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## Some remarks on Einstein spaces and spaces of constant curvature.

Dedicated to Professor Z. Suetuna on his 60th birthday.

By Kentaro YANO and Tsunero TAKAHASHI

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## §1. Preliminaries.

The object of the present paper is to generalise some of recent results of André Avez [1]\* to the case of non-compact Einstein spaces and to the case of spaces of constant curvature.

We shall here give notations and the formulas which will be used in the sequel.

Let *M* be an *n* dimensional Riemannian space of class C<sup>4</sup> with the fundamental metric tensor  $g_{\mu\lambda}$  whose signature is not necessarily positive definite. We denote by  $\mathcal{V}_{\mu}$  the covariant differentiation with respect to the Christoffel symbols  $\{\mu_{\lambda}^{\kappa}\}$ , by  $K_{\nu\mu\lambda\kappa}$  the curvature tensor, by  $K_{\mu\lambda}$  the Ricci tensor and by *K* the curvature scalar.

For an arbitrary skew-symmetric tensor field  $w: w_{\lambda_1 \lambda_2 \cdots \lambda_p}$  of order p, we write

(1.1) 
$$(dw)_{\mu\lambda_1\lambda_2\cdots\lambda_p} = (p+1)\mathcal{V}_{[\mu}w_{\lambda_1\lambda_3\cdots\lambda_p]}$$

and

(1.2)  $(\delta w)_{\lambda_1 \lambda_2 \cdots \lambda_p} = \nabla_{\mu} w^{\mu}{}_{\lambda_2 \lambda_3 \cdots \lambda_p} \,.$ 

Then the de Rham operator  $\Delta = d\delta + \delta d$  applied to w gives [2]

$$(\Delta w)_{\lambda_1 \lambda_2 \cdots \lambda_p} = g^{\nu \mu} \nabla_{\nu} \nabla_{\mu} w_{\lambda_1 \lambda_2 \cdots \lambda_p} -p K_{[\lambda_1}{}^{\mu} w_{[\mu] \lambda_2 \cdots \lambda_p]} - \frac{p(p-1)}{2} K_{[\lambda_1 \lambda_2}{}^{\nu \mu} w_{[\nu \mu] \lambda_2 \cdots \lambda_p]}.$$

Especially, if w is a vector field,

(1.3) 
$$(\Delta w)_{\lambda} = g^{\nu \mu} \nabla_{\nu} \nabla_{\mu} w_{\lambda} - K_{\lambda}^{\kappa} w_{\kappa}$$

and if w is a skew-symmetric tensor field of order two,

$$(\varDelta w)_{\lambda\kappa} = g^{\nu\mu} \nabla_{\!\!\nu} \nabla_{\!\!\mu} w_{\lambda\kappa} - 2K_{[\lambda}{}^{\mu} w_{[\mu]\kappa]} - K_{\lambda\kappa}{}^{\nu\mu} w_{\nu\mu}$$

or

(1.4) 
$$(\varDelta w)_{\lambda\kappa} = g^{\nu\mu} \nabla_{\nu} \nabla_{\mu} w_{\lambda\kappa} - (2K_{[\lambda}{}^{[\nu}A^{\mu]}_{\kappa]} + K_{\lambda\kappa}{}^{\nu\mu}) w_{\nu\mu} .$$

<sup>\*</sup> See the Bibliography at the end of the paper.