# Umbilics of surfaces in the Minkowski 3-space 

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#### Abstract

We prove that any closed and convex surface in the Minkowski 3 -space of class $C^{3}$ has at least two umbilic points. This shows that the Carathéodory conjecture for surfaces in the Euclidean 3 -space is true for surfaces in the Minkowski 3-space.


## 1. Introduction.

The Carathéodory conjecture states that any smooth closed and convex surface in the Euclidean 3 -space has at least two umbilic points. Various attempts were made to prove this conjecture (see for example [4] for a survey article and [3] for the latest results on the problem using the mean curvature flow on the space of oriented lines in $\mathbb{R}^{3}$ ).

We prove in this paper that any closed and convex surface in the Minkowski 3 -space of class $C^{3}$ has at least two umbilic points (Theorem 3.3). For ovaloids, we can even specify the nature of the umbilic points (Theorem 3.4). We give some preliminaries in Section 2 and prove the main results in Section 3.

## 2. Preliminaries.

The Minkowski space $\left(\mathbb{R}_{1}^{3},\langle\rangle,\right)$ is the vector space $\mathbb{R}^{3}$ endowed with the pseudoscalar product $\langle\boldsymbol{u}, \boldsymbol{v}\rangle=-u_{0} v_{0}+u_{1} v_{1}+u_{2} v_{2}$, for any $\boldsymbol{u}=\left(u_{0}, u_{1}, u_{2}\right)$ and $\boldsymbol{v}=$ $\left(v_{0}, v_{1}, v_{2}\right)$ in $\mathbb{R}_{1}^{3}$. We say that a non-zero vector $\boldsymbol{u} \in \mathbb{R}_{1}^{3}$ is spacelike if $\langle\boldsymbol{u}, \boldsymbol{u}\rangle>0$, lightlike if $\langle\boldsymbol{u}, \boldsymbol{u}\rangle=0$ and timelike if $\langle\boldsymbol{u}, \boldsymbol{u}\rangle<0$. The norm of a vector $\boldsymbol{u} \in \mathbb{R}_{1}^{3}$ is defined by $\|\boldsymbol{u}\|=\sqrt{|\langle\boldsymbol{u}, \boldsymbol{u}\rangle|}$. The set of all lightlike vectors form the lightcone

$$
L C^{*}=\left\{\boldsymbol{u} \in \mathbb{R}_{1}^{3} \backslash\{\underline{0}\} \mid\langle\boldsymbol{u}, \boldsymbol{u}\rangle=0\right\} .
$$

The lightcone can be considered as the cone in $\mathbb{R}^{3}$ minus the origin given by

$$
\left\{\left(u_{0}, u_{1}, u_{2}\right) \in \mathbb{R}^{3} \backslash\{(0,0,0)\} \mid-u_{0}^{2}+u_{1}^{2}+u_{2}^{2}=0\right\} .
$$

[^0]A plane $P_{c}^{\boldsymbol{v}}=\left\{\boldsymbol{u} \in \mathbb{R}_{1}^{3} \mid\langle\boldsymbol{u}, \boldsymbol{v}\rangle=c\right\}$, for some constant $c \in \mathbb{R}$, is called respectively, spacelike, timelike or lightlike if $\boldsymbol{v}$ is timelike, spacelike or lightlike. Fixing $\boldsymbol{v}$ and varying $c$ gives a family of parallel planes with $P_{0}^{\boldsymbol{v}}$ passing through the origin (i.e., is a vector space). The vector $\boldsymbol{v}$ is called the "normal" vector to $P_{c}^{v}$. Every non-zero vector in a spacelike plane $P_{0}^{v}$ is spacelike. There are two linearly independent lightlike vectors in a timelike plane $P_{0}^{v}$ and a unique lightlike vector in a lightlike plane $P_{0}^{v}$. The normal vector $\boldsymbol{v}$ is transverse to $P_{c}^{v}$ if this plane is spacelike or timelike but determines the unique lightlike direction in $P_{0}^{v}$ if the plane $P_{c}^{v}$ is lightlike.

Let $S$ be a surface in $\mathbb{R}_{1}^{3}$ (of class $C^{3}$ ). The pseudo-scalar product in $\mathbb{R}_{1}^{3}$ induces a metric on $S$. If $S$ is closed, then this metric must be degenerate at some point on $S$ (see for example Lemma 3.1). This happens at points $p$ on $S$ where the tangent space $T_{p} S$ is a lightlike plane. We call the locus of points where the induced metric on $S$ is degenerate the Locus of Degeneracy and denote it by $L D$.

Let $\boldsymbol{x}: U \subset \mathbb{R}^{2} \rightarrow S$ be a local parametrisation of $S$ and let

$$
E=\left\langle\boldsymbol{x}_{u}, \boldsymbol{x}_{u}\right\rangle, \quad F=\left\langle\boldsymbol{x}_{u}, \boldsymbol{x}_{v}\right\rangle, \quad G=\left\langle\boldsymbol{x}_{v}, \boldsymbol{x}_{v}\right\rangle
$$

denote the coefficients of the first fundamental form of $S$ with respect to $\boldsymbol{x}$. We identify the $L D$ and its pre-image in $U$ by $\boldsymbol{x}$. Then the $L D$ (in $U$ ) is the zero set of the $C^{2}$-function $\delta(u, v)=\left(F^{2}-E G\right)(u, v)$. Therefore, the $L D$ is a closed subset of $S$. We observe that the $L D$ of a generic closed surface is a smooth curve, but we do not make the genericity assumption here. We can have, for instance, a convex surface with an $L D$ that has interior points.

Pei [6] defined an $\mathbb{R} P^{2}$-valued Gauss map on $S$. In $\boldsymbol{x}(U)$, this is simply the map $P N: \boldsymbol{x}(U) \rightarrow \mathbb{R} P^{2}$ which associates to a point $p=\boldsymbol{x}(q)$ the projectivisation of the vector $\left(\boldsymbol{x}_{u} \times \boldsymbol{x}_{v}\right)(q)$, where " $\times$ " denotes the wedge product in $\mathbb{R}_{1}^{3}$. Away from the $L D$, the $\mathbb{R} P^{2}$-valued Gauss map can be identified with the de Sitter Gauss map $\boldsymbol{x}\left(U_{1}\right) \rightarrow S_{1}^{2}$ on the Lorentzian part of the surface and with the hyperbolic Gauss map $\boldsymbol{x}\left(U_{2}\right) \rightarrow H^{2}(-1)$ on its Riemannian part. (Here $U_{1}$ and $U_{2}$ are open sets with $U=U_{1} \cup U_{2} \cup L D$.) Both maps are given by $\boldsymbol{N}=\boldsymbol{x}_{u} \times \boldsymbol{x}_{v} /\left\|\boldsymbol{x}_{u} \times \boldsymbol{x}_{v}\right\|$. The shape operator $A_{p}(\boldsymbol{v})=-d \boldsymbol{N}_{p}(\boldsymbol{v})$ is a self-adjoint operator on $\boldsymbol{x}(U) \backslash L D$. We denote by

$$
\begin{aligned}
l & =-\left\langle\boldsymbol{N}_{u}, \boldsymbol{x}_{u}\right\rangle \\
m & =\left\langle\boldsymbol{N}, \boldsymbol{x}_{u u}\right\rangle, \\
n & =-\left\langle\boldsymbol{N}_{u}, \boldsymbol{x}_{v}\right\rangle
\end{aligned}=\left\langle\boldsymbol{N}, \boldsymbol{x}_{u v}\right\rangle, ~\left\{\boldsymbol{N}_{v}, \boldsymbol{x}_{v}\right\rangle=\left\langle\boldsymbol{N}, \boldsymbol{x}_{v v}\right\rangle, ~ \$
$$

the coefficients of the second fundamental form on $\boldsymbol{x}(U) \backslash L D$. When $A_{p}$ has real
eigenvalues, we call them the principal curvatures and their associated eigenvectors the principal directions of $S$ at $p$. We observe that there are always two principal curvatures at points in the Riemannian part of $S$ but this is not true at points in its Lorentzian part ([5]). The lines of principal curvature, which are the integral curves of the principal directions, are solutions of the binary quadratic differential equation (BDE for short)

$$
\begin{equation*}
(G m-F n) d v^{2}+(G l-E n) d v d u+(F l-E m) d u^{2}=0 \tag{1}
\end{equation*}
$$

The discriminant of the above equation, which is the set points in $U \backslash L D$ where it determines a unique direction, is denoted the Lightlike Principal Locus (LPL) in [5]. It is the zero set of the function $\left((G l-E n)^{2}-4(G m-F n)(F l-E m)\right)(u, v)$.

A spacelike umbilic point (resp. timelike umbilic point) is a point in the Riemannian part (resp. Lorentzian part) of the surface where the coefficients of equation (1) vanish simultaneously. (The coefficients of a BDE are its coefficients when viewed as a quadratic form in $d u$ and $d v$.) Spacelike and timelike umbilic points can also be characterised as the points $p$ where the shape operator $A_{p}$ is a multiple of the identity map.

One can extend the lines of principal curvature across the $L D$ as follows ([5]). As equation (1) is homogeneous in $l, m, n$, we can multiply these coefficients by $\left\|\boldsymbol{x}_{u} \times \boldsymbol{x}_{v}\right\|$ and substitute them by

$$
\bar{l}=\left\langle\boldsymbol{x}_{u} \times \boldsymbol{x}_{v}, \boldsymbol{x}_{u u}\right\rangle, \quad \bar{m}=\left\langle\boldsymbol{x}_{u} \times \boldsymbol{x}_{v}, \boldsymbol{x}_{u v}\right\rangle, \quad \bar{n}=\left\langle\boldsymbol{x}_{u} \times \boldsymbol{x}_{v}, \boldsymbol{x}_{v v}\right\rangle .
$$

This substitution does not alter the pair of foliations on $\boldsymbol{x}(U) \backslash L D$. The new equation is defined on the $L D$ and defines the same pair of foliations associated to the de Sitter (resp. hyperbolic) Gauss map on the Lorentzian (resp. Riemannian) part of $\boldsymbol{x}(U)$. The extended lines of principal curvature are the solution curves of the BDE

$$
\begin{equation*}
(G \bar{m}-F \bar{n}) d v^{2}+(G \bar{l}-E \bar{n}) d u d v+(F \bar{l}-E \bar{m}) d u^{2}=0 . \tag{2}
\end{equation*}
$$

We still call the directions determined by equation (2) at points on the $L D$ principal directions. We do not have a shape operator at points on the $L D$. For this reason, we define a lightlike umbilic point as a point on the $L D$ where the coefficients of equation (2) vanish simultaneously.

We say that a point on $S$ is an umbilic point if it is either a spacelike, timelike or lightlike umbilic point. Thus, a point $p=\boldsymbol{x}(q)$ is an umbilic point if and only if all the coefficients of equation (2) vanish at $q$.

Remark 2.1. The lines of principal curvatures on a generic surface in $\mathbb{R}_{1}^{3}$ are studied in [5]. On the Riemannian part of a generic surface, the $L P L$, when not empty, consists of isolated points which are spacelike umbilic points. Away from these points, there are always two orthogonal spacelike principal directions. On the Lorentzian part of a generic surface, the $L P L$, when not empty, is a smooth curve except at isolated points where it has Morse singularities of node type. The singular points of the $L P L$ are precisely the timelike umbilic points. The regular points of the $L P L$ consist of points where the principal directions coincide and become lightlike. There are two principal directions on one side of the $L P L$ and none on the other. When there are two of them, they are orthogonal and one is spacelike while the other is timelike.

Equation (2) determines two directions in $T_{p} S$ at most points $p$ on the $L D$. One of these directions is the unique lightlike direction in $T_{p} S$ and the other is a spacelike. The two directions coincide and become the unique lightlike direction in $T_{p} S$ at isolated points $p$ on the $L D$. Generic surfaces do not have lightlike umbilic points. The generic local topological configurations of the lines of principal curvature at points on the $L P L$ and on the $L D$ are given in [5].

We consider here closed and convex surfaces in $\mathbb{R}_{1}^{3}$. Convexity is an affine property so is independent of the metric (Euclidean or Lorentzian) in $\mathbb{R}^{3}$.

We also consider some special convex surfaces. An ovaloid in the Euclidean space $\mathbb{R}^{3}$ is defined as a surface with everywhere strictly positive Gaussian curvature $K$. We do not have the concept of Gaussian curvature of a surface in the Minkowski space $\mathbb{R}_{1}^{3}$ at point on the $L D$. (In fact, for generic surfaces, the Gaussian curvature tends to infinity as a point on the $L D$ is approached from either the Riemannian or the Lorentzian part of the surface; see [7].) However we can still define the concept of ovaloids using the contact of the surface with planes (which is an affine property of the surface).

Let $P_{c}^{\boldsymbol{v}}=\left\{p \in \mathbb{R}^{3} \mid\langle p, \boldsymbol{v}\rangle=c\right\}$ be a plane in $\mathbb{R}_{1}^{3}$. The contact of a surface $S$ with $P_{c}^{v}$ is measured by the singularities of the height function $h: S \rightarrow \mathbb{R}$, given by

$$
h(p)=\langle p, \boldsymbol{v}\rangle .
$$

We say that the contact is of type $A_{1}^{+}$at $p \in S$ if $p \in P_{c}^{\boldsymbol{v}}$ and the height function $h$ has a Morse singularity of index 0 or 2 at $p$, i.e., $h$ can be written in some local coordinate system at $p$ in $S$ in the form $\pm\left(u^{2}+v^{2}\right)$. For this, it is necessary and sufficient for the Taylor polynomial of degree 2 of $h$ at $p$ to be a strictly positive or a strictly negative quadratic form.

We say that a closed and convex surface $S$ is an ovaloid if it has an $A_{1}^{+}$-contact with its tangent plane $T_{p} S$ at all $p \in S$. An example of an ovaloid in the Minkowski
space $\mathbb{R}_{1}^{3}$ is (the "Euclidean sphere")

$$
S^{2}=\left\{\left(u_{0}, u_{1}, u_{2}\right) \in \mathbb{R}_{1}^{3} \mid u_{0}^{2}+u_{1}^{2}+u_{2}^{2}=1\right\} .
$$

The surface $S^{2} \subset \mathbb{R}_{1}^{3}$ has two umbilic points ([5, Section 4.4]), so is not a totally umbilic surface. (See [2] for the study of geodesics on an ellipsoid in $\mathbb{R}_{1}^{3}$.)

A surface $S$ is locally convex at $p \in S$ if there exists a neighbourhood $V$ of $p$ in $S$ such that $V$ is contained in one of the closed half-spaces determined by the tangent plane $T_{p} S$. A convex surface is of course locally convex. Given a local parametrisation $\boldsymbol{x}: U \rightarrow S$ of the surface $S$ and $q_{0}=\left(u_{0}, v_{0}\right) \in U$, the height function $h$ along the "normal vector" $\boldsymbol{v}=\left(\boldsymbol{x}_{u} \times \boldsymbol{x}_{v}\right)\left(q_{0}\right)$ at $q_{0}$ can be considered locally as a map $U \rightarrow \mathbb{R}$, given by $h(u, v)=\langle\boldsymbol{x}(u, v), \boldsymbol{v}\rangle$. The Taylor polynomial of degree 2 of $h$ at $q_{0}=\left(u_{0}, v_{0}\right) \in U$ is given by

$$
\frac{1}{2}\left(h_{u u}\left(q_{0}\right)\left(u-u_{0}\right)^{2}+2 h_{u v}\left(q_{0}\right)\left(u-u_{0}\right)\left(v-v_{0}\right)+h_{v v}\left(q_{0}\right)\left(v-v_{0}\right)^{2}\right)
$$

and a necessary condition for $S$ to be locally convex at $p_{0}=\boldsymbol{x}\left(q_{0}\right)$ is that

$$
\left(h_{u v}^{2}-h_{u u} h_{v v}\right)\left(q_{0}\right) \leq 0 .
$$

The above condition is true at any point on $S$ including points on the $L D$.

## 3. The main results.

The proof of the main result relies on the structure of the $L D$ and on the directions determined by equation (2) on this set.

Lemma 3.1. The LD of a closed surface $S$ in $\mathbb{R}_{1}^{3}$ of class $C^{1}$ is the union of at least two disjoint non-empty closed subsets of $S$.

Proof. The $L D$ is the set of points on $S$ where the tangent plane to $S$ is lightlike. Lightlike planes are those tangent to the lightcone $L C^{*}$ and a key observation is that these planes can be captured by changing the metric on $\mathbb{R}^{3}$.

We change the metric in $\mathbb{R}^{3}$ and consider $S \subset \mathbb{R}_{1}^{3}$ as a surface $\tilde{S}$ in the Euclidean space $\mathbb{R}^{3}$. Since $\tilde{S}$ is closed, the image of its Gauss map $N: \tilde{S} \rightarrow S^{2}$ is the whole sphere $S^{2}$.

The unit Euclidean normals to the tangent planes to $L C^{*}$ (viewed as a cone in $\mathbb{R}^{3}$, see Section 1) trace the two circles $u_{0}= \pm 1 / \sqrt{2}$ on $S^{2}$. The $L D$ of $S$ is precisely the pre-image of the two circles $u_{0}= \pm 1 / \sqrt{2}$ by the Gauss map $N$ on $\tilde{S}$. Therefore, the $L D$ consists of at least two disjoint non-empty closed subsets of $S$.

Lemma 3.2. Let $S$ be a closed and convex surface in $\mathbb{R}_{1}^{3}$ of class $C^{3}$ and $\boldsymbol{x}: U \rightarrow S$ a local parametrisation of $S$.
(1) The singular points of $\delta=F^{2}-E G$ on the $L D$ are lightlike umbilic points.
(2) The unique lightlike principal direction in $T_{p} S$ at the regular points of $\delta$ on the $L D$ is transverse to the $L D$.

Proof. If $E(q)=0$ or $G(q)=0$ at $q \in U$ with $\boldsymbol{x}(q) \in L D$, then $F(q)=0$. Therefore, we cannot have $E(q)=G(q)=0$ at points on the $L D$. We assume, without loss of generality, that $G \neq 0$ on $U$.

The lightlike directions at points in $\boldsymbol{x}(U)$ are solutions of the equation

$$
G d v^{2}+2 F d u d v+E d u^{2}=0
$$

and the unique lightlike direction on the $L D$ is parallel to $G \boldsymbol{x}_{u}-F \boldsymbol{x}_{v}$. This is a smooth vector field on $\boldsymbol{x}(U)$, so we can re-parametrise $\boldsymbol{x}(U)$ so that one of the coordinate curves are the integral curves of this vector field. That is, we can choose a local parametrisation of $S$, that we still denote by $\boldsymbol{x}$, so that the unique lightlike direction on the $L D$ is along $\boldsymbol{x}_{u}$. With this parametrisation, that we use in the rest of the proof, $E=F=0$ on the $L D$.
(1) The function $\delta$ is singular on the $L D$ if and only if $\left(-E_{u} G,-E_{v} G\right)=(0,0)$. The coefficients of equation (2) become ( $G \bar{m}, G \bar{l}, 0$ ) on the $L D$, and on this set we also have $\boldsymbol{x}_{u} \times \boldsymbol{x}_{v}=\lambda \boldsymbol{x}_{u}$ for some non-zero function $\lambda$. Therefore,

$$
\bar{l}=\left\langle\boldsymbol{x}_{u} \times \boldsymbol{x}_{v}, \boldsymbol{x}_{u u}\right\rangle=\lambda\left\langle\boldsymbol{x}_{u}, \boldsymbol{x}_{u u}\right\rangle=\frac{1}{2} \lambda E_{u},
$$

and similarly,

$$
\bar{m}=\left\langle\boldsymbol{x}_{u} \times \boldsymbol{x}_{v}, \boldsymbol{x}_{u v}\right\rangle=\lambda\left\langle\boldsymbol{x}_{u}, \boldsymbol{x}_{u v}\right\rangle=\frac{1}{2} \lambda E_{v} .
$$

Thus, the coefficients of equation (2) at points on the $L D$ are $\left(\lambda E_{v} G, \lambda E_{u} G, 0\right)$, and this is $(0,0,0)$ at a point $q \in L D$ if and only if the $\delta$ is singular at $q$.
(2) Suppose now that $\delta$ is regular on the $L D$ (so the $L D$ is a regular curve). Then we have either $E_{u} \neq 0$ or $E_{v} \neq 0$ on this curve. We consider the contact of $S$ with its tangent plane $T_{p_{0}} S$ at $p_{0}=\boldsymbol{x}\left(q_{0}\right) \in L D$. The Taylor polynomial of degree 2 of the height function $h(u, v)=\left\langle\boldsymbol{x}(u, v), \boldsymbol{x}_{u}\left(q_{0}\right)\right\rangle$ along the lightlike "normal vector" $\boldsymbol{x}_{u}\left(q_{0}\right)$ at $q_{0}$ is given by

$$
\frac{1}{2}\left(h_{u u}\left(q_{0}\right)\left(u-u_{0}\right)^{2}+2 h_{u v}\left(q_{0}\right)\left(u-u_{0}\right)\left(v-v_{0}\right)+h_{v v}\left(q_{0}\right)\left(v-v_{0}\right)^{2}\right)
$$

with

$$
\begin{aligned}
& h_{u u}\left(q_{0}\right)=\left\langle\boldsymbol{x}_{u u}\left(q_{0}\right), \boldsymbol{x}_{u}\left(q_{0}\right)\right\rangle=\frac{1}{2} E_{u}\left(q_{0}\right), \\
& h_{u v}\left(q_{0}\right)=\left\langle\boldsymbol{x}_{u v}\left(q_{0}\right), \boldsymbol{x}_{u}\left(q_{0}\right)\right\rangle=\frac{1}{2} E_{v}\left(q_{0}\right), \\
& h_{v v}\left(q_{0}\right)=\left\langle\boldsymbol{x}_{v v}\left(q_{0}\right), \boldsymbol{x}_{u}\left(q_{0}\right)\right\rangle=\left(F_{v}-\frac{1}{2} G_{u}\right)\left(q_{0}\right) .
\end{aligned}
$$

The lightlike direction $\boldsymbol{x}_{u}\left(q_{0}\right)$ is tangent to the $L D$ at $p_{0}=\boldsymbol{x}\left(q_{0}\right)$ if and only if $E_{u}\left(q_{0}\right)=0$. But as $S$ is convex, $\left(h_{u v}^{2}-h_{u u} h_{v v}\right)\left(q_{0}\right)=\left(E_{v}^{2}-2\left(F_{v}-\right.\right.$ $\left.\left.\left(G_{u} / 2\right)\right) E_{u}\right)\left(q_{0}\right) / 4 \leq 0$, so $E_{u}\left(q_{0}\right)=0$ implies $E_{v}\left(q_{0}\right)=0$, and consequently the $L D$ is singular. Therefore, $E_{u} \neq 0$ at regular points of $\delta$ on the $L D$, that is, the lightlike principal direction is transverse to the $L D$ at the regular points of $\delta$ on this set.

Theorem 3.3. Let $S$ be a closed and convex surface of class $C^{3}$ in $\mathbb{R}_{1}^{3}$. Then $S$ has at least two umbilic points.

Proof. Consider the $C^{3}$-function $f: S \rightarrow \mathbb{R}$ given by $f(p)=p_{0}$ for any $p=\left(p_{0}, p_{1}, p_{2}\right) \in S$. It has a global minimum $p_{\min }$ and a global maximum $p_{\max }$ (these points need not be unique). The tangent planes to $S$ at $p_{\min }$ and $p_{\max }$ are spacelike (both are given by $u_{0}=0$ ). Therefore, $p_{\min }$ and $p_{\max }$ belong to the Riemannian part of $S$. Suppose that they belong to the same Riemannian connected component $R$ of $S$. Let $\gamma:[0,1] \rightarrow R$ be a continuous path in $R$ with $\gamma(0)=p_{\min }$ and $\gamma(1)=p_{\text {max }}$, and consider the Gauss map $N: \tilde{S} \rightarrow S^{2}$ as in the proof of Lemma 3.1. The continuous curve $N \circ \gamma$ satisfies $N \circ \gamma(0)=(-1,0,0)$ and $N \circ \gamma(1)=(1,0,0)$, so there exists $t_{0} \in(0,1)$ such that $N \circ \gamma\left(t_{0}\right)$ belongs the equator $u_{0}=0$ on $S^{2}$. Therefore, the tangent space to $S$ at $\gamma\left(t_{0}\right)$ is a timelike plane, which is a contradiction as $R$ is supposed to be Riemannian.

Let $R_{1}$ (resp. $R_{2}$ ) denotes the Riemannian connected component of $S$ which contains $p_{\text {min }}$ (resp. $p_{\text {max }}$ ) and let $L_{1}$ (resp. $L_{2}$ ) be its boundary. The sets $L_{1}$ and $L_{2}$ are part of the $L D$. It follows from the proof of Lemma 3.1 that $L_{1}$ and $L_{2}$ are disjoint sets ( $L_{1}$ is part of the pre-image of the circle $u_{0}=-1 / \sqrt{2}$ by the Gauss map $N$, and $L_{2}$ is part of the pre-image of the circle $u_{0}=1 / \sqrt{2}$ by $N$ ).

We consider local parametrisations of $S$ at points on $L_{1}$ and $L_{2}$. If $\delta=$ $F^{2}-E G$ is singular on $L_{1}$ and $L_{2}$, then the singular points are lightlike umbilic points (Lemma 3.2(1)). As $L_{1}$ and $L_{2}$ are disjoint, we get at least two umbilic
points on $S$.
Suppose that $\delta$ is regular on $L_{1}$ (so $L_{1}$ is a regular curve; it is also a closed curve). The surface $S$ being closed and convex is homeomorphic to a 2 -sphere. Thus $R_{1}$ is homeomorphic to a disc. Consider the direction field in $R_{1}$ given by equation (2) and which agrees with the unique lightlike direction in $T_{p} S$ for all $p \in L_{1}$. This direction field is transverse to $L_{1}$ (Lemma 3.2(2)), so by PoincaréHopf theorem it must have at least one singularity in $R_{1}$. This singularity is a spacelike umbilic point as $R_{1}$ is a Riemannian region. We proceed similarly if $\delta$ is regular on $L_{2}$ to get a second umbilic point of $S$. If $\delta$ is singular at a point on $L_{2}$, the singularity is a lightlike umbilic point and gives a second umbilic point of $S$.

We showed in [5, Section 4.4] that (the Euclidean sphere) $S^{2}$ has exactly two umbilic points and both of them are spacelike. We have the following general result.

Theorem 3.4. The umbilic points of an ovaloid in $\mathbb{R}_{1}^{3}$ of class $C^{3}$ are all spacelike and there are at least two of them.

Proof. We change the metric in $\mathbb{R}^{3}$ and consider an ovaloid $S \subset \mathbb{R}_{1}^{3}$ as a surface $\tilde{S}$ in the Euclidean space $\mathbb{R}^{3}$. The fact that the contact of $S$ with its tangent plane (which is independent of the metric) is $A_{1}^{+}$implies that the Gaussian curvature of $\tilde{S}$ is strictly positive. By Hadamard's theorem, the Gauss $\operatorname{map} N: \tilde{S} \rightarrow S^{2}$ is a diffeomorphism ([1]). This implies that the $L D$ of $S$ is the union of two regular (non-empty) disjoint closed curves. These split the surface into three parts, two of them are Riemannian and one is Lorentzian. By Lemma $3.2(2)$, the unique lightlike principal direction on the $L D$ is transverse to the $L D$. By Poincaré-Hopf theorem, there is at least one spacelike umbilic point in each Riemannian disc of $S$.

We now show that there are no timelike umbilic points on $S$. The timelike umbilic points occur in the Lorentzian part of the surface, so we can take a local parametrisation $\boldsymbol{x}: U \subset \mathbb{R}^{2} \rightarrow \mathbb{R}_{1}^{3}$ where the coordinate curves are lightlike (Theorem 3.1 in [5]). Then $E=G=0$ in $U$. The equation of the lines of principal curvature simplifies to $n d v^{2}-l d u^{2}=0$, so the timelike umbilic points are the solutions of $l=n=0$.

Let $q_{0}=\left(u_{0}, v_{0}\right) \in U$ and consider the height function $h$ on $\boldsymbol{x}(U)$ along the unit normal vector $\boldsymbol{N}\left(q_{0}\right)=\left(\boldsymbol{x}_{u} \times \boldsymbol{x}_{v} /\left\|\boldsymbol{x}_{u} \times \boldsymbol{x}_{v}\right\|\right)\left(q_{0}\right)$. The Taylor polynomial of $h(u, v)=\left\langle\boldsymbol{x}(u, v), \boldsymbol{N}\left(q_{0}\right)\right\rangle$ at $q_{0}$ is given by

$$
\frac{1}{2}\left(l\left(q_{0}\right)\left(u-u_{0}\right)^{2}+2 m\left(q_{0}\right)\left(u-u_{0}\right)\left(v-v_{0}\right)+n\left(q_{0}\right)\left(v-v_{0}\right)^{2}\right)
$$

where $l, m, n$ are the coefficients of the second fundamental form. As $S$ is an ovaloid, $\left(m^{2}-n l\right)(q)<0$ for any $q \in U$ and consequently $l(q) n(q) \neq 0$ at any $q \in U$. This proves that there are no timelike umbilic points on $S$.

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