# Einstein H-umbilical submanifolds with parallel mean curvatures in complex space forms 

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#### Abstract

In this paper we determine H -umbilical Einstein submanifolds with parallel mean curvatures in complex space forms with non-negative holomorphic sectional curvatures.


## 1 Introduction

In Riemannian Geometry, Einstein manifolds are very important subject. When we focus our attention to submanifolds in complex space forms, there are many interesting results (cf. [1]). There are two important classes of submanifolds of a complex space form. One is the class of holomorphic submanifolds and another is the class of totally real submanifolds. A submanifold in a complex space form is said to be totally real if the complex structure of the ambient space carries each tangent vector to a normal vector. A totally real submanifold is called a Lagrangian submanifold if its real dimension is equal to the complex dimension of the ambient space. The classification of Lagrangian Einstein submanifolds of a complex space form is still open. We know the fact that a non-flat complex space form of complex dimension $\geq 2$ admits no totally umbilical Lagrangian submanifolds except the totally geodesic ones. So, B. Y. Chen [3] introduced the notion of H-umbilical submanifolds which are the simplest Lagrangian submanifolds next to the totally geodesic ones in a complex space form (for the definition see $\S 2$ ).

In this paper we investigate H -umbilical Einstein submanifolds of complex space forms with non-negative holomorphic sectional curvatures and give the following theorem:

[^0]Theorem 1.1 Let $M^{n}$ be a complete $n(\geq 3)$ dimensional Einstein $H$-umbilical submanifold with parallel mean curvature in an $n$-dimensional complex space form $\widetilde{M}_{n}(\tilde{c})$ with constant holomorphic sectional curvature $\tilde{c} \geq 0$. Then $M^{n}$ is congruent to a totally geodesic Lagrangian submanifold of $\widetilde{M}_{n}(\tilde{c})$ or $S^{1}\left(\frac{1}{\sqrt{\lambda}}\right) \times \mathbb{R}^{n-1}$ in $\mathbb{C}^{n}$, where we denote the radius of sphere in the parentheses (for the definition of $\lambda$ see §2).

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## 2 Preliminaries

In this section we explain preliminary results concerning Riemannian submanifolds and give some definitions.

Let $\widetilde{M}_{n}(\tilde{c})$ be an $n$-dimensional complex space form with constant holomorphic sectional curvature $\tilde{c}$. Let $\langle\cdot, \cdot\rangle$ and $J$ be its Riemannian metric and complex structure, respectively. Then, the curvature tensor $\bar{R}$ of $\widetilde{M}_{n}(\tilde{c})$ is obtained by

$$
\begin{align*}
\bar{R}(X, Y) Z= & \frac{\tilde{c}}{4}\{\langle Y, Z\rangle X-\langle X, Z\rangle Y+\langle J Y, Z\rangle J X-\langle J X, Z\rangle J Y \\
& -2\langle J X, Y\rangle J Z\}, X, Y, Z \in T \widetilde{M} \tag{2.1}
\end{align*}
$$

where $T \widetilde{M}$ denotes the tangent bundle of $\widetilde{M}$.
Let $M^{n}$ be a connected $n$-dimensional real submanifold of $\widetilde{M}_{n}(\tilde{c})$. Then, $M^{n}$ is said to be a Lagrangian submanifold if the complex structure $J$ of $\widetilde{M}_{n}(\tilde{c})$ carries each tangent space of $M^{n}$ into its corresponding normal space.

For a Lagrangian submanifold $M^{n}$ of $\widetilde{M}_{n}(\tilde{c})$, the Gauss and Ricci equations become

$$
\begin{align*}
&\langle R(X, Y) Z, W\rangle= \frac{\tilde{c}}{4}(\langle Y, Z\rangle\langle X, W\rangle-\langle X, Z\rangle\langle Y, W\rangle) \\
&+\langle\sigma(Y, Z), \sigma(X, W)\rangle-\langle\sigma(X, Z), \sigma(Y, W)\rangle  \tag{2.2}\\
& X, Y, Z, W \in T M
\end{aligned}, ~ \begin{aligned}
\left\langle R^{\perp}(X, Y) J Z, J W\right\rangle= & \left\langle\left[A_{J Z}, A_{J W}\right] X, Y\right\rangle+\frac{\tilde{c}}{4}(\langle Y, Z\rangle\langle X, W\rangle \\
& -\langle X, Z\rangle\langle Y, W\rangle) \\
& X, Y, Z, W \in T M
\end{align*}
$$

where $\sigma$ and $A_{\nu}$ denote the second fundamental form of $M^{n}$ and the shape operator in the direction $\nu$, respectively, and $R^{\perp}$ is the curvature tensor of the normal bundle $T^{\perp} M^{n}$ of $M^{n}$.

The Gauss and Weingarten formulas are the following:

$$
\begin{gathered}
\bar{\nabla}_{X} Y=\nabla_{X} Y+\sigma(X, Y), X, Y \in T M^{n} \\
\bar{\nabla}_{X} \nu=-A_{\nu} X+\nabla_{X}^{\perp} \nu, X \in T M^{n}, \nu \in T^{\perp} M^{n}
\end{gathered}
$$

where $\bar{\nabla}, \nabla$ and $\nabla^{\perp}$ denote the Levi Civita connection of $\widetilde{M}_{n}(\tilde{c})$, that of $M^{n}$ and the normal connection of $T^{\perp} M^{n}$.

Lagrangian submanifold $M^{n}$ of $\widetilde{M}_{n}(\tilde{c})$ is called H-umbilical if the second fundamental form $\sigma$ of $M^{n}$ takes the following form for some functions $\lambda$ and $\mu$ with respect to some local orthnormal frame field $e_{1}, \ldots, e_{n}$ on $M^{n}$ :

$$
\begin{align*}
& \sigma\left(e_{1}, e_{1}\right)=\lambda J e_{1}, \sigma\left(e_{2}, e_{2}\right)=\cdots=\sigma\left(e_{n}, e_{n}\right)=\mu J e_{1}  \tag{2.4}\\
& \sigma\left(e_{1}, e_{j}\right)=\mu J e_{j}, \sigma\left(e_{j}, e_{k}\right)=0, j \frac{1}{\tau} k, j, k=2, \ldots, n .
\end{align*}
$$

The mean curvature vector $\mathfrak{H}$ of $M^{n}$ is defined by

$$
\begin{equation*}
\mathfrak{H}=\sum_{i=1}^{n} \sigma\left(e_{i}, e_{i}\right), \tag{2.5}
\end{equation*}
$$

where $e_{1}, \ldots, e_{n}$ is a orthonormal frame field of $M^{n}$.
$M^{n}$ is said to be of parallel mean curvature if $\nabla^{\perp} \mathfrak{H}=0$ is satisfied. And $M^{n}$ is said to be of constant mean curvature if $\langle\mathfrak{H}, \mathfrak{H}\rangle$ is constant on $M^{n}$. If $M^{n}$ has the parallel mean curvature, then $\langle\mathfrak{H}, \mathfrak{H}\rangle$ is always constant.

According to (2.2), the Ricci tensor $S$ of $M^{n}$ is given by

$$
\begin{align*}
S(X, Y)= & \frac{\tilde{\tilde{c}}}{4}(n-1)\langle X, Y\rangle+\langle\sigma(X, Y), \mathfrak{H}\rangle  \tag{2.6}\\
& -\sum_{i=1}^{n}\left\langle\sigma\left(X, e_{i}\right), \sigma\left(Y, e_{i}\right)\right\rangle, X, Y \in T M^{n}
\end{align*}
$$

For the scalar curvature $\rho$ of $M^{n}$, using (2.5) and (2.6), we obtain

$$
\begin{equation*}
\rho=\frac{\tilde{c}}{4} n(n-1)+\langle\mathfrak{H}, \mathfrak{H}\rangle-\|\sigma\|^{2} \tag{2.7}
\end{equation*}
$$

where $\|\sigma\|^{2}$ denotes the square of the length of $\sigma$.

## 3 Lemmas

In this section we present two lemmas to prove our main theorem.
We explain the equation of the Laplacian $\Delta\|\sigma\|^{2}$ of the function $\|\sigma\|^{2}$ given by J. Simons [8]. Let $E_{1}, \ldots, E_{n}$ be a local orthonormal frame field of $M^{n}$ around a point $p \in M^{n}$ which satisfy $E_{i}(p)=e_{i}$ and $\nabla E_{i}(p)=0(i=$ $1, \ldots, n)$. Then, we have

Lemma 3.1 ([8], [4]) Let $\iota: M^{n} \rightarrow \bar{M}^{n+k}$ be an isometric immersion of an $n$-dimensional Riemannian manifold into an $(n+k)$-dimensional Riemannian manifold $\bar{M}^{n+k}$. Then, the following equation holds:

$$
\begin{align*}
\frac{1}{2} \Delta\|\sigma\|^{2}= & \left\|\nabla^{\prime} \sigma\right\|^{2}+\sum_{i, j, k}\left\langle\bar{\nabla}_{e_{j}}\left(\left(\nabla_{E_{k}}^{\prime} \sigma\right)\left(E_{i}, E_{i}\right)\right), \sigma\left(e_{j}, e_{k}\right)\right\rangle \\
& +\sum_{i, j, k}\left\langle\bar{\nabla}_{e_{i}}\left(\bar{R}\left(E_{j}, E_{k}\right) E_{i}\right) N, \sigma\left(e_{j}, e_{k}\right)\right\rangle \\
& +\sum_{i, j, k}\left\langle R^{\perp}\left(e_{i}, e_{j}\right) \sigma\left(e_{i}, e_{k}\right), \sigma\left(e_{j}, e_{k}\right)\right\rangle  \tag{3.1}\\
& -\sum_{i, j, k}\left\langle\sigma\left(R\left(e_{i}, e_{j}\right) e_{i}, e_{k}\right), \sigma\left(e_{j}, e_{k}\right)\right\rangle \\
& -\sum_{i, j, k}\left\langle\sigma\left(e_{i}, R\left(e_{i}, e_{j}\right) e_{k}\right), \sigma\left(e_{j}, e_{k}\right)\right\rangle \\
& \left.+\sum_{i, j, k}\left\langle\bar{\nabla}_{e_{i}}\left(\bar{R}\left(E_{i}, E_{j}\right) E_{k}\right), \sigma\left(e_{j} \cdot e_{k}\right)\right\rangle\right\rangle
\end{align*}
$$

where $\langle\cdot, \cdot\rangle, R$ and $\bar{R}$ denote the metric tensor of $\bar{M}$, the Riemannian curvature tensor of $M$ and the Riemannan curvature tensor of $\bar{M}$, respectively, and ( $)^{N}$ denotes the normal component of a vector.

Using Lemma 3.1, (2.1), (2.2) and (2.3), we get the following:
Lemma 3.2 Let $M^{n}$ be a Lagrangian submanifold in a complex space form $\widetilde{M}_{n}(\tilde{c})$. Then, we have the following equation:

$$
\begin{align*}
\frac{1}{2} \Delta\|\sigma\|^{2}= & \left\|\nabla^{\prime} \sigma\right\|^{2}+\frac{\tilde{c}}{4}(n+1)\|\sigma\|^{2}-\frac{\tilde{c}}{2}\langle\mathfrak{H}, \mathfrak{H}\rangle \\
& +\sum_{i, j, k}\left\langle\bar{\nabla}_{e_{j}}\left(\left(\nabla_{E_{k}}^{\prime} \sigma\right)\left(E_{i}, E_{i}\right)\right), \sigma\left(e_{j} . e_{k}\right)\right\rangle  \tag{3.2}\\
& +\sum_{\alpha, \beta=1}^{n} \operatorname{Tr}\left(A_{\alpha} A_{\beta}-A_{\beta} A_{\alpha}\right)^{2}-\sum_{\alpha, \beta=1}^{n}\left(\operatorname{Tr} A_{\alpha} A_{\beta}\right)^{2} \\
& +\sum_{j, k}\left\langle A_{\mathfrak{H}} e_{j}, A_{\sigma\left(e_{j}, e_{k}\right)} e_{k}\right\rangle
\end{align*}
$$

where $A_{\alpha}$ is the shape operator of $M^{n}$ in the direction $J e_{\alpha}$.

## 4 Proof of Theorem 1.1

Let $e_{1}, \ldots, e_{n}$ be an orthonormal basis of $T_{p} M$ which satisfies (2.4). Using (2.4) and (2.6), we have the following for the Ricci tensor $S$ of $M$ :

$$
\begin{align*}
& S\left(e_{1}, e_{1}\right)=\frac{n-1}{4} \tilde{c}+(n-1) \mu(\lambda-\mu) \\
& S\left(e_{j}, e_{j}\right)=\frac{n-1}{4} \tilde{c}+\mu(\lambda+(n-3) \mu) \quad(j \geq 2)  \tag{4.1}\\
& S\left(e_{i}, e_{k}\right)=0 \quad(i \neq k)
\end{align*}
$$

Since $M^{n}$ is Einstein and $n \geq 3$, using (4.1), we are led to

$$
\begin{equation*}
\mu(\lambda-2 \mu)=0 \tag{4.2}
\end{equation*}
$$

Because of the Ricci curvature of Einstein manifold is constant, we have $\mu(\lambda-\mu)=$ constant. Using this fact and (4.2), we deduce that $\mu=$ constant. So, either $\mu \equiv 0$ or $\lambda=2 \mu=$ constant $\frac{1}{\tau} 0$ is satisfied on $M$.

We discuss dividing into the following two cases:
Case $1 \mu \equiv 0$; Case $2 \lambda=2 \mu \neq 0$.
Case 1 Using (2.4) and $\mu \equiv 0$, we have $\mathfrak{H}=\lambda J e_{1}$. Since $\mathfrak{H}$ is parallel, $\lambda$ is constant. Because of the scalar curvature $\rho$ of Einstein manifold is constant, we conclude that $\|\sigma\|^{2}$ is constant from (2.7). According to (2.4) and (3.2) we obtain the following:

$$
\begin{equation*}
0=\frac{1}{2} \Delta\|\sigma\|^{2}=\frac{\tilde{c}}{4}(n-1) \lambda^{2}+\left\|\nabla^{\prime} \sigma\right\|^{2} . \tag{4.3}
\end{equation*}
$$

When $\tilde{c}$ is positive, we have $\lambda^{2}=0$ and $\left\|\nabla^{\prime} \sigma\right\|^{2}=0$ from (4.3). So $M^{n}$ is totally geodesic. When $\tilde{c}=0$, we deduce that $M^{n}$ is a parallel submanifold in $\mathbb{C}^{n}$ from (4.3). According to the classification in [5], we conclude that either $M^{n}$ is congruent to $S^{1}\left(\frac{1}{\sqrt{\lambda}}\right) \times \mathbb{R}^{n-1}$ or a totally geodesic submanifold.

Case 2 In the following we shall show that this case cannot occur.
In this case, using the fact that $\lambda=2 \mu=$ constant and (2.4), we can deduce that both $\langle\mathfrak{H}, \mathfrak{H}\rangle$ and $\|\sigma\|^{2}$ are constant on $M$. According to (3.2), we get the following equation:

$$
\begin{aligned}
0=\frac{1}{2} \Delta\|\sigma\|^{2}= & \left(n^{2}-1\right) \mu^{2}\left(\frac{\tilde{c}}{4}+\mu^{2}\right)+\left\|\nabla^{\prime} \sigma\right\|^{2} \\
& +\sum_{i, j, k}\left\langle\bar{\nabla}_{e_{j}}\left(\left(\nabla_{E_{k}}^{\prime} \sigma\right)\left(E_{i}, E_{i}\right)\right), \sigma\left(e_{j}, e_{k}\right)\right\rangle .
\end{aligned}
$$

From this equation, we have $\sum_{i}\left(\nabla_{E_{k}}^{\prime} \sigma\right)\left(E_{i}, E_{i}\right) \neq 0$. This contradicts our assumption $\nabla^{\perp} \mathfrak{H}=0$. So this case cannot occur. We have thus proved the theorem.

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