# LOCALLY UNIVALENT FUNCTIONS AND COEFFICIENT DISTORTIONS 

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#### Abstract

We look at functions $f(z)=z+\sum_{n=2}^{\infty} a_{n} z^{n}$ satisfying $\sum_{n=2}^{\infty} n\left|a_{n}\right|>1$ and determine conditions for which the arguments of the coefficients may vary without affecting the univalence of the function. A bound on the radius of starlikeness for the convolution of functions taken from the closed convex hull of convex functions and a special subclass of starlike functions is also obtained.


1. Introduction. Let $L S$ denote the class of functions of the form $z+\sum_{n=2}^{\infty} a_{n} z^{n}$ that are analytic and locally univalent in the unit disk $U$, and let $S$ denote the subclass of univalent functions. It is well known that a sufficient condition for $z+\sum_{n=2}^{\infty} a_{n} z^{n}$ to be in $S$ is that

$$
\begin{equation*}
\sum_{n=2}^{\infty} n\left|a_{n}\right| \leqq 1 \tag{1}
\end{equation*}
$$

For functions of the form

$$
\begin{equation*}
z-\sum_{n=2}^{\infty}\left|a_{n}\right| z^{n} \tag{2}
\end{equation*}
$$

the condition (1) is also necessary. This follows because functions that fail to satisfy (1) are not even in $L S$. The necessary and sufficient condition (1) for functions of the form (2) to be in $S$ makes extremal problems much more manageable. Very little is known for functions in $S$ of the form

$$
\begin{equation*}
f_{k}(z)=z-e^{i \lambda} \sum_{n=2}^{\infty}\left|a_{n}\right| z^{n}, \tag{3}
\end{equation*}
$$

where the coefficients are not necessarily real but have constant argument. In [6] it is asked if a function $g(\lambda, n)$ can be found for which the inequality $\left|\alpha_{n}\right| \leqq g(\lambda, n)$ is sharp. Note that $g(0, n)=1 / n$ and $g(\pi, n)=n$, with extremal functions $z-z^{n} / n$ and $z /(1-z)^{2}$ respectively.

In this paper we show that a function in $L S$ must satisfy (1) when its coefficients are "close to" negative. Since the degree of closeness depends on both $\lambda$ and the coefficients in (3), rather than on $\lambda$ alone, we cannot conclude that $g(\lambda, n) \leqq 1 / n$ for any positive $\lambda$. We also examine the extent to which a violation of condition (1)
enables us to distort the arguments of some of the coefficients to construct functions that are not in $L S$.

The Hadamard product or convolution of two power series $f(z)=$ $\sum_{n=0}^{\infty} a_{n} z^{n}$ and $g(z)=\sum_{n=0}^{\infty} b_{n} z^{n}$ is defined as the power series $(f * g)(z)=$ $\sum_{n=0}^{\infty} a_{n} b_{n} z^{n}$. For

$$
\begin{equation*}
h_{\lambda}(z)=z+e^{i \lambda} \sum_{n=2}^{\infty} z^{n} \tag{4}
\end{equation*}
$$

we may express (3) as $f_{0} * h_{\lambda}$. When $\lambda$ is sufficiently small $h_{\lambda}(z)$ is starlike and we find a bound on the radius of starlikeness for $h_{2} * f$, where $\operatorname{Re} f(z) / z>1 / 2(z \in U)$. This generalizes a result of MacGregor [4].
2. Coefficient distortions. Given a function $f(z)=z+\sum_{n=2}^{\infty} b_{n} z^{n}$, the function $g(z)=z+\sum_{n=2}^{\infty} e^{i \lambda_{n}} b_{n} z^{n}$ is said to be in $F_{\varepsilon}(f)$ for some $\varepsilon>0$ if $-\varepsilon \leqq \lambda_{n} \leqq \varepsilon$ for all $n$.

Lemma 1. If $f(z)=z+\sum_{n=2}^{\infty} b_{n} z^{n} \notin L S$, then there exists an $\varepsilon>0$ such that $g \in F_{\epsilon}(f)$ implies that $g \notin L S$.

Proof. Suppose, on the contrary, that there is a sequence $\varepsilon(n)$ tending to 0 for which we can find a corresponding sequence of functions $g_{n} \in F_{\varepsilon(n)}$ such that $g_{n} \in L S$ for all $n$. Since $f \notin L S$, there exists a point $z_{0} \in U$ such that $f^{\prime}\left(z_{0}\right)=0$. Note that $\left\{g_{n}^{\prime}\right\}$ converges uniformly to $f^{\prime}$ in some neighborhood $D$ of $z_{0}$. Since $g_{n}^{\prime} \neq 0$ in $D$ for any $n$, it follows by Hurwitz's theorem that $f^{\prime} \neq 0$ in $D$. This contradicts our assumption that $f^{\prime}\left(z_{0}\right)=0$.

Theorem 1. If $f(z)=z-\sum_{n=2}^{\infty}\left|a_{n}\right| z^{n}$ with $\sum_{n=2}^{\infty} n\left|a_{n}\right|>1$, then there exists an $\varepsilon>0$ such that $g \in F_{\varepsilon}(f)$ implies that $g \notin L S$.

Proof. Since $f^{\prime}(0)=1$ and $f^{\prime}(r)<0$ for $r(<1)$ sufficiently close to 1 , there must exist a point $r_{0}, 0<r_{0}<1$, such that $f^{\prime}\left(r_{0}\right)=0$. The result now follows from the lemma, with $b_{n}=-\left|a_{n}\right|$.

Corollary. If $\sum_{n=2}^{\infty} n\left|a_{n}\right|>1$, then $f_{2}(z)$, defined by (3), is not in LS for $\lambda$ sufficiently small.

Our next theorem shows that if (1) is violated, then we can always construct a nonunivalent function by distorting the arguments of finitely many coefficients.

THEOREM 2. If $\sum_{n=2}^{N} n\left|a_{n}\right|>1$, then there exist real numbers $\alpha_{2}, \cdots, \alpha_{N}\left(-\pi<\alpha_{j} \leqq \pi\right)$ such that

$$
f(z)=z+\sum_{n=2}^{N} a_{n} e^{i \alpha_{n}} z^{n}+\sum_{n=N+1}^{\infty} a_{n} z^{n} \notin L S
$$

Proof. Let $C=z(t)$ be an arc of increasing modulus from the origin to the boundary of $U$ such that

$$
g(z)=\sum_{n=N+1}^{\infty} n a_{n} z^{n-1} \leqq 0 \quad \text { for } \quad z \in C
$$

For each point $z=z(t) \in C,|z|=r$, choose $\alpha_{n}(t)$ so that

$$
n a_{n} e^{i \alpha_{n}(t)} z^{n-1}=-n\left|a_{n}\right| r^{n-1} \quad(n=2, \cdots, N)
$$

Next define $f_{t}(z)$ by

$$
f_{t}(z)=z+\sum_{n=2}^{N} a_{n} e^{i \alpha_{n}(t)} z^{n}+\sum_{n=N+1}^{\infty} a_{n} z^{n},
$$

where $\alpha_{n}(t)$ varies continuously with $t$. Since $g(z) \leqq 0$ for $z \in C$, $|z|=r$, we have for all $t$ that

$$
f_{t}^{\prime}(z) \leqq 1-\sum_{n=2}^{N} n\left|a_{n}\right| r^{n-1} \quad(z \in C)
$$

Since $f_{t}^{\prime}(0)=1$ and $f_{t}^{\prime}(z)<0$ for $z \in C$ sufficiently close to the boundary of $U$, it follows for some $\xi=z\left(t_{0}\right)$ on $C$ that $f_{t_{0}}^{\prime}(\xi)=0$. Setting $f(z)=f_{t_{0}}(z)$ and $\alpha_{n}=\alpha_{n}\left(t_{0}\right)$, the result follows.

Corollary 1. If $\sum_{n=2}^{N} n\left|a_{n}\right|>1$, then there exist real numbers $\alpha_{2}, \cdots, \alpha_{N}\left(-\pi<\alpha_{j} \leqq \pi\right)$ and an $\varepsilon>0$ such that for each $\lambda,-\varepsilon \leqq$ $\lambda \leqq \varepsilon$,

$$
f_{2}(z)=z+e^{i \lambda}\left(\sum_{n=2}^{N} a_{n} e^{i \alpha_{n}} z^{n}+\sum_{n=N+1}^{\infty} a_{n} z^{n}\right) \notin L S .
$$

Proof. The corollary is established upon applying Lemma 1 to Theorem 2.

Corollary 2. If $\sum_{k=2}^{N} n_{k}\left|a_{n_{k}}\right|>1$, then there exist real numbers $\alpha_{2}, \cdots, \alpha_{N}\left(-\pi<\alpha_{j} \leqq \pi\right)$ and $\varepsilon>0$ such that for each $\lambda,-\varepsilon \leqq \lambda \leqq \varepsilon$,

$$
f_{\lambda}(z)=z+e^{i \lambda}\left(\sum_{k=2}^{N} a_{n_{k}} e^{i \alpha_{k}} z^{n_{k}}+\sum_{\substack{n=2 \\ n \neq n_{k}}}^{\infty} a_{n} z^{n}\right) \notin L S .
$$

Proof. The proof is the same as that of Corollary 1 except for a rearrangement of terms.
3. Extremal examples. One might ask for conditions under which $\sum_{n=2}^{\infty} n\left|a_{n}\right|>1$ guarantees the existence of a real number $\lambda$
such that $f_{\lambda}(z)=z+e^{i \lambda} \sum_{n=2}^{\infty} a_{n} z^{n} \notin S$. Before answering this question, we need

Lemma 2. For $n \geqq 3$ there exist real numbers $\alpha_{1}, \alpha_{2}, \cdots, \alpha_{n}$ such that $\left|\sum_{k=1}^{n} e^{i \alpha_{k}} z^{k}\right|<n(|z| \leqq 1)$.

Proof. We have $\left|\sum_{k=1}^{n} e^{i \alpha_{k}} e^{i k \theta}\right|=n$ if and only if there is a $\theta$ for which

$$
\alpha_{1}+\theta=\alpha_{k}+k \theta \quad(\bmod 2 \pi), \quad k=2, \cdots, n
$$

Clearly the $\alpha_{k}$ 's can be chosen to preclude the existence of such a $\theta$.

Theorem 3. (i) If $\sum_{k=2}^{3} n_{k}\left|a_{n_{k}}\right|>1$, then $f_{\lambda}(z)=z+e^{i \lambda}\left(a_{n_{2}} z^{n_{2}}+\right.$ $\left.a_{n_{3}} z^{n_{3}}\right) \notin L S$ for some $\lambda$.
(ii) If $N>3$, then we can find $\left\{a_{n_{k}} \mid k=2,3, \cdots, N\right\}$ such that $\sum_{k=2}^{N} n_{k}\left|a_{n_{k}}\right|>1$ and $f_{\lambda}(z)=z+e^{i \lambda} \sum_{k=2}^{N} a_{n_{k}} z^{n_{k}} \in S$ for all $\lambda$.

Proof. To prove (i), assume $\arg a_{n_{2}}=\alpha_{2}+\pi$ and $\arg a_{n_{3}}=\alpha_{3}+\pi$. Then for $\theta=\left(\alpha_{2}-\alpha_{3}\right) /\left(n_{3}-n_{2}\right)$, we have

$$
f_{\lambda}^{\prime}\left(r e^{i \theta}\right)=1-e^{i \lambda}\left(n_{2}\left|a_{n_{2}}\right| r^{n_{2}-1}+n_{3}\left|a_{n_{3}}\right| r^{n_{3}-1}\right) e^{i \beta},
$$

where

$$
\beta=\alpha_{2}+\frac{\left(n_{2}-1\right)\left(\alpha_{2}-\alpha_{3}\right)}{n_{3}-n_{2}} .
$$

Hence $f_{\lambda}(z) \notin L S$ for $\lambda=-\beta$.
To prove (ii), we choose $\left\{\alpha_{k}\right\}$ so that

$$
\left|\sum_{k=2}^{N} e^{i \alpha_{k}} z^{n_{k}-1}\right| \leqq A<N-1 \quad(|z| \leqq 1)
$$

For $1<B \leqq(N-1) / A$, set

$$
f_{\lambda}(z)=z+e^{i \lambda} \sum_{k=2}^{N} \frac{B e^{i \alpha_{k}}}{(N-1) n_{k}} z^{n_{k}} .
$$

Then

$$
f_{\lambda}^{\prime}(z)=1+\frac{B}{N-1} e^{i \lambda} \sum_{k=2}^{N-1} e^{i \alpha_{k}} z^{n_{k}-1}
$$

so that

$$
\operatorname{Re} f_{\lambda}^{\prime}(z) \geqq 1-\frac{B}{N-1}\left|\sum_{k=2}^{N-1} e^{i \alpha_{k}} z^{n_{k}-1}\right|>1-\frac{A B}{N-1} \geqq 0 \quad(z \in U)
$$

By a criterion of Kaplan [3], $f_{\lambda}(z) \in S$ for all $\lambda$.
4. A radius of starlikeness theorem. Denote by $F$ functions of the form $f(z)=z+\sum_{n=2}^{\infty} a_{n} z^{n}$ that are analytic in $U$ and satisfy $\operatorname{Re} f(z) / z>1 / 2$. MacGregor has shown [4] that the radius of starlikeness of $F$ is $1 / \sqrt{2}$. In this section we generalize this result. It is known [1] that the family $F$ is the closed convex hull of convex functions, and that a function $f(z)$ is in $F$ if and only if it can be expressed as

$$
\begin{equation*}
f(z)=\int_{x} \frac{z}{1-x z} d \mu(x) \tag{5}
\end{equation*}
$$

for some probability measure $\mu$ defined on the unit circle $X$.
In [2] Campbell proves
Lemma A. Let $M$ be a class of starlike functions with $b(r)=$ $\max \{\arg f(z) / z:|z|=r, f \in M\}$. If $\left|z\left(f^{\prime}(z) / f(z)\right)-a(r)\right| \leqq d(r)$ for $f \in M$, where $a(r)$ and $d(r)$ are continuous functions of $r$ satisfying $\lim _{r \rightarrow 1} a(r) / d(r)>0$, then the radius of starlikeness of the closed convex hull of $M$ is at least as large as the first positive root of $a(r)$ $d(r) \sec b(r)=0$.

In the sequel, we let

$$
\begin{equation*}
H_{\varepsilon}=\left\{z+e^{i \lambda} \sum_{n=2}^{\infty} z^{n} \| e^{i \lambda}-1 \mid \leqq \varepsilon<1\right\} \tag{6}
\end{equation*}
$$

and

$$
\begin{equation*}
G_{\varepsilon}=\left\{\frac{h(x z)}{x}\left|h \in H_{\varepsilon},|x|=1\right\}\right. \tag{7}
\end{equation*}
$$

We shall also need
Lemma 3. Set $a(r)=1 /\left(1-r^{2}\right)-\varepsilon^{2} r^{2} /\left(1-\varepsilon^{2} r^{2}\right)$ and $d(r)=r /\left(1-r^{2}\right)+$ $\varepsilon r /\left(1-\varepsilon^{2} r^{2}\right)$. For $g \in G_{\varepsilon}$, the values for $\left(z g^{\prime}(z) / g(z)\right),|z| \leqq r$, lie in a disk centered at $a(r)$ and having radius $d(r)$.

Proof. It suffices to prove the result for $h \in H_{\varepsilon}$ since the class $G_{\varepsilon}$ consists of rotations of these functions. Writing $h(z)=$ $\left(z+\left(e^{i \lambda}-1\right) z^{2}\right) /(1-z)$ for $h \in H_{c}$, we see that $z h^{\prime}(z) / h(z)=1 /(1-z)+$ $\left(\left(e^{i \lambda}-1\right) z\right) /\left(1+\left(e^{i \lambda}-1\right) z\right)$. Since

$$
\frac{-\varepsilon r}{1-\varepsilon r} \leqq \operatorname{Re} \frac{\left(e^{i \lambda}-1\right) z}{1+\left(e^{i \lambda}-1\right) z} \leqq \frac{\varepsilon r}{1+\varepsilon r} \quad(|z| \leqq r),
$$

it follows that

$$
\begin{equation*}
\frac{1}{1+r}-\frac{\varepsilon r}{1-\varepsilon r} \leqq \operatorname{Re} \frac{z h^{\prime}(z)}{h(z)} \leqq \frac{1}{1-r}+\frac{\varepsilon r}{1+\varepsilon r} . \tag{8}
\end{equation*}
$$

Thus the values for $z h^{\prime}(z) / h(z)$ are contained in a disk centered at

$$
\frac{1}{2}\left[\left(\frac{1}{1-r}+\frac{\varepsilon r}{1+\varepsilon r}\right)+\left(\frac{1}{1+r}-\frac{\varepsilon r}{1-\varepsilon r}\right)\right]=a(r)
$$

whose radius is

$$
\frac{1}{2}\left[\left(\frac{1}{1-r}+\frac{\varepsilon r}{1+\varepsilon r}\right)-\left(\frac{1}{1+r}-\frac{\varepsilon r}{1-\varepsilon r}\right)\right]=d(r) .
$$

Lemma 4. A function $g \in G_{\varepsilon}$ defined by (7) is starlike if and only if $\varepsilon \leqq 1 / 3$.

Proof. It suffices to consider $h \in H_{\varepsilon}$ defined by (6). Letting $r \rightarrow 1$ in the left-hand side of (8), we see that $\operatorname{Re} z h^{\prime}(z) / h(z) \geqq 0$ when $\varepsilon \leqq 1 / 3$. For $e^{i \lambda}-1=\varepsilon e^{i \sigma}$ and $\beta=\pi-\sigma$, we have $e^{i \beta} h^{\prime}\left(e^{i \beta}\right) / h\left(e^{i \beta}\right)=$ $1 / 2-\varepsilon /(1-\varepsilon)<0$ when $\varepsilon>1 / 3$.

Remark. Since the Pólya-Schoenberg conjecture is true [5], we know that for $f$ convex and $h \in H_{\varepsilon}(\varepsilon \leqq 1 / 3)$, the Hadamard product $h * f$ is starlike. However if $f$ is only required to be starlike, then $h_{\lambda} * f$ need not be in $S$ for any $h_{\lambda}$ defined by (4), $\lambda \neq 0$. To see this, observe that $h_{\lambda} * z /\left((1-z)^{2}\right)=z+e^{i \lambda} \sum_{n=2}^{\infty} n z^{n}$ is not in $S$ for $\lambda \neq 0$ because $f(z)=z+\sum_{n=2}^{\infty} a_{n} z^{n},\left|a_{2}\right|=2$, is in $S$ only if $f(z)=$ $z /\left((1-x z)^{2}\right),|x|=1$.

We now give a bound for the radius of starlikeness for $h * f$, $f \in F, h \in H_{\varepsilon}$ with $\varepsilon \leqq 1 / 3$.

THEOREM 4. Suppose $h(z)=z+e^{i \lambda} \sum_{n=2}^{\infty} z^{n} \in H_{s}, \quad \varepsilon \leqq 1 / 3$, and $f \in F$, with $a(r)$ and $d(r)$ defined in Lemma 3. Then $h * f$ is starlike in a disk $|z|<r_{0}$, where $r_{0}$ is the first positive root of

$$
\begin{equation*}
a(r)-\frac{d(r)}{\sqrt{\left(1-\varepsilon^{2} r^{2}\right)\left(1-r^{2}\right)}-\varepsilon r^{2}}=0 . \tag{9}
\end{equation*}
$$

Proof. In view of (5),

$$
(h * f)(z)=\int_{X} h * \frac{z}{1-x z} d \mu(x)=\int_{X} \frac{h(x z)}{x} d \mu(x), \quad|x|=1 .
$$

By Lemma 4, the kernel functions are all starlike. An application of Lemma 3 to Lemma A shows that $h * f$ is starlike in a disk whose radius is at least as large as the first positive root of

$$
\begin{equation*}
a(r)-d(r) \sec b(r)=0 \tag{10}
\end{equation*}
$$

where

$$
b(r)=\max _{|z|=r}\left\{\left.\arg \frac{g(z)}{z} \right\rvert\, g \in G_{\varepsilon}\right\}
$$

Since

$$
\begin{align*}
b(r) & \leqq \max _{|z|=r}|\arg (1+\varepsilon z)|+\max _{|z|=r}|\arg (1-z)| \\
& =\sin ^{-1}(\varepsilon r)+\sin ^{-1}(r)=t(r) \\
\sec b(r) & \leqq \sec t(r)=\frac{1}{\cos t(r)}=\frac{1}{\sqrt{\left(1-\varepsilon^{2} r^{2}\right)\left(1-r^{2}\right)-\varepsilon r^{2}}} \tag{11}
\end{align*}
$$

A substitution of the right-hand side of (11) into (10) yields the desired result.

Remark. If $\varepsilon=0$ then (9) reduces to $1 /\left(1-r^{2}\right)-r /\left(\left(1-r^{2}\right)^{3 / 2}\right)=0$, whose smallest positive root is $1 / \sqrt{2}$. This coincides with a sharp result of MacGregor [4].

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