

### 138. The Relativity Theory in the Einstein Space under the Extended Lorentz Transformation Group

By Tsurusaburo TAKASU

Tohoku University, Sendai

(Comm. by Zyoiti SUETUNA, M.J.A., Nov. 12, 1963)

The general theory of relativity of A. Einstein was based on the non-definite quadratic differential form

$$(1) \quad dS^2 = g_{\mu\nu}(x^\sigma) dx^\mu dx^\nu, \quad (\lambda, \mu, \nu, \dots = 1, 2, 3, 4)$$

and grasped as the *Riemannian geometry of the Einstein space*:

$$(i) \quad R_{\mu\nu} = 0, \quad \left| \quad (ii) \quad R_{\mu\nu} = \frac{R}{4} g_{\mu\nu}, \right.$$

the path of a free particle being the geodesic curve:

$$(2) \quad \frac{d^2 x^\lambda}{dS^2} + \left\{ \begin{matrix} \lambda \\ \mu\nu \end{matrix} \right\} \frac{dx^\mu}{dS} \frac{dx^\nu}{dS} = 0.$$

The fundamental assumption was the so-called *principle of equivalence*. The merit was the *geometrization of physics*. But the demerit was the obscurity of the physical side caused by the laborious calculations in terms of  $g_{\mu\nu}$  and  $\left\{ \begin{matrix} \lambda \\ \mu\nu \end{matrix} \right\}$  as well as by too much forcing physical interpretations. Thus the Einstein's theory has remained merely as a *conjecture* for the last 47 years without becoming a decisive immortal theory.

With the hope to make it a decisive theory comparable with the Newton's theory, the present author ([1]–[14]) started with the expressibility of (1) in the form

$$(3) \quad dS^2 = g_{\mu\nu} dx^\mu dx^\nu = (-1)^{1+\delta_i^4} \omega^i \omega^i, \quad (\omega^i = \omega_\mu^i(x^\sigma) dx^\mu, |\omega_\mu^i| \neq 0)$$

except undergoing *extended* orthogonal transformations of  $\frac{1}{2}(1 + \delta_i^4)\omega^i$ , having discovered the *extended* orthogonal transformations with functions of coordinates ( $x^\sigma$ ) as coefficients and simplified calculations extremely by taking  $\omega_\mu^i(x^\sigma)$  and  $A_{\mu\nu}^\lambda$  in place of  $g_{\mu\nu} = \omega_\mu^i \omega_\nu^i$  and  $\left\{ \begin{matrix} \lambda \\ \mu\nu \end{matrix} \right\}$  respectively, where

$$(4) \quad A_{\mu\nu}^\lambda \stackrel{\text{def}}{=} \Omega_i^\lambda \frac{\partial \omega_\mu^i}{\partial x^\nu} \equiv -\omega_\mu^i \frac{\partial \Omega_i^\lambda}{\partial x^\nu}$$

is the *parameter of teleparallelism* of  $\omega_\mu^i(x^\sigma)$  and  $\Omega_i^\lambda(x^\sigma)$ , and

$$(5) \quad \Omega_i^\lambda \omega_\mu^i = \delta_\mu^\lambda \iff \Omega_m^\lambda \omega_\lambda^i = \delta_m^i,$$

the  $\delta$ 's being the Kronecker deltas. The equations of motion of a free particle were

$$(6) \quad \frac{d^2 \xi^i}{dS^2} = \frac{d}{dS} \frac{\omega^i}{dS} \equiv \omega_\lambda^i \left\{ \frac{d^2 x^\lambda}{dS^2} + A_{\mu\nu}^\lambda \frac{dx^\mu}{dS} \frac{dx^\nu}{dS} \right\} = 0,$$

whose finite equations are