

# The Spinor Heat Kernel in Maximally Symmetric Spaces

**Roberto Camporesi**

Theoretical Physics Institute, Department of Physics, University of Alberta, Edmonton, Alberta T6G 2J1, Canada

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**Abstract.** The heat kernel  $K(x, x', t)$  of the iterated Dirac operator on an  $N$ -dimensional simply connected maximally symmetric Riemannian manifold is calculated. On the odd-dimensional hyperbolic spaces  $K$  is a Minakshisundaram–DeWitt expansion which terminates to the coefficient  $a_{(N-1)/2}$  and is exact. On the odd spheres the heat kernel may be written as an image sum of WKB kernels, each term corresponding to a classical path (geodesic). In the even dimensional case the WKB approximation is not exact, but a closed form of  $K$  is derived both in terms of (spherical) eigenfunctions and of a “sum over classical paths.” The spinor Plancherel measure  $\mu(\lambda)$  and  $\zeta$  function in the hyperbolic case are also calculated. A simple relation between the analytic structure of  $\mu$  on  $H^N$  and the degeneracies of the Dirac operator on  $S^N$  is found.

## 1. Introduction

A maximally symmetric Riemannian manifold  $M$  of dimension  $N$  has an isometry group of maximum dimension  $N(N+1)/2$ .  $M$  is also a constant curvature space, i.e., the Riemann tensor takes the form

$$R_{abcd} = k(g_{ad}g_{bc} - g_{ac}g_{bd}), \quad (1.1)$$

where  $k$  is a constant. The Ricci tensor and curvature scalar are given by  $R_{ad} = k(N-1)g_{ad}$ , and  $R = kN(N-1)$ . Moreover,  $M$  is necessarily isometric to one of the following spaces: a) Euclidean space  $R^N$  ( $k=0$ ); b) the sphere  $S^N$  of radius  $a$  ( $k=1/a^2$ ); c) the real projective space  $P^N(R) = S^N/\sim$ , where  $\sim$  is the antipodal points identification ( $k$  is the same as for  $S^N$ ); d) the real hyperbolic space  $H^N(R)$  of radius  $a$  ( $k=-1/a^2$ ) (see ref. [24], vol. 1, p. 308). These spaces are all simply connected except for  $P^N(R)$ , which is doubly connected.

In the pseudo-Riemannian Lorentzian case [signature  $(-, +, \dots, +)$ ] we have, similarly, that the maximally symmetric spacetimes are Minkowski spacetime  $M^N$  (zero curvature), de Sitter spacetime  $(dS)_N$  (positive curvature), and anti-de Sitter spacetime  $(AdS)_N$  (negative curvature). In the Euclidean approach to