

# NOTE ON THE BINOMIAL DISTRIBUTION

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The purpose of this note is to show that

$$(1) \quad f(x) = (-1)^n \frac{q^n n!}{\pi} \left(\frac{p}{q}\right)^x \frac{\sin \pi x}{x^{(n+1)}}$$

where  $n$  is an integer  $\geq 0$ ,  $0 < p < 1$ ,  $p + q = 1$ , and  $x^{(n+1)} = x(x-1)(x-2) \cdots (x-n)$ , is a function whose values at  $x = 0, 1, 2, \dots, n$  are the successive terms of the expansion of  $(q + p)^n$ , and also to consider the problem of fitting  $f(x)$  to an observed frequency distribution.

The statement made about (1) can be verified by evaluating (1) as an indeterminate form. On the other hand, (1) can be derived by observing that the  $x$ -th term ( $x$  an integer) of the expansion of  $(q + p)^n$  is

$$(2) \quad \frac{n!}{x!(n-x)!} p^x q^{n-x} = \frac{\Gamma(n+1) p^x q^{n-x}}{\Gamma(x+1) \Gamma(n-x+1)};$$

then (1) can be derived from (2) by means of the product expansions for  $\Gamma(x)$  and  $\sin x$ . This derivation of (1) from (2) can also be carried out by expressing (2) as a Beta function and then using

$$B(x+1, n-x+1) = \int_0^1 \frac{t^x}{(1+t)^{n+2}} dt = (-1)^n \frac{\pi}{(n+1)!} \frac{x^{(n+1)}}{\sin \pi x}.$$

This integration can be performed by means of the theory of residues.

Consider the problem of fitting (1) to an observed frequency distribution. We shall write (1) in the form

$$(3) \quad F(z) = ab^z \frac{\sin \pi x}{x^{(n+1)}}, \quad x = \frac{nb}{1+b} + h(z - \bar{z})$$

and determine the constants  $a$ ,  $b$ ,  $n$ , and  $h$  so that, when  $\bar{z}$  is the mean of the observed distribution,  $F(z)$  will fit the distribution.

The values of  $a$ ,  $b$ ,  $n$ , and  $h$  can be determined by the method of moments. Let  $\nu_2$ ,  $\nu_3$ , and  $\nu_4$ , denote the usual second, third, and fourth moments of the distribution, which are calculated in the usual way (as in W. P. Elderton, *Frequency-Curves and Correlation*) and not adjusted by any procedure such as Sheppard's adjustments. Also, use the usual notation  $\beta_1 = \frac{\nu_3^2}{\nu_2^3}$  and  $\beta_2 = \frac{\nu_4}{\nu_2^2}$ .