## **Decomposable Differential Operators** in a Cosmological Context

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**Abstract.** The integrability conditions for a certain second order ordinary differential equation in two variables are studied via the concept of decomposability of the associated differential operator. The results are applied to regain in a unified manner the known exact solutions for locally rotationally symmetric, spatially homogeneous cosmological models. In addition, new solutions are obtained.

## 1. Introduction

Einstein's equations

$$R_{ab} - \frac{1}{2}Rg_{ab} + Ag_{ab} = T_{ab} \tag{1.1}$$

for a spacetime metric tensor  $g_{ab}$ , where  $R_{ab}$  is the Ricci tensor,  $R = R_a^a$ , and  $\Lambda$  is the cosmological constant, can be solved exactly only in cases of rather high spacetime symmetry, and for relatively simple forms of the energy-momentum tensor  $T_{ab}$ .

In this paper, we consider exact solutions of (1.1) for spacetimes in which local coordinates  $(x^{\alpha}) = (t, x^{\alpha}) (a = 0, ..., 3; \alpha = 1, ..., 3)$  may be chosen so that one or more of the field equations, or combinations thereof, take the generic form

$$A_{1}\frac{\ddot{X}}{X} + A_{2}\frac{\ddot{Y}}{Y} + A_{3}\frac{\dot{X}^{2}}{X^{2}} + A_{4}\frac{\dot{Y}^{2}}{Y^{2}} + A_{5}\frac{\dot{X}}{X} + A_{6}\frac{\dot{Y}}{Y} + A_{7}\frac{\dot{X}\dot{Y}}{XY} + H(X, Y, t) = 0$$
(1.2)

where  $A_i \in \mathbf{R}(i = 1, ..., 7)$ , a dot denotes differentiation with respect to t, and X(t), Y(t) are metric component functions. This is the case, for example, when the spacetime is locally rotationally symmetric and admits a 1-parameter family of homogeneous hypersurfaces (see, e.g., Refs. [1-3]). If a first integral of (1.2) can be found, then together with the remaining field equations and the conservation equations  $T_{ib}^{ab} = 0$ , this often allows us to obtain a reduction of the system of equations to quadratures, or to a single ordinary differential equation in one variable, plus quadratures.