Chiral models and the Einstein-Maxwell field equations

Hideo DOI and Ryuichi SAWAE

(Received December 25, 1989)

1. Introduction

The main objective in this paper is to provide a geometric picture of solutions of a (1 + 1)-dimensional reduction for the (1 + 3)-dimensional principal chiral model taking values in an arbitrary linear algebraic group.

Let G be a closed subgroup of the group scheme GL_N and assume that G is defined over **R**. The equations of motion for the SO(1, 2)-invariant chiral model on flat Minkowski space can be written

$$(1.1) d(t * d\sigma \cdot \sigma^{-1}) = 0$$

for $\sigma \in G(C[[t, z]])$. Here t, z are real variables, d is exterior differentiation, and * is the Hodge operator with respect to the Lorentz metric $(dt)^2 - (dz)^2$.

Let λ be a real parameter. Let \mathscr{A} denote an algebra $\{a = \sum_{n \in \mathbb{Z}} a_n \lambda^n \in C[[t, z, \lambda, \lambda^{-1}]]; \text{ ord } a_n \geq n\}$, where $\operatorname{ord} \varphi = \sup \{k \in \mathbb{Z}; \varphi \in (C[[t, z]]t + C[[t, z]]z)^k\}$. Set $\mathscr{A}^{\pm} = \mathscr{A} \cap C[[t, z, \lambda^{\pm 1}]], \mathscr{P}_G = G(\mathscr{A}^+) \text{ and } \mathscr{N}_G = \{g \in G(\mathscr{A}^-); g(t, z, \infty) = 1\}$. Then $G(\mathscr{A}) = \mathscr{N}_G \mathscr{P}_G$ (Lemma 2.3 and K. Takasaki [6, (3.17)]). This decomposition is used for solving (1.1).

THEOREM 1.1. There exist $w \in \mathcal{N}_G$ and $p \in \mathcal{P}_G$ such that $w^{-1}p = \gamma(z + \lambda t^2/2 + 1/2\lambda)$ for each $\gamma \in G(\mathbb{C}[[z]])$. Furthermore, if we set $\sigma = p(t, z, 0)$, then σ is a unique solution of (1.1) with $\sigma(0, z) = \gamma(z)$.

We give a proof of the theorem in §2 and derive an explicit formula for the solution σ with $\sigma(0, z) \in G(\mathbb{C}[z])$. Also we consider a transformation group for solutions of (1.1). As an application, we show in §3 a variant of the Geroch conjecture [3], that is to say, a real form $\mathscr{SU}(1, 2)$ of $SL_3(\mathbb{C}[[z]])$ acts transitively on the space of plane wave solutions of the Einstein-Maxwell field equations.

The authors would like to thank Prof. T. Kako for his aid to use the computer algebra system REDUCE 3.3.

2. The chiral models

To start with, we consider a manifest invariance of (1.1). We note that $d(t * d\tau^{-1} \cdot \tau) = -\operatorname{Ad} \tau^{-1}(d(t * d\tau \cdot \tau^{-1}))$ for any $\tau \in G(C[[t, z]])$. The following result is obvious.