

Non-relativistic global limits of weak solutions of the relativistic Euler equation

By

Lu MIN and Seiji UKAI

1. Introduction

The relativistic Euler equation for a perfect fluid in two dimensional Minkowski space-time has the form ([9], [10])

$$(1.1) \quad \begin{aligned} \frac{\partial}{\partial t} \left\{ \frac{(p + \rho c^2)}{c^2} \frac{v^2}{c^2 - v^2} + \rho \right\} + \frac{\partial}{\partial x} \left\{ (p + \rho c^2) \frac{v}{c^2 - v^2} \right\} &= 0, \\ \frac{\partial}{\partial t} \left\{ (p + \rho c^2) \frac{v}{c^2 - v^2} \right\} + \frac{\partial}{\partial x} \left\{ (p + \rho c^2) \frac{v^2}{c^2 - v^2} + p \right\} &= 0. \end{aligned}$$

Here $v = v(x, t)$ is the classical coordinate velocity, $\rho = \rho(x, t)$ is the mass-energy density of the fluid, $p = p(\rho)$ is the pressure and c is the speed of light. On the other hand, the non-relativistic Euler equation is

$$(1.2) \quad \begin{aligned} \frac{\partial}{\partial t} \rho + \frac{\partial}{\partial x} (\rho v) &= 0, \\ \frac{\partial}{\partial t} (\rho v) + \frac{\partial}{\partial x} (\rho v^2 + p) &= 0. \end{aligned}$$

For the systems (1.1) and (1.2), the local existence theorems are known for the smooth solutions (see [4] and [5] for the full-dimensional case). Also, the global existence theorems are established for the one-dimensional isentropic motions $p = \rho^\gamma$, $\gamma > 1$ ([1] and [7]). In the case of the isothermal motions $p = \sigma^2 \rho$, where the sound speed σ is assumed to be the constant, the existence theorems with arbitrary initial data have been obtained both for (1.1) and (1.2), by J. Smoller and B. Temple [9] and by T. Nishida [6] respectively.

In physics, it is well-known that the classical mechanics reappears as the limit of the relativistic mechanics when $c \rightarrow \infty$, and in particular, it is easy to check that the relativistic Euler equation (1.1) reduces formally to the non-relativistic Euler equation (1.2) when $c \rightarrow \infty$. However, until now there are only local results for the limit of smooth solutions of the relativistic Euler equation ([5]). The aim of this paper is to discuss the convergence of weak solutions of (1.1) as $c \rightarrow \infty$. Since