{X}\xi - D_{[X,Y]}\xi.$$

The normal curvature K^D is given by

(2.9)
$$K^D = \langle R^D(e_1, e_2)e_3, e_4 \rangle.$$

A surface M in $R_2^4(c)$ is called *parallel* (resp., *minimal*) if $\overline{\nabla}h = 0$ (resp., H = 0) holds identically. An immersion $\psi : M \to R_2^4(c)$ is called *full* if the image $\psi(M)$ does not lies in any totally geodesic submanifold of $R_2^4(c)$. A surface M in $R_2^4(c)$ is called *isotropic* if, at each point $p \in M$, |h(u, u)| is independent of the choice of the unit vector $u \in T_p M$.

For an immersion $\psi : M \to H_2^4(-1)$, let $L = \iota \circ \psi : M \to \mathbb{E}_3^5$ be the composition of ψ with the standard inclusion $\iota : H_2^4(-1) \to \mathbb{E}_3^5$ via (1.2). Since $H_2^4(-1)$ is totally umbilical with mean curvature one in \mathbb{E}_3^5 , we have

(2.10)
$$\nabla_X Y = \nabla_X Y + h(X, Y) + \langle X, Y \rangle L$$

for X, Y tangent to M, where h is the second fundamental form of ψ and $\hat{\nabla}$ denotes the Levi-Civita connection of \mathbb{E}_3^5 .

3. Surfaces with Null Normal Curvature in \mathbb{E}_2^4

Theorem 3.1. Let M be a space-like surface in the pseudo-Euclidean 4-space \mathbb{E}_2^4 . If M has constant mean and Gauss curvatures and null normal curvature, then M is congruent to an open part of one of the following six types of surfaces:

(1) A totally geodesic plane in \mathbb{E}_2^4 defined by (0, 0, x, y);

(2) a totally umbilical hyperbolic plane $H^2(-\frac{1}{a^2}) \subset \mathbb{E}^3_1 \subset \mathbb{E}^4_2$ given by

 $(0, a \cosh u, a \sinh u \cos v, a \sinh u \sin v),$

where *a* is a positive number;

(3) A flat surface in \mathbb{E}_2^4 defined by

$$\frac{1}{\sqrt{2m}}\Big(\cosh(\sqrt{2mx}),\cosh(\sqrt{2my}),\sinh(\sqrt{2mx}),\sinh(\sqrt{2my})\Big),$$

where m is a positive number;

(4) A flat surface in \mathbb{E}_2^4 defined by

$$\left(0, \frac{1}{a}\cosh(ax), \frac{1}{a}\sinh(ax), y\right),\$$

where a is a positive number;

(5) A flat surface in \mathbb{E}_2^4 defined by

$$\left(\frac{\cosh(\sqrt{2}x)}{\sqrt{2mr}}, \frac{\cosh(\sqrt{2}y)}{\sqrt{2m(2m-r)}}, \frac{\sinh(\sqrt{2}x)}{\sqrt{2mr}}, \frac{\sinh(\sqrt{2}y)}{\sqrt{2m(2m-r)}}\right),$$

where m and r are positive numbers satisfying 2m > r > 0;

(6) A surface of negative curvature $-b^2$ in \mathbb{E}_2^4 defined by

$$\left(\frac{1}{b}\cosh(bx)\cosh(by), \int_0^y \cosh(by)\sinh\left(\frac{4\sqrt{m^2-b^2}}{b}\tan^{-1}\left(\tanh\frac{by}{2}\right)\right)dy, \\ \frac{1}{b}\sinh(bx)\cosh(by), \int_0^y \cosh(by)\cosh\left(\frac{4\sqrt{m^2-b^2}}{b}\tan^{-1}\left(\tanh\frac{by}{2}\right)\right)dy \right)$$

where b and m are real numbers satisfying 0 < b < m.

Proof. Assume that $L: M \to \mathbb{E}_2^4$ is an isometric immersion of a space-like surface M into \mathbb{E}_2^4 . If M is totally geodesic in \mathbb{E}_2^4 , we obtain case (1). Thus, from now on, we assume that M is non-totally geodesic in \mathbb{E}_2^4 .

Let us choose an orthonormal tangent frame $\{e_1, e_2\}$ of the tangent bundle and an orthonormal normal frame $\{e_3, e_4\}$ of the normal bundle of M which satisfy

(3.1)
$$\langle e_1, e_1 \rangle = \langle e_2, e_2 \rangle = 1, \ \langle e_1, e_2 \rangle = 0,$$

(3.2)
$$\langle e_3, e_3 \rangle = \langle e_4, e_4 \rangle = -1, \ \langle e_3, e_4 \rangle = 0$$

We may also choose e_1, e_2 which diagonalize A_{e_3} so that the shape operator satisfies

(3.3)
$$A_{e_3} = \begin{pmatrix} \alpha & 0 \\ 0 & \mu \end{pmatrix}, \ A_{e_4} = \begin{pmatrix} \delta & \gamma \\ \gamma & -\delta \end{pmatrix}$$

for some functions $\alpha, \gamma, \delta, \mu$.

By definition, the *normal curvature* K^D of M is defined by

(3.4)
$$K^D = \langle [A_{e_3}, A_{e_4}]e_1, e_2 \rangle$$
.

For the orthonormal frame $\{e_1, e_2, e_3, e_4\}$, we put

(3.5)
$$\nabla_X e_1 = \omega_1^2(X) e_2, \quad D_X e_3 = \omega_3^4(X) e_4.$$

From (2.3), (3.2) and (3.3) we have

(3.6)
$$h(e_1, e_1) = -\alpha e_3 - \delta e_4, \ h(e_1, e_2) = -\gamma e_4, \ h(e_2, e_2) = -\mu e_3 + \delta e_4.$$

Thus, the mean curvature vector, the Gauss curvature and the normal curvature are given respectively by

(3.7)
$$H = -\frac{\alpha + \mu}{2} e_3, \ K = \gamma^2 + \delta^2 - \alpha \mu, \ K^D = \gamma(\mu - \alpha).$$

It follows from (3.5), (3.6) and the equation of Codazzi that

(3.8)
$$e_1\gamma - e_2\delta = \alpha\omega_3^4(e_2) - 2\gamma\omega_1^2(e_2) - 2\delta\omega_1^2(e_1).$$

(3.9)
$$e_2 \alpha = -\gamma \omega_3^4(e_1) + \delta \omega_3^4(e_2) + (\alpha - \mu) \omega_1^2(e_1),$$

(3.10)
$$e_2\gamma + e_1\delta = \mu\omega_3^4(e_1) - 2\delta\omega_1^2(e_2) + 2\gamma\omega_1^2(e_1),$$

(3.11)
$$e_1\mu = -\delta\omega_3^4(e_1) - \gamma\omega_3^4(e_2) + (\alpha - \mu)\omega_1^2(e_2).$$

Since M has null normal curvature, we may also assume that $\gamma = 0$. Thus, by the constancy of mean and Gauss curvatures, we obtain from (3.7) that

(3.12)
$$\mu = 2m - \alpha, \quad k = \delta^2 + \alpha^2 - 2m\alpha$$

for some constants k, m. Without loss of generality, we may assume $m \ge 0$.

Case (i). $\mu = \alpha$. In this case, $\mu = \alpha = m$ is a constant, which gives $A_{e_3} = mI$. Moreover, (3.12) gives

(3.13)
$$\delta^2 = m^2 + k \ge 0.$$

Case (i.1). $m^2 = -k$. From (3.13), we get $\delta = 0$. Hence, M is a totally umbilical surfaces in \mathbb{E}_2^4 . Such a surface has parallel second fundamental form. Therefore, after applying Proposition 4.3 of [5], we obtain case (2) of the theorem,

Case (i.2). $m^2 > -k$. Without loss of generality, we may put $\delta = \sqrt{m^2 + k}$, which is a nonzero constant. Thus, we find from (3.8)-(3.11) that $\omega_1^2 = \omega_3^4 = 0$. Hence, M must be flat. So, we have k = 0. Because $\omega_1^2 = 0$, we may choose coordinates $\{x, y\}$ such that $e_1 = \partial/\partial x, e_2 = \partial/\partial y$. The metric tensor is then given by $g = dx^2 + dy^2$. Moreover, we know that the second fundamental form satisfies

$$(3.14) \quad h(e_1, e_1) = -me_3 - me_4, \ h(e_1, e_2) = 0, \ h(e_2, e_2) = -me_3 + me_4.$$

Now, it follows from (2.1), (3.14) that the immersion $L: M \to \mathbb{E}_2^4$ satisfies

$$L_{xx} = -me_3 - me_4, \ L_{xy} = 0, \ L_{yy} = -me_3 + me_4,$$

(3.15) $\tilde{\nabla}_{\frac{\partial}{\partial x}}e_3 = -mL_x, \ \tilde{\nabla}_{\frac{\partial}{\partial y}}e_3 = -mL_y, \ \tilde{\nabla}_{\frac{\partial}{\partial x}}e_4 = -mL_x, \ \tilde{\nabla}_{\frac{\partial}{\partial y}}e_4 = mL_y.$

After solving this system and choosing suitable initial conditions, we get case (3).

Case (ii). $\mu \neq \alpha$. It follows from (3.12) that

(3.16)
$$\mu = 2m - \alpha, \ \delta = \sqrt{k + 2m\alpha - \alpha^2}.$$

Case (ii.1). $\delta = 0$. In this case, we have $k = \alpha^2 - 2m\alpha$ which is constant. Hence, α is also a constant. Thus, we derive from (3.8)-(3.11) that

(3.17)
$$\omega_1^2 = 0, \ \alpha \omega_3^4(e_2) = \mu \omega_3^4(e_1) = 0.$$

Therefore, M is flat and $\alpha \mu = 0$. Since M is non-totally geodesic, without loss of generality we may assume that $\alpha \neq 0$ and $\mu = 0$. Since $\omega_1^2 = 0$, we may choose coordinates $\{x, y\}$ such that $e_1 = \partial/\partial x$, $e_2 = \partial/\partial y$. So, we obtain

(3.18)
$$h(e_1, e_1) = -\alpha e_3, \ h(e_1, e_2) = h(e_2, e_2) = 0.$$

It follows from (3.17) and (3.18) that immersion $L: M \to \mathbb{E}_2^4$ satisfies

(3.19)
$$L_{xx} = -\alpha e_3, \quad L_{xy} = L_{yy} = 0,$$
$$\tilde{\nabla}_{\frac{\partial}{\partial x}} e_3 = -\alpha L_x, \quad \tilde{\nabla}_{\frac{\partial}{\partial y}} e_3 = 0.$$

After solving this system and choosing suitable initial conditions, we get case (4).

Case (ii.2). $\delta \neq 0$. We have

(3.20)
$$\mu = 2m - \alpha, \ \gamma = 0, \ \delta = \sqrt{k + 2m\alpha - \alpha^2} \neq 0.$$

Case (ii.2.1). $m\alpha = -k$. In this case, α and δ are constant. Moreover, we have

(3.21)
$$\alpha = -\frac{k}{m}, \ \delta = \frac{\sqrt{-k(k+m^2)}}{m}, \ \mu = 2m + \frac{k}{m}, \ \gamma = 0$$

Because δ is a real nonzero number, we must have $-m^2 < k < 0$. Thus, we may put $k = -b^2$ with 0 < b < m. Substituting (3.21) into (3.8)-(3.11) yields

(3.22)
$$\omega_1^2(e_2) = \omega_3^4(e_1) = 0, \ \ \omega_3^4(e_2) = \frac{2\sqrt{m^2 - b^2}}{b}\omega_1^2(e_1)$$

Thus, if f is a function satisfying $e_2(\ln f) = \omega_1^2(e_1)$, then we get $[fe_1, e_2] = 0$, which implies that there exist coordinates $\{x, y\}$ such that

(3.23)
$$\frac{\partial}{\partial x} = fe_1, \quad \frac{\partial}{\partial y} = e_2.$$

Therefore, the metric tensor is given by

(3.24)
$$g = f^2 dx^2 + dy^2.$$

Consequently, the Levi-Civita connection satisfies

(3.25)
$$\nabla_{\frac{\partial}{\partial x}\frac{\partial}{\partial x}} = \frac{f_x}{f}\frac{\partial}{\partial x} - ff_y\frac{\partial}{\partial y}, \quad \nabla_{\frac{\partial}{\partial x}\frac{\partial}{\partial y}} = \frac{f_y}{f}\frac{\partial}{\partial x}, \quad \nabla_{\frac{\partial}{\partial y}\frac{\partial}{\partial y}} = 0.$$

From (3.22), (3.23) and (3.25), we derive that

(3.26)
$$\omega_1^2(e_1) = -\frac{f_y}{f}, \ \omega_1^2(e_2) = \omega_3^4(e_1) = 0, \ \omega_3^4(e_2) = -\frac{2f_y\sqrt{m^2 - b^2}}{bf}.$$

Moreover, it follow from (3.24) and $K = -b^2$ that f satisfies

$$(3.27) f_{yy} = b^2 f_z$$

By solving (3.27) we obtain $f = u(x) \cosh(by+v(x))$ for some functions u(x), v(x). After replacing x by an anti-derivative of u(x), we find from (3.24) and (3.25) that

(3.28)
$$g = \cosh^2(by + v(x))dx^2 + dy^2,$$

(3.29)
$$\nabla_{\frac{\partial}{\partial x}} \frac{\partial}{\partial x} = v' \tanh(by+v) \frac{\partial}{\partial x} - \frac{b}{2} \sinh(2by+2v) \frac{\partial}{\partial y},$$
$$\nabla_{\frac{\partial}{\partial x}} \frac{\partial}{\partial y} = b \tanh(by+v) \frac{\partial}{\partial x}, \quad \nabla_{\frac{\partial}{\partial y}} \frac{\partial}{\partial y} = 0.$$

Also, it follows from (3.6), (3.21), and (3.28) that

(3.30)
$$h\left(\frac{\partial}{\partial x}, \frac{\partial}{\partial x}\right) = -\frac{\cosh^2(by+v)}{m} \{b^2 e_3 + b\sqrt{m^2 - b^2} e_4\},$$
$$h\left(\frac{\partial}{\partial x}, \frac{\partial}{\partial y}\right) = 0,$$
$$h\left(\frac{\partial}{\partial y}, \frac{\partial}{\partial y}\right) = \frac{(b^2 - 2m^2)e_3 + b\sqrt{m^2 - b^2} e_4}{m}.$$

Therefore, the immersion $L: M \to \mathbb{E}_2^4$ satisfies

$$L_{xx} = v' \tan(by+v)L_x - \frac{b}{2}\sinh(2by+2v)L_y - \frac{\cosh^2(by+v)}{m} \{b^2e_3 + b\sqrt{m^2 - b^2}e_4\},$$

$$L_{xy} = b \tanh(by+v)L_x,$$

$$L_{yy} = \frac{(b^2 - 2m^2)e_3 + b\sqrt{m^2 - b^2}e_4}{m},$$

$$\tilde{\nabla}_{\frac{\partial}{\partial x}}e_3 = -\frac{b^2}{m}L_x, \quad \tilde{\nabla}_{\frac{\partial}{\partial x}}e_4 = -\frac{b\sqrt{m^2 - b^2}}{m}L_x,$$

$$\tilde{\nabla}_{\frac{\partial}{\partial y}}e_3 = \frac{b^2 - 2m^2}{m}L_y - 2\sqrt{m^2 - b^2}\tanh(by+v)e_4,$$

$$\tilde{\nabla}_{\frac{\partial}{\partial y}}e_4 = \frac{b\sqrt{m^2 - b^2}}{m}L_y + 2\sqrt{m^2 - b^2}\tanh(by+v)e_3.$$

The compatibility condition of (3.31) is given by v'(x) = 0. Then, after applying a suitable translation in y, we may put v = 0. Therefore, system (3.31) reduces to

$$L_{xx} = -\frac{b}{2}\sinh(2by)L_y - \frac{\cosh^2(by)}{m} \{b^2e_3 + b\sqrt{m^2 - b^2}e_4\},$$

$$L_{xy} = b\tanh(by)L_x,$$

$$L_{yy} = \frac{(b^2 - 2m^2)e_3 + b\sqrt{m^2 - b^2}e_4}{m},$$
(3.32)
$$\tilde{\nabla}_{\frac{\partial}{\partial x}}e_3 = -\frac{b^2}{m}L_x, \quad \tilde{\nabla}_{\frac{\partial}{\partial x}}e_4 = -\frac{b\sqrt{m^2 - b^2}}{m}L_x,$$

$$\tilde{\nabla}_{\frac{\partial}{\partial y}}e_3 = \frac{b^2 - 2m^2}{m}L_y - 2\sqrt{m^2 - b^2}\tanh(by)e_4,$$

$$\tilde{\nabla}_{\frac{\partial}{\partial y}}e_4 = \frac{b\sqrt{m^2 - b^2}}{m}L_y + 2\sqrt{m^2 - b^2}\tanh(by)e_3.$$

Solving the second equation in (3.32) gives

$$(3.33) L = A(x)\cosh by + B(y)$$

for some vector-valued functions A(x), B(y). Substituting this into the first, third and fourth equations in (3.32) gives $A'''(x) = b^2 A'(x)$. Thus, we get

(3.34)
$$A(x) = c_5 + c_1 \cosh(bx) + c_2 \sinh(bx)$$

for some vectors c_5 , c_1 , c_2 . Combining this with (3.33) gives

(3.35)
$$L = (c_5 + c_1 \cosh(bx) + c_2 \sinh(bx)) \cosh by + B(y).$$

By substituting (3.35) into the first, third and fifth equations in (3.32), we find

(3.36)
$$\cosh^2(by)B''' - \frac{b}{2}\sinh(2by)B'' + (3b^2 - 4m^2)B' \\ = c_5b(3b^2 - 4m^2)\sinh(by).$$

A direct computation shows that $B_p = -c_5 \cosh(by)$ is a particular solution of (3.36). Thus, it follows from (3.35) and (3.36) that

(3.37)
$$L = (c_1 \cosh(bx) + c_2 \sinh(bx)) \cosh by + C(y),$$

where C(y) satisfies the homogeneous differential equation:

(3.38)
$$\cosh^2(by)C'''(y) - \frac{b}{2}\sinh(2by)C''(y) + (3b^2 - 4m^2)C'(y) = 0$$

After solving this differential equation, we have

(3.39)
$$C(y) = c_3 \int_0^y \cosh(by) \cosh\left(\frac{4\sqrt{m^2 - b^2}}{b} \tan^{-1}\left(\tanh\frac{by}{2}\right)\right) dy + c_4 \int_0^y \cosh(by) \sinh\left(\frac{4\sqrt{m^2 - b^2}}{b} \tan^{-1}\left(\tanh\frac{by}{2}\right)\right) dy + c_0$$

for some vectors $c_3, c_4, c_5 \in \mathbb{E}_2^4$. Combining this with (3.37) yields

$$L = c_0 + (c_1 \cosh(bx) + c_2 \sinh(bx)) \cosh by$$

+ $c_3 \int_0^y \cosh(by) \cosh\left(\frac{4\sqrt{m^2 - b^2}}{b} \tan^{-1}\left(\tanh\frac{by}{2}\right)\right) dy$
+ $c_4 \int_0^y \cosh(by) \sinh\left(\frac{4\sqrt{m^2 - b^2}}{b} \tan^{-1}\left(\tanh\frac{by}{2}\right)\right) dy.$

Therefore, after choosing suitable initial conditions, we obtain case (6).

Case (ii.2.2). $m\alpha \neq -k$. By substituting (3.20) into (3.8)-(3.11) we obtain

(3.40)
$$\omega_3^4(e_1) = \frac{2(k+m^2)\omega_1^2(e_2)}{m\sqrt{k+2m\alpha-\alpha^2}}, \ \ \omega_3^4(e_2) = \frac{2(k+m^2)\omega_1^2(e_1)}{m\sqrt{k+2m\alpha-\alpha^2}}.$$

(3.41)
$$\omega_1^2(e_1) = e_2(\ln\sqrt{k+m\alpha}), \ \omega_2^1(e_2) = e_1(\ln\sqrt{k+2m^2-m\alpha}).$$

It follows from (3.41) that $[e_1/\sqrt{k+m\alpha}, e_2/\sqrt{k+2m^2-m\alpha}] = 0$. Thus, there exist coordinates $\{x, y\}$ such that

(3.42)
$$\frac{\partial}{\partial x} = \frac{e_1}{\sqrt{k+m\alpha}}, \quad \frac{\partial}{\partial y} = \frac{e_2}{\sqrt{k+2m^2-m\alpha}}.$$

Hence, the metric tensor is given by

(3.43)
$$g = \frac{dx^2}{k+m\alpha} + \frac{dy^2}{k+2m^2-m\alpha}.$$

From (3.43) we have

(3.44)
$$\nabla_{\frac{\partial}{\partial x}} \frac{\partial}{\partial x} = \frac{-m\alpha_x}{2(k+m\alpha)} \frac{\partial}{\partial x} + \frac{m(k+2m^2-m\alpha)\alpha_y}{2(k+m\alpha)^2} \frac{\partial}{\partial y},$$
$$\nabla_{\frac{\partial}{\partial y}} \frac{\partial}{\partial y} = \frac{-m\alpha_y}{2(k+m\alpha)} \frac{\partial}{\partial x} + \frac{m\alpha_x}{2(k+2m^2-m\alpha)} \frac{\partial}{\partial y},$$
$$\nabla_{\frac{\partial}{\partial y}} \frac{\partial}{\partial y} = \frac{-m(k+m\alpha)\alpha_x}{2(k+2m^2-m\alpha)^2} \frac{\partial}{\partial x} + \frac{m\alpha_y}{2(k+2m^2-m\alpha)} \frac{\partial}{\partial y}.$$

It follows from (3.6), (3.20) and (3.23) that the second fundamental form satisfies

$$(3.45) \qquad h\left(\frac{\partial}{\partial x}, \frac{\partial}{\partial x}\right) = \frac{-\alpha e_3 - \sqrt{k + 2m\alpha - \alpha^2}e_4}{k + m\alpha}, \quad h\left(\frac{\partial}{\partial x}, \frac{\partial}{\partial y}\right) = 0,$$
$$h\left(\frac{\partial}{\partial y}, \frac{\partial}{\partial y}\right) = \frac{(\alpha - 2m)e_3 + \sqrt{k + 2m\alpha - \alpha^2}e_4}{k + 2m^2 - m\alpha}.$$

By applying (2.1), (3.24), (3.25) and (3.26) we obtain

$$\begin{split} L_{xx} &= \frac{-m\alpha_x L_x}{2(k+m\alpha)} + \frac{m(k+2m^2-m\alpha)\alpha_y L_y}{2(k+m\alpha)^2} - \frac{\alpha e_3 + \sqrt{k+2m\alpha - \alpha^2} e_4}{k+m\alpha}, \\ L_{xy} &= \frac{-m\alpha_y L_x}{2(k+m\alpha)} + \frac{m\alpha_x L_y}{2(k+2m^2-m\alpha)}, \\ L_{yy} &= \frac{-m(k+m\alpha)\alpha_x L_x}{2(k+2m^2-m\alpha)^2} + \frac{m\alpha_y L_y}{2(k+2m^2-m\alpha)} \\ &+ \frac{(\alpha-2m)e_3 + \sqrt{k+2m\alpha - \alpha^2} e_4}{k+2m^2-m\alpha}, \end{split}$$

$$(3.46) \quad \tilde{\nabla}_{\frac{\partial}{\partial x}} e_3 &= -\alpha L_x + \frac{(k+m^2)\alpha_x}{(k+2m^2-m\alpha)\sqrt{k+2m\alpha - \alpha^2}} e_4, \\ \tilde{\nabla}_{\frac{\partial}{\partial y}} e_3 &= (\alpha-2m)L_y + \frac{(k+m^2)\alpha_y}{(k+m\alpha)\sqrt{k+2m\alpha - \alpha^2}} e_4, \\ \tilde{\nabla}_{\frac{\partial}{\partial y}} e_4 &= -\sqrt{k+2m\alpha - \alpha^2}L_x - \frac{(k+m^2)\alpha_x}{(k+2m^2-m\alpha)\sqrt{k+2m\alpha - \alpha^2}} e_3, \\ \tilde{\nabla}_{\frac{\partial}{\partial y}} e_4 &= \sqrt{k+2m\alpha - \alpha^2}L_y - \frac{(k+m^2)\alpha_y}{(k+m\alpha)\sqrt{k+2m\alpha - \alpha^2}} e_3. \end{split}$$

After applying (3.46) and a long computation, we find from $\langle L_{xxy}, L_y \rangle = \langle L_{xyx}, L_y \rangle$ and from $\langle L_{xyy}, L_x \rangle = \langle L_{yyx}, L_x \rangle$ that

(3.47)
$$\alpha_y\{(k+m\alpha)\alpha_x+(k+2m^2-m\alpha)\alpha_y\}=0.$$

Hence, we have either

(1) $\alpha_y = 0$ or (2) $(k + m\alpha)\alpha_x + (k + 2m^2 - m\alpha)\alpha_y = 0.$

Case (ii.2.2.a). $\alpha_y = 0$. In this case, system (3.46) reduces to

$$L_{xx} = \frac{-m\alpha_x L_x}{2(k+m\alpha)} - \frac{\alpha e_3 + \sqrt{k+2m\alpha - \alpha^2} e_4}{k+m\alpha},$$

$$L_{xy} = \frac{m\alpha_x L_y}{2(k+2m^2 - m\alpha)},$$

$$L_{yy} = \frac{-m(k+m\alpha)\alpha_x L_x}{2(k+2m^2 - m\alpha)^2} + \frac{(\alpha - 2m)e_3 + \sqrt{k+2m\alpha - \alpha^2} e_4}{k+2m^2 - m\alpha},$$
(3.48) $\tilde{\nabla}_{\frac{\partial}{\partial x}} e_3 = -\alpha L_x + \frac{(k+m^2)\alpha_x}{(k+2m^2 - m\alpha)\sqrt{k+2m\alpha - \alpha^2}} e_4,$
 $\tilde{\nabla}_{\frac{\partial}{\partial y}} e_3 = (\alpha - 2m)L_y,$
 $\tilde{\nabla}_{\frac{\partial}{\partial y}} e_4 = -\sqrt{k+2m\alpha - \alpha^2} L_x - \frac{(k+m^2)\alpha_x}{(k+2m^2 - m\alpha)\sqrt{k+2m\alpha - \alpha^2}} e_3,$
 $\tilde{\nabla}_{\frac{\partial}{\partial y}} e_4 = \sqrt{k+2m\alpha - \alpha^2} L_y.$

Now, after applying (3.48), $\langle L_{xxy}, L_y \rangle = \langle L_{xyx}, L_y \rangle$ and $\langle L_{xyy}, L_y \rangle = \langle L_{yyx}, L_y \rangle$, we obtain that

(3.49)
$$a_{xx} = -\frac{2k(k+2m^2-m\alpha)^2 - m^2(2k+m^2+m\alpha)\alpha_x^2}{m(k+2m^2-m\alpha)(k+m\alpha)}$$

(3.50)
$$\alpha_x^2 = \frac{2k(k+2m^2-m\alpha)^2}{m^2(k+m^2)}$$

Next, by differentiating (3.50) and by applying (3.49), we find $\alpha_x = 0$. Thus, α is a constant, say $\alpha = r$. Because $\delta \neq 0$, (3.50) gives k = 0. Therefore, system (3.48) becomes

$$L_{xx} = -\frac{re_3 + \sqrt{2mr - r^2 e_4}}{mr},$$

$$L_{xy} = 0,$$

$$L_{yy} = \frac{(r - 2m)e_3 + \sqrt{2mr - r^2 e_4}}{2m^2 - mr},$$

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$$\begin{split} \tilde{\nabla}_{\frac{\partial}{\partial x}} e_3 &= -\alpha L_x, \ \tilde{\nabla}_{\frac{\partial}{\partial y}} e_3 = (\alpha - 2m) L_y, \\ \tilde{\nabla}_{\frac{\partial}{\partial x}} e_4 &= -\sqrt{2mr - r^2} L_x, \ \tilde{\nabla}_{\frac{\partial}{\partial y}} e_4 = \sqrt{2mr - r^2} L_y \end{split}$$

After solving this system and choosing suitable initial conditions, we have case (5) of the theorem.

Case (ii.2.2.b). $(k + m\alpha)\alpha_x + (k + 2m^2 - m\alpha)\alpha_y = 0$. In this case, we have

(3.51)
$$\alpha_y = \frac{(k+m\alpha)\alpha_x}{m\alpha - k - 2m^2}.$$

Thus, system (3.46) becomes

$$\begin{split} L_{xx} &= -\frac{m\alpha_x(L_x + L_y)}{2(k + m\alpha)} - \frac{\alpha e_3 + \sqrt{k + 2m\alpha - \alpha^2} e_4}{k + m\alpha}, \\ L_{xy} &= \frac{m\alpha_x(L_x + L_y)}{2(k + 2m^2 - m\alpha)}, \\ L_{yy} &= -\frac{m(k + m\alpha)\alpha_x(L_x + L_y)}{2(k + 2m^2 - m\alpha)^2} + \frac{(\alpha - 2m)e_3 + \sqrt{k + 2m\alpha - \alpha^2}e_4}{k + 2m^2 - m\alpha}, \\ (3.52) \quad \tilde{\nabla}_{\frac{\partial}{\partial x}} e_3 &= -\alpha L_x + \frac{(k + m^2)\alpha_x e_4}{(k + 2m^2 - m\alpha)\sqrt{k + 2m\alpha - \alpha^2}}, \\ \tilde{\nabla}_{\frac{\partial}{\partial y}} e_3 &= (\alpha - 2m)L_y - \frac{(k + m^2)\alpha_x e_4}{(k + 2m^2 - m\alpha)\sqrt{k + 2m\alpha - \alpha^2}}, \\ \tilde{\nabla}_{\frac{\partial}{\partial x}} e_4 &= -\sqrt{k + 2m\alpha - \alpha^2}L_x - \frac{(k + m^2)\alpha_x e_3}{(k + 2m^2 - m\alpha)\sqrt{k + 2m\alpha - \alpha^2}}, \\ \tilde{\nabla}_{\frac{\partial}{\partial y}} e_4 &= \sqrt{k + 2m\alpha - \alpha^2}L_y + \frac{(k + m^2)\alpha_x e_4}{(k + 2m^2 - m\alpha)\sqrt{k + 2m\alpha - \alpha^2}}. \end{split}$$

Now, from $L_{xxy} = L_{xyx}$ we find $(k+m\alpha)\alpha_x = 0$. Also, we find from $L_{xyy} = L_{yyx}$ that k=0. Thus, α is a constant and k=0. Hence, this case reduces to (ii.2.a).

4. SPACELIKE MINIMAL SURFACES WITH CONSTANT GAUSS CURVATURE

From the equation of Gauss, we have

Lemma 4.1. Let M be a space-like minimal surface in $R_2^4(c)$. Then $K \ge c$. In particular, if K = c holds identically, then M is totally geodesic.

For space-like *minimal surfaces* in $R_2^4(c)$, Theorem 1 of [12] implies that M has constant Gauss curvature if and only if it has constant normal curvature.

We recall the following result of Sasaki from [12].

Theorem 4.2. Let M be a space-like minimal surface in $R_2^4(c)$. If M has constant Gauss curvature, then either

(1) K = c and M is a totally geodesic surface in $R_2^4(c)$;

(2) c < 0, K = 0 and M is congruent to an open part of the minimal surface defined by $\frac{1}{\sqrt{2}} (\cosh u, \cosh v, 0, \sinh u, \sinh v)$, or

(3) c < 0, K = c/3 and M is isotropic.

Let \mathbf{R}^2 be a plane with coordinates s, t. Consider a map $\mathcal{B} : \mathbf{R}^2 \to \mathbb{E}_3^5$ given by

(4.1)
$$\mathcal{B}(s,t) = \left(\sinh\left(\frac{2s}{\sqrt{3}}\right) - \frac{t^2}{3} - \left(\frac{7}{8} + \frac{t^4}{18}\right)e^{\frac{2s}{\sqrt{3}}}, t + \left(\frac{t^3}{3} - \frac{t}{4}\right)e^{\frac{2s}{\sqrt{3}}}, \\ \frac{1}{2} + \frac{t^2}{2}e^{\frac{2s}{\sqrt{3}}}, t + \left(\frac{t^3}{3} + \frac{t}{4}\right)e^{\frac{2s}{\sqrt{3}}}, \sinh\left(\frac{2s}{\sqrt{3}}\right) - \frac{t^2}{3} - \left(\frac{1}{8} + \frac{t^4}{18}\right)e^{\frac{2s}{\sqrt{3}}}\right)$$

The first author proved in [4] that \mathcal{B} defines a full isometric parallel immersion

(4.2)
$$\psi_{\mathcal{B}}: H^2(-\frac{1}{3}) \to H^4_2(-1)$$

or

of the hyperbolic plane $H^2(-\frac{1}{3})$ of curvature $-\frac{1}{3}$ into $H_2^4(-1)$. The following result was also obtained in [4].

Theorem 4.3. Let $\psi : M \to H_2^4(-1)$ be a parallel full immersion of a spacelike surface M into $H_2^4(-1)$. Then M is minimal in $H_2^4(-1)$ if and only if M is congruent to an open part of the surface defined by

$$\left(\sinh\left(\frac{2s}{\sqrt{3}}\right) - \frac{t^2}{3} - \left(\frac{7}{8} + \frac{t^4}{18}\right)e^{\frac{2s}{\sqrt{3}}}, t + \left(\frac{t^3}{3} - \frac{t}{4}\right)e^{\frac{2s}{\sqrt{3}}}, \\ \frac{1}{2} + \frac{t^2}{2}e^{\frac{2s}{\sqrt{3}}}, t + \left(\frac{t^3}{3} + \frac{t}{4}\right)e^{\frac{2s}{\sqrt{3}}}, \sinh\left(\frac{2s}{\sqrt{3}}\right) - \frac{t^2}{3} - \left(\frac{1}{8} + \frac{t^4}{18}\right)e^{\frac{2s}{\sqrt{3}}}\right).$$

Combining Theorem 4.2 and Theorem 4.3, we obtain the following.

Theorem 4.4. Let M be a non-totally geodesic space-like minimal surface in $H_2^4(-1)$. If M has constant Gauss curvature K, then either

(1) K = 0 and M is congruent to an open part of the surface defined by

$$\frac{1}{\sqrt{2}}\left(\cosh u, \cosh v, 0, \sinh u, \sinh v\right),\,$$

(2) $K = -\frac{1}{3}$ and M is is congruent to an open part of the surface defined by

$$\left(\sinh\left(\frac{2s}{\sqrt{3}}\right) - \frac{t^2}{3} - \left(\frac{7}{8} + \frac{t^4}{18}\right)e^{\frac{2s}{\sqrt{3}}}, t + \left(\frac{t^3}{3} - \frac{t}{4}\right)e^{\frac{2s}{\sqrt{3}}}, \\ \frac{1}{2} + \frac{t^2}{2}e^{\frac{2s}{\sqrt{3}}}, t + \left(\frac{t^3}{3} + \frac{t}{4}\right)e^{\frac{2s}{\sqrt{3}}}, \sinh\left(\frac{2s}{\sqrt{3}}\right) - \frac{t^2}{3} - \left(\frac{1}{8} + \frac{t^4}{18}\right)e^{\frac{2s}{\sqrt{3}}}\right).$$

5. WINTGEN IDEAL SURFACES SATISFYING $K^D = -2K$

In 1979, P. Wintgen [13] proved a basic relationship between Gauss curvature K, normal curvature K^D , and mean curvature vector H of a surface M in a Euclidean 4-space \mathbb{E}^4 ; namely,

(5.1)
$$K + |K^D| \le \langle H, H \rangle,$$

with the equality holding if and only if the curvature ellipse is a circle.

The following Wintgen type inequality for space-like surfaces in $R_2^4(c)$ can be found in [7].

Theorem 5.1. Let M be a space-like surface in a 4-dimensional indefinite space form $R_2^4(c)$ of constant sectional curvature c and index two. Then we have

(5.2)
$$K + K^D \ge \langle H, H \rangle + c$$

at every point. Moreover, the equality sign of (5.2) holds at a point $p \in M$ if and only if, with respect to some suitable orthonormal frame $\{e_1, e_2, e_3, e_4\}$, the shape operator at p satisfies

(5.3)
$$A_{e_3} = \begin{pmatrix} \mu + 2\gamma & 0 \\ 0 & \mu \end{pmatrix}, \ A_{e_4} = \begin{pmatrix} 0 & \gamma \\ \gamma & 0 \end{pmatrix}.$$

Following [6, 9, 11], we call a surface in $R_2^4(c)$ Wintgen ideal if it satisfies the equality case of (5.2) identically. Wintgen ideal surfaces in \mathbb{E}_2^4 satisfying $|K| = |K^D|$ are classified by the first author in [7] (see [6] for the classification of Wintgen ideal surfaces in \mathbb{E}^4 satisfying $|K| = |K^D|$).

We need the following existence result.

Theorem 5.2. Let c be a real number and γ with $3\gamma^2 > -c$ be a positive solution of the second order elliptic differential equation

(5.4)
$$\frac{\partial}{\partial x} \left(\frac{(3\gamma\sqrt{c+3\gamma^2}-c)(6\gamma+2\sqrt{3c+9\gamma^2})^{\sqrt{3}}\gamma_x}{2\gamma(c+3\gamma^2)} \right) \\ -\frac{\partial}{\partial y} \left(\frac{(3\gamma\sqrt{c+3\gamma^2}-c)\gamma_y}{2\gamma(c+3\gamma^2)(6\gamma+2\sqrt{3c+9\gamma^2})^{\sqrt{3}}} \right) = \gamma\sqrt{c+3\gamma^2}$$

defined on a simply-connected domain $D \subset \mathbf{R}^2$. Then $M_{\gamma} = (D, g_{\gamma})$ with the metric

(5.5)
$$g_{\gamma} = \frac{\sqrt{c+3\gamma^2}}{\gamma(6\gamma+2\sqrt{3c+9\gamma^2})^{\sqrt{3}}} \left(dx^2 + (6\gamma+2\sqrt{3c+9\gamma^2})^{2\sqrt{3}} dy^2 \right)$$

admits a non-minimal Wintgen ideal immersion $\psi_{\gamma}: M_{\gamma} \to R_2^4(c)$ into a complete simply-connected indefinite space form $R_2^4(c)$ satisfying $K^D = 2K$ identically. Space-like Surfaces

Proof. Let c be a real number and γ be positive solution of (5.4) with $3\gamma^2 > -c$ defined on a simply-connected domain D. Consider the surface $M_{\gamma} = (D, g_{\gamma})$ with metric g_{γ} given by (5.5). Then the Levi-Civita connection of g_{γ} satisfies

$$\nabla_{\frac{\partial}{\partial x}} \frac{\partial}{\partial x} = -\frac{(3\gamma\sqrt{c+3\gamma^2}+c)\gamma_x}{2\gamma(c+3\gamma^2)} \frac{\partial}{\partial x} + \frac{(3\gamma\sqrt{c+3\gamma^2}+c)\gamma_y}{2\gamma(c+3\gamma^2)(6\gamma+2\sqrt{3}c+9\gamma^2)^{2\sqrt{3}}} \frac{\partial}{\partial y}$$

$$\nabla_{\frac{\partial}{\partial x}} \frac{\partial}{\partial y} = -\frac{(3\gamma\sqrt{c+3\gamma^2}+c)\gamma_x}{2\gamma(c+3\gamma^2)} \frac{\partial}{\partial x} + \frac{(3\gamma\sqrt{c+3\gamma^2}-c)\gamma_x}{2\gamma(c+3\gamma^2)} \frac{\partial}{\partial y},$$

$$\nabla_{\frac{\partial}{\partial y}} \frac{\partial}{\partial y} = \frac{(6\gamma+2\sqrt{3}c+9\gamma^2)^{2\sqrt{3}}(c-3\gamma\sqrt{c+3\gamma^2})\gamma_x}{2\gamma(c+3\gamma^2)} \frac{\partial}{\partial x}$$

$$+ \frac{(3\gamma\sqrt{c+3\gamma^2}-c)\gamma_y}{2\gamma(c+3\gamma^2)} \frac{\partial}{\partial y}.$$

Let us define a bilinear map: $h: TM \to NM$ by

$$h\left(\frac{\partial}{\partial x}, \frac{\partial}{\partial x}\right) = -\frac{(\gamma + \sqrt{c + 3\gamma^2})\sqrt{c + 3\gamma^2}}{\gamma(6\gamma + 2\sqrt{3c + 9\gamma^2})^{\sqrt{3}}}e_3,$$
(5.7)
$$h\left(\frac{\partial}{\partial x}, \frac{\partial}{\partial y}\right) = -\sqrt{c + 3\gamma^2}e_4,$$

$$h\left(\frac{\partial}{\partial y}, \frac{\partial}{\partial y}\right) = \frac{(\gamma - \sqrt{c + 3\gamma^2})\sqrt{c + 3\gamma^2}(6\gamma + 2\sqrt{3c + 9\gamma^2})^{\sqrt{3}}}{\gamma}e_3,$$

where NM is the plane bundle over M spanned by an orthonormal time-like frame $\{e_3, e_4\}$. Define a linear metric connection D on NM by

$$D_{\frac{\partial}{\partial x}}e_{3} = \frac{-3\gamma\gamma_{y}e_{4}}{(c+3\gamma^{2})(6\gamma+2\sqrt{3c+9\gamma^{2}})^{\sqrt{3}}},$$

$$D_{\frac{\partial}{\partial y}}e_{3} = \frac{3\gamma(6\gamma+2\sqrt{3c+9\gamma^{2}})^{\sqrt{3}}\gamma_{x}}{c+3\gamma^{2}}e_{4},$$

$$D_{\frac{\partial}{\partial x}}e_{4} = \frac{3\gamma\gamma_{y}e_{3}}{(c+3\gamma^{2})(6\gamma+2\sqrt{3c+9\gamma^{2}})^{\sqrt{3}}},$$

$$D_{\frac{\partial}{\partial y}}e_{3} = -\frac{3\gamma(6\gamma+2\sqrt{3c+9\gamma^{2}})^{\sqrt{3}}\gamma_{x}}{c+3\gamma^{2}}e_{3}.$$

Then it follows from a very long direct computation that $(M_{\gamma}, g_{\gamma}, D, h)$ satisfies the equations of Gauss, Codazzi and Ricci. Hence, the fundamental existence and uniqueness theorem of submanifolds implies that, up to rigid motions, there exists a unique isometric immersion from M_{γ} into $R_2^4(c)$ whose second fundamental form and normal connection are given by h and D, respectively. By applying (5.5), (5.7) and $c + 3\gamma^2 > 0$ we see that M is a non-minimal Wintgen ideal surface in $R_2^4(c)$.

Now, we classify Wintgen ideal surfaces in $R_2^4(c)$ which satisfy $K^D = 2K$.

Theorem 5.3. Let M be a Wintgen ideal surface in a complete simply-connected indefinite space form $R_2^4(c)$ with c = 1, 0 or -1. If M satisfies $K^D = 2K$ identically, then one of following three cases occurs:

- (1) c = 0 and M is a totally geodesic surface in \mathbb{E}_{2}^{4} ;
- (2) c = -1 and M is a minimal surface in $H_2^4(-1)$ congruent to an open part of $\psi_{\mathcal{B}} : H^2(-\frac{1}{3}) \to H_2^4(-1) \subset \mathbb{E}_3^5$ defined by

$$\left(\sinh\left(\frac{2s}{\sqrt{3}}\right) - \frac{t^2}{3} - \left(\frac{7}{8} + \frac{t^4}{18}\right)e^{\frac{2s}{\sqrt{3}}}, t + \left(\frac{t^3}{3} - \frac{t}{4}\right)e^{\frac{2s}{\sqrt{3}}}, \\ \frac{1}{2} + \frac{t^2}{2}e^{\frac{2s}{\sqrt{3}}}, t + \left(\frac{t^3}{3} + \frac{t}{4}\right)e^{\frac{2s}{\sqrt{3}}}, \sinh\left(\frac{2s}{\sqrt{3}}\right) - \frac{t^2}{3} - \left(\frac{1}{8} + \frac{t^4}{18}\right)e^{\frac{2s}{\sqrt{3}}}\right);$$

(3) *M* is a non-minimal surface in $R_2^4(c)$ which is congruent to an open part of $\psi_{\gamma} : M_{\gamma} \to R_2^4(c)$ associated with a positive solution γ of the elliptic differential equation (5.4) as described in Theorem 5.2.

Proof. Let M be a Wintgen surface in $R_2^4(c)$. Then, according to Theorem 5.1, there exist an orthonormal frame $\{e_1, e_2, e_3, e_4\}$ such that shape operator satisfies (5.3) for some functions γ, μ . Thus, the Gauss and normal curvatures are given by

(5.9)
$$K = c + \gamma^2 - \mu^2 - 2\gamma\mu, \quad K^D = -2\gamma^2.$$

It follows from the condition $K^D = 2K$ that $\mu = -\gamma \pm \sqrt{c + 3\gamma^2}$. Without loss of generality, we may assume $\gamma \ge 0$.

Case (i). $\mu = -\gamma + \sqrt{c + 3\gamma^2}$. We divide this into two subcases.

Case (i.1). $c + 3\gamma^2 = 0$. We have $\mu = -\gamma$ and $c \le 0$. Thus, M is a minimal surface.

If c = 0, we get $\gamma = \mu = 0$, which implies that M is totally geodesic. So, we get case (1) of the theorem.

If c = -1, we have $\gamma = -\mu = \frac{1}{\sqrt{3}}$. Thus, by (5.9) *M* is a minimal surface with curvature $-\frac{1}{3}$. Hence, we obtain case (2) of the theorem according to Theorem 4.4.

Case (i.2). $c + 3\gamma^2 \neq 0$. From (5.3) we obtain

(5.10)
$$h(e_1, e_1) = -(\gamma + \sqrt{c + 3\gamma^2})e_3,$$
$$h(e_1, e_2) = -\gamma e_4,$$
$$h(e_2, e_2) = (\gamma - \sqrt{c + 3\gamma^2})e_3.$$

Thus, it follows from Codazzi's equation that

(5.11)
$$\omega_1^2(e_1) = \frac{3\gamma\sqrt{c+3\gamma^2+c}}{2\gamma(c+3\gamma^2)}e_2\gamma, \ \omega_1^2(e_2) = \frac{3\gamma\sqrt{c+3\gamma^2-c}}{2\gamma(c+3\gamma^2)}e_1\gamma,$$

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(5.12)
$$\omega_3^4(e_1) = -\frac{3\gamma e_2 \gamma}{c+3\gamma^2}, \ \ \omega_3^4(e_2) = \frac{3\gamma e_1 \gamma}{c+3\gamma^2}.$$

After applying (5.11) we derive that

$$\left[\frac{(c+3\gamma^2)^{1/4}}{\sqrt{\gamma}(6\gamma+2\sqrt{3c+9\gamma^2})^{\sqrt{3}/2}}e_1,\frac{(c+3\gamma^2)^{1/4}(6\gamma+2\sqrt{3c+9\gamma^2})^{\sqrt{3}/2}}{\sqrt{\gamma}}e_2\right]=0.$$

Hence there exist coordinates $\{x,y\}$ such that

(5.13)
$$\frac{\partial}{\partial x} = \frac{(c+3\gamma^2)^{1/4}}{\sqrt{\gamma}(6\gamma+2\sqrt{3c+9\gamma^2})^{\sqrt{3}/2}}e_1,$$
$$\frac{\partial}{\partial y} = \frac{(c+3\gamma^2)^{1/4}(6\gamma+2\sqrt{3c+9\gamma^2})^{\sqrt{3}/2}}{\sqrt{\gamma}}e_2.$$

By using (5.13) we know that the metric tensor is given by

(5.14)
$$g = \frac{\sqrt{c+3\gamma^2}}{\gamma(6\gamma+2\sqrt{3c+9\gamma^2})^{\sqrt{3}}} dx^2 + \frac{\sqrt{c+3\gamma^2}(6\gamma+2\sqrt{3c+9\gamma^2})^{\sqrt{3}}}{\gamma} dy^2,$$

which implies that the Levi-Civita connection satisfies

$$\nabla_{\frac{\partial}{\partial x}} \frac{\partial}{\partial x} = -\frac{(3\gamma\sqrt{c+3\gamma^2}+c)\gamma_x}{2\gamma(c+3\gamma^2)} \frac{\partial}{\partial x} + \frac{(3\gamma\sqrt{c+3\gamma^2}+c)\gamma_y}{2\gamma(c+3\gamma^2)(6\gamma+2\sqrt{3c+9\gamma^2})^{2\sqrt{3}}} \frac{\partial}{\partial y},$$

$$\nabla_{\frac{\partial}{\partial x}} \frac{\partial}{\partial y} = -\frac{(3\gamma\sqrt{c+3\gamma^2}+c)\gamma_x}{2\gamma(c+3\gamma^2)} \frac{\partial}{\partial x} + \frac{(3\gamma\sqrt{c+3\gamma^2}-c)\gamma_x}{2\gamma(c+3\gamma^2)} \frac{\partial}{\partial y},$$

$$\nabla_{\frac{\partial}{\partial y}} \frac{\partial}{\partial y} = \frac{(6\gamma+2\sqrt{3c+9\gamma^2})^{2\sqrt{3}}(c-3\gamma\sqrt{c+3\gamma^2})\gamma_x}{2\gamma(c+3\gamma^2)} \frac{\partial}{\partial x}$$

$$+ \frac{(3\gamma\sqrt{c+3\gamma^2}-c)\gamma_y}{2\gamma(c+3\gamma^2)} \frac{\partial}{\partial y}.$$

From (5.12) and (5.13) we find

(5.16)
$$\omega_3^4 \left(\frac{\partial}{\partial x}\right) = \frac{-3\gamma\gamma_y}{(c+3\gamma^2)(6\gamma+2\sqrt{3c+9\gamma^2})^{\sqrt{3}}},$$
$$\omega_3^4 \left(\frac{\partial}{\partial y}\right) = \frac{3\gamma(6\gamma+2\sqrt{3c+9\gamma^2})^{\sqrt{3}}}{c+3\gamma^2}\gamma_x.$$

Also, it follows from (5.10) and (5.13) that

$$h\left(\frac{\partial}{\partial x}, \frac{\partial}{\partial x}\right) = -\frac{(\gamma + \sqrt{c + 3\gamma^2})\sqrt{c + 3\gamma^2}}{\gamma(6\gamma + 2\sqrt{3c + 9\gamma^2})^{\sqrt{3}}}e_3,$$
(5.17)
$$h\left(\frac{\partial}{\partial x}, \frac{\partial}{\partial y}\right) = -\sqrt{c + 3\gamma^2}e_4,$$

$$h\left(\frac{\partial}{\partial y}, \frac{\partial}{\partial y}\right) = \frac{(\gamma - \sqrt{c + 3\gamma^2})\sqrt{c + 3\gamma^2}(6\gamma + 2\sqrt{3c + 9\gamma^2})^{\sqrt{3}}}{\gamma}e_3.$$

Moreover, from (5.10), (5.15) and the equation of Gauss we know that γ satisfies the elliptic differential equation (5.4). Consequently, after applying Theorem 5.2 we obtain case (3) of the theorem.

Case (ii). $\mu = -\gamma - \sqrt{c + 3\gamma^2}$. After replacing e_3, e_4 by $-e_3, -e_4$, respectively, this reduces to (i).

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