

Cosmology in Terms of Wave Geometry (VII) Some Characteristics of the Universe.

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§ 1. **Introduction and summary.** Cosmology in terms of wave geometry is characterized by the equations:

$$(1.1) \quad \frac{\partial \Psi}{\partial x^i} = (\Gamma_i + \sum_i) \Psi$$

$$(1.2) \quad u^i = \Psi^\dagger A \gamma^i \Psi$$

$$(1.3) \quad u^i \nabla_i u^j = Q u^j.$$

(1.1) is the fundamental equation of wave geometry. (1.2) is the definition of u^i , where Ψ is the solution of (1.1) and A an hermitian matrix which makes $A \gamma^i$ hermitian.⁽¹⁾ (1.3) is the condition that u^i may generate a geodesic. From these equations it results that the universe is either (i) of de Sitter type or (ii) of Einstein type. We have here, on plausible grounds,⁽²⁾ adopted the de Sitter type as representing the universe.

u^i thus defined has been identified with the momentum-density vector of a constituent matter of the universe, and the phenomena of red-shifts of the spectral lines in nebulae, i. e. Hubble's velocity-distance relation, have been successfully explained.

The physical interpretations of u^i , however, present some difficulties, which are considered in detail in § 2. In § 3 we proceed to further properties of u^i and other physical quantities in our cosmology, defining a material energy tensor analogously to relativistic hydrodynamics.

§ 2. **The physical meanings of u^i and M .** In cosmology in terms of wave geometry, a nebula was considered as having two aspects, one being that of a particle to detect the structure of the universe, the other being that of probability-existence, all the nebulae being regarded as almost smeared out all over the universe.⁽³⁾ Therefore it seems natural to study the physical properties of the universe in a manner analogous to the gas theory or fluid mechanics in the classical theory.

(1) T. Sibata: this Journ., **8** (1938), 172, (W. G. No. 26).

(2) K. Itimaru: this Journ., **8** (1938), 240, (W. G. No. 31).

(3) Y. Mimura and T. Iwatsuki: this Journ., **8** (1938), 194 (W. G. No. 28).